COST OF A MUON COOLING TEST RING

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Muon Ring Cooler Workshop Neutrino Factory and Muon Collider Collaboration Lewis Hall, Room 228 Dept. of Physics and Astronomy University of Mississippi-Oxford University, Mississippi 38677 11–12 March 2004

Popular Experimental Papers -22/40 Feature Neutrinos

$\begin{array}{llllllllllllllllllllllllllllllllllll$	Experiment	Subject	Journal	Cites
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	SuperK BNL598 SPEAR EMC	Atmospheric ν Oscillations Heavy Particle J Narrow Resonance in $e^+ e^-$ Spin Asymmetry and g_1	PRL 81 (1998) 1562 PRL 33 (1974) 1404 PRL 33 (1974) 1406 PL B206 (1988) 364	$1989 \\ 1305 \\ 1225 \\ 1186$
$\begin{array}{llllllllllllllllllllllllllllllllllll$	$\operatorname{EMC}_{\operatorname{CDF}}$	Spin Structure of the Proton Observation of the Top Quark	NP B328 (1989) 1 PRL 74 (1995) 2626	$\begin{array}{c} 1028\\ 966 \end{array}$
BNL181 $K_2^{V} \to \pi\pi$ PRL 13 (1964) 138892FNAL2889.5 GeV/c² Dimuon ResonancePRL 39 (1977) 252882CHOOZLimits on Neutrino OscillationsPL B466 (1999) 415821SNO ⁸ B Solar NeutrinosPRL 87 (2001) 071301819UA1Large E_T Electrons + Missing EPL B122 (1983) 103816SuperKAtmospheric μ/e Neutrino RatioPL B335 (1994) 237788MARK ILepton Production in e^+e^- PRL 35 (1975) 1489780UA1Lepton Pairs around 95 GeV/c²PL B126 (1983) 398779HomestakeSolar Electron Neutrino FluxApJ 496 (1998) 505754UA2 $Z^0 \to e^+e^-$ PL B129 (1983) 130746CHOOZLimits on Neutrino OscillationsPL B123 (1983) 275702UA2Large E_T Electrons + Missing EPL B122 (1983) 476681CLEORadiative Penguin DecayPRL 74 (1995) 2885656JADEMulti Jets in e^+e^- ZP C33 (1986) 23654SuperKAtmospheric μ/e Neutrino RatioPL B436 (1998) 9610LSNDEvidence for $\nu_{\mu} \to \nu_e$ PRL 77 (1996) 3082601SuperKAtmospheric Neutrino FluxPL B436 (1998) 33596SuperKSolar Neutrino DataPRL 77 (1996) 1683573CDFEvidence for $\nu_{\mu} \to \nu_e$ PRL 417 (1999) 127552SuperKSolar Neutrino ObservationsPL B447 (1992) 1774552GargamelleNeutrino Interactions with Z^0PRL 81 (1922) 1774552	ARGUS D0	Observation of $B^0 - \overline{B}^0$ Mixing Observation of the Top Quark	PL B192 (1987) 245 PRL 74 (1995) 2632	$\begin{array}{c} 924\\920\end{array}$
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	BNL181	$K_2^0 \to \pi \pi$	PRL 13 (1964) 138	892
CHOOZLimits on Neutrino OscillationsPL B466 (1999) 415821SNO 8B Solar NeutrinosPRL 87 (2001) 071301819UA1Large E_T Electrons + Missing EPL B122 (1983) 103816SuperKAtmospheric μ/e Neutrino RatioPL B335 (1994) 237788MARK ILepton Production in $e^+ e^-$ PRL 35 (1975) 1489780UA1Lepton Pairs around 95 GeV/c ² PL B126 (1983) 398779HomestakeSolar Electron Neutrino FluxApJ 496 (1998) 505754UA2 $Z^0 \rightarrow e^+e^-$ PL B129 (1983) 130746CHOOZLimits on Neutrino OscillationsPL B122 (1983) 476681CLEORadiative Penguin DecayPRL 74 (1995) 2885656JADEMulti Jets in $e^+ e^-$ ZP C33 (1986) 23654SuperKAtmospheric μ/e Neutrino RatioPL B280 (1992) 146649SNONeutrino Oscillations and Z^0 PRL 89 (2002) 011301637SuperKAtmospheric μ/e Neutrino RatioPL B436 (1998) 33596SuperKAtmospheric Neutrino FluxPL B436 (1998) 33596SuperKNeutrino Strom SN1987APRL 58 (1987) 1490593SPEARScond Resonance in $e^+ e^-$ PRL 81 (1992) 3720561GallexSolar Neutrino ObservationsPL B447 (1999) 127558LSNDEvidence for $\nu_{\mu} \rightarrow \nu_{e}$ PRL 81 (1992) 1774552GargamelleNeutrino Interactions with Z^0 PL B46 (1973) 138543IMBNeu	FNAL288	9.5 GeV/c^2 Dimuon Resonance	PRL 39 (1977) 252	882
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$	Homestake	Solar Electron Neutrino Flux	ApJ 496 (1998) 505	754
CHOO2Limits on Neutrino OscillationsPL B420 (1998) 397735EMCNucleon Structure Function F_2^N PL B123 (1983) 275702UA2Large E_T Electrons + Missing EPL B122 (1983) 476681CLEORadiative Penguin DecayPRL 74 (1995) 2885656JADEMulti Jets in $e^+ e^-$ ZP C33 (1986) 23654SuperKAtmospheric μ/e Neutrino RatioPL B280 (1992) 146649SNONeutrino Oscillations and Z^0 PRL 89 (2002) 011301637SuperKAtmospheric μ/e Neutrino RatioPL B433 (1998) 9610LSNDEvidence for $\nu_{\mu} \rightarrow \nu_{e}$ PRL 77 (1996) 3082601SuperKAtmospheric Neutrino FluxPL B436 (1998) 33596SuperKNeutrinos from SN1987APRL 58 (1987) 1490593SPEARSecond Resonance in $e^+ e^-$ PRL 33 (1974) 1453579SuperKSolar Neutrino DataPRL 77 (1996) 1683573CDFEvidence for the Top QuarkPR D50 (1994) 2966565SuperKsolar Neutrino ObservationsPL B447 (1999) 127558LSNDEvidence for $\nu_{\mu} \rightarrow \nu_{e}$ PRL 81 (1992) 1774552GargamelleNeutrinos from SN1987APRL 86 (2001) 5651524SMCDeuteron Structure $g_1(x)$ PL B302 (1993) 533519SuperKAtmospheric Neutrino FluxPL B205 (1988) 416501	UA2	$Z^0 \rightarrow e^+ e^-$	PL B129 (1983) 130	$746 \\ 725$
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$OA2$ Large E_T Electrons + Missing E PL B122 (1983) 476 $O81$ $CLEO$ Radiative Penguin DecayPRL 74 (1995) 2885 656 JADEMulti Jets in $e^+ e^-$ ZP C33 (1986) 23 654 SuperKAtmospheric μ/e Neutrino RatioPL B280 (1992) 146 649 SNONeutrino Oscillations and Z^0 PRL 89 (2002) 011301 637 SuperKAtmospheric μ/e Neutrino RatioPL B433 (1998) 9 610 LSNDEvidence for $\nu_{\mu} \rightarrow \nu_{e}$ PRL 77 (1996) 3082 601 SuperKAtmospheric Neutrino FluxPL B436 (1998) 33 596 SuperKNeutrinos from SN1987APRL 58 (1987) 1490 593 SPEARSecond Resonance in $e^+ e^-$ PRL 33 (1974) 1453 579 SuperKSolar Neutrino DataPRL 77 (1996) 1683 573 CDFEvidence for the Top QuarkPR D50 (1994) 2966 565 SuperKSolar Neutrino ObservationsPL B447 (1992) 127 558 LSNDEvidence for $\nu_{\mu} \rightarrow \nu_{e}$ PRL 81 (1992) 1774 552 GargamelleNeutrino Interactions with Z^0 PL B46 (1973) 138 543 IMBNeutrinos from SN1987APRL 86 (2001) 5651 524 SMCDeuteron Structure $g_1(x)$ PL B302 (1993) 533 519 SuperKAtmospheric Neutrino FluxPL B205 (1988) 416 501		Nucleon Structure Function F_2	PL B123 (1983) 275 DI B123 (1983) 476	(UZ
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SuperKAtmospheric Neutrino FluxPL $B205$ (1988) 416501	SMC	Deuteron Structure $a_1(x)$	PL B302 (1993) 533	$524 \\ 519$
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KJ. Kim	Cooling Formulas–Solenoids	PRL 85 (2000) 760	24
G. Penn	μ Cooling Envelopes–Solenoids	PRL 85 (2000) 764	23

Physics with Black Holes of Known Mass at a Muon Collider

- Only the Muon Collider can produce black holes of known mass.
- Known mass could be critical in measuring: Quantum Black Hole Remnants, Scanning production turn on, Gravitons as missing energy.
- CLIC e^+e^- suffers from beamstrahlung 5 TeV spectrum from Greg Landsberg



Oak Ridge Sector Cyclotron, AIP 9 (1972) 54

Uranium energy (MeV/u)	10
Relativistic Energy Limit	100
Min. q/A (for 10 MeV/u)	0.15
$B\rho_{max}$ (kG-cm)	3018
E Constant, K (E = q^2/A)	440
Max. magnetic field (kG)	16.0
Magnet fraction (52^0 hills)	0.58
Number of sectors	4
Injection E, U ion (MeV/u)	0.6
Radius Ratio (R_f/R_i)	3 - 4.3
Extraction mean R (m)	3.15
RF freq. range, (MHz)	6 - 14
Magnet weight, tons	2300

Peak V, fundamental, kV	250
2nd Harmonic V, % fund.	26
Power, fundamental, kW	400
Power, 2nd harmonic, kW	100
Resonator Length, m	8.6
Resonator dia. (max.), m	3.3
Amplitude Stability	$1 \text{ in } 10^4$
Phase stability, deg	± 0.1
Energy Ratio (E_f/E_i)	9 - 19
Injection $R, R_f/R_i=3$ (m)	1.05
RF freq., 10 MeV/u (MHz)	13.22
Harmonic $\#$ (10 MeV/u U)	6



Cost of Four 201.25 MHz RF

Cavities

- Length = 42 cm
 Radius (Aperture) = 20 cm
 Gradient = 16 MV/m
 A wedge shape may be needed to fit
- 26 MV/orbit \rightarrow 2 GV / 80 orbits If I lose and gain 2 GeV, I can cool.
- Rough Estimates from Al Moretti Four cells at \$91000 each \rightarrow \$360000 4 MW RF Power Supply \rightarrow \$1300000

Cost of Four 1.6 T Dipole Magnets

Coil Circumference = 170 cm "Solenoid" Height = 30 cm 400 000 Ampere turns per magnet 2500 Amps in 160 turns 15mm square copper conductor Two 135mm square coils/magnet + H₂O Coil Length = 81 turns \times 170cm = 140 m $(1.7 \times 10^{-8}) (140) / .015^2 = 0.011 \Omega$ $Volts = IR = 2500 \times 0.011 = 28 V$ Power = $I^2 R = 2500^2 \times 0.011 = 70 kW$ 70 kW \times 8 coils = 560 kW 24 hrs \times 560 kW \times \$.05/kW-hr = \$670/day Copper Volume = $.015^2 \times 140 \times 8 = .25 \text{ m}^3$ Copper Mass = 8900 kg/m³ × .25 = 2200 kg Cost at 20 / kg of Copper = 44000Cost of 0.5 MW Dynapower Supply = \$60,000 Steel Vol. = $4 \times 0.5 \text{m} \times 1 \text{m} \times 1.3 \text{m} = 2.6 \text{ m}^3$ Steel Mass = 7900 kg/m³ × 2.6 m³ = 21000 kg Steel Cost = $21000 \text{ kg} \times \$4/\text{kg} = \$84000$

Cost of Scintillating Fiber Tracker

at Room Temperature

Four X – Y – Stereo Planes Multi Anode Hamamatsu PMTs Multi Hit TDCs (e.g. LeCroy 3377)
1.5mm scintillating fibers Resolution = 1.5mm / √12 = 0.45mm Must check multiple scattering in the ring!!!
200 fibers / plane → 30cm coverage
2400 channels × \$200/channel = \$480000







B-Field

- Magnet 2.9 Tesla
 4 concentric coils
 Weak focusing
 Azimuthally symmetric field
- dE/dx Injection radius = 120 mm, p = 105 MeV/c, 0.3 mbar hydrogen
- Anti-protons adiabatically spiral to the center
- dE/dx cannot be too high

B-Field Continue...

- Final anti-proton swarm r = 15 mm h = 40 mmKE = 2 keV
- Pulsed electric kicker in Z 80 ns pulse 20 ns rise
- 20 microsecond spiral time
- A long bunch train is coalesced into one swarm

Challenges

- more focusing \implies more dE/dx \implies faster spiral \implies still adiabatic
- more focusing \Rightarrow greater beam acceptance \Rightarrow accept high emittance muons





- Cyclotron frequency $f = \omega/2\pi = qB/2\pi m$
- $f_{\overline{p}} / f_{\mu} = 938/106 = 8.8$ 20 μ S spiral $\rightarrow 2.3 \ \mu$ S spiral Also can increase B

Muon Swarm Size Estimate

Put 10¹² muons at a point
 Take B = 2.9 Tesla
 Set electric repulsion = Lorentz force
 Find radius
 Estimate, not orbit!!!

•
$$\mathbf{E} = \mathbf{vB}; \mathbf{v} = \mathbf{qBr/m}$$

 $\mathbf{10^{12}} \mathbf{q} / (4\pi\epsilon_0 \mathbf{r}^2) = \mathbf{q} \mathbf{r} \mathbf{B}^2 / \mathbf{m}$
 $\mathbf{r} = [\mathbf{10^{12}} \mathbf{m} / (4\pi\epsilon_0 \mathbf{B}^2)]^{1/3}$
 $= \mathbf{6mm}$

- Put a wire through the muon swarm Neutralize the charge! 10^{12} electrons move in a $\mu S \rightarrow 1$ Amp
- Anti-proton Anti-cyclotron momentum p = .3 B R = $.3 \times 2.9 \times 0.12 = .105 \text{ GeV/c}$

Damped Harmonic Oscillator

- Generalized Angular Momentum $L_g = L_z - er A_{\theta},$ NIM A278 (1989) 368
- Quasipotential Well, $\eta = e/M$ U(r,z) = V(r,z) - (1/2 η r²) (L_g/M + η r A_{θ})²
- (a) U'(r,0)[MeV] vs r [mm] for various L_g
 a, b, c, and d are stable orbit radii
- (b) $[U'(r_0,z) U'(r_0,0)]$ [MeV] vs z [mm] for various r_0

