

# Neutrino oscillation experiments: Present and future

**Ryan Patterson**  
Caltech

Conference on Flavor Physics and CP Violation  
May 25, 2022

# Neutrino oscillations

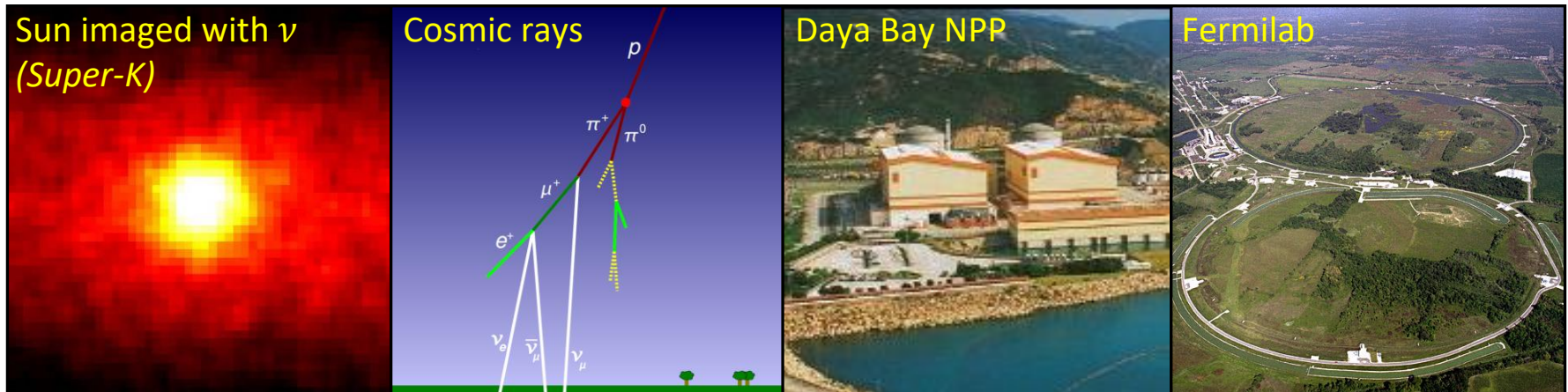
$$P(\nu_\alpha \rightarrow \nu_\beta) \sim \sin^2 \left( \Delta m_{ij}^2 \frac{L}{4E} \right)$$

travel distance ("baseline")  
neutrino energy

A given source has characteristic  $E$  (and  $\nu_\alpha$ ).  
Suggests optimal  $L$  for maximizing oscillations.

(simplified picture  $\Rightarrow$  exceptions)

**solar, atmospheric, reactor, and accelerator  $\nu$  sources**



# Experimental overview

## 3-flavor oscillations

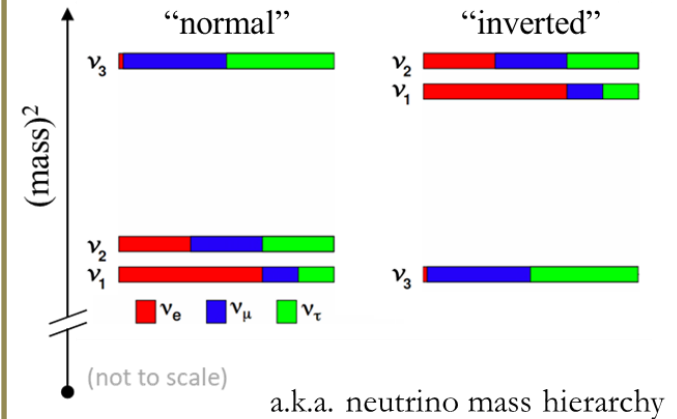
*Parameters best constrained by various sources*

Solar (MSW)	}	$\Delta m_{21}^2$	$\theta_{12}$						
Reactor ( $\sim 60$ km)									
Reactor ( $\sim 2$ km)	$\rightarrow$	$ \Delta m_{32}^2 $	$\theta_{13}$						
Atmospheric	$\rightarrow$	$ \Delta m_{32}^2 $	$\theta_{23}$	[ $\nu$ MO]					
Accelerator (LBL)	$\rightarrow$	$ \Delta m_{32}^2 $	$\theta_{23}$	$\nu$ MO	$\delta_{CP}$	[ $\theta_{13}$ ]			

$\rightarrow \sin^2\theta_{ij}$  and  $|\Delta m_{ij}^2|$  all measured to  $\sim$ few percent

Next “structural” questions

Neutrino mass ordering ( $\nu$ MO)?



Leptonic  $CP$  violation?

$$\sin \delta_{CP} \neq 0$$

Maximal mixing?

$$\theta_{23} = 45^\circ$$
$$(|U_{\mu 3}| = |U_{\tau 3}|)$$

# Experimental overview

## 3-flavor oscillations

*Parameters best constrained by various sources*

Solar (MSW)	}	$\Delta m_{21}^2$	$\theta_{12}$						
Reactor ( $\sim 60$ km)									
Reactor ( $\sim 2$ km)	$\rightarrow$	$ \Delta m_{32}^2 $	$\theta_{13}$						
Atmospheric	$\rightarrow$	$ \Delta m_{32}^2 $	$\theta_{23}$	[ $\nu$ MO]					
Accelerator (LBL)	$\rightarrow$	$ \Delta m_{32}^2 $	$\theta_{23}$	$\nu$ MO	$\delta_{CP}$	[ $\theta_{13}$ ]			

$\rightarrow \sin^2\theta_{ij}$  and  $|\Delta m_{ij}^2|$  all measured to  $\sim$ few percent

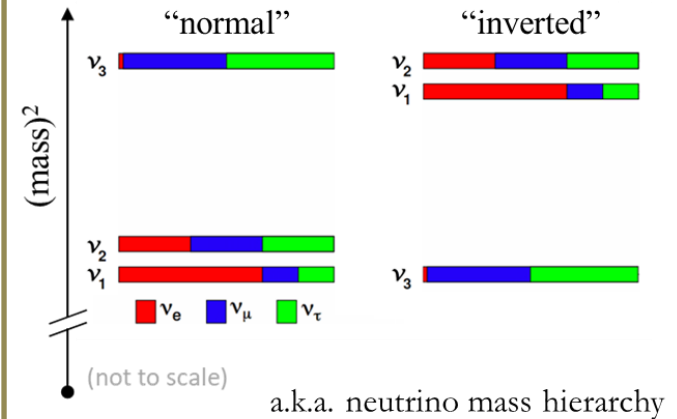
Needs:

- Address structural unknowns (e.g., at right)
- Testing mass/flavor models  $\leftrightarrow$  improved precision
- Direct unitarity tests (currently rather poor)

(And outside the scope of oscillation experiments:  $m_\nu$ ? Majorana or Dirac?)

Next “structural” questions

Neutrino mass ordering ( $\nu$ MO)?



Leptonic  $CP$  violation?

$$\sin \delta_{CP} \neq 0$$

Maximal mixing?

$$\theta_{23} = 45^\circ$$
$$(|U_{\mu 3}| = |U_{\tau 3}|)$$

# Experimental overview

## 3-flavor oscillations

*Parameters best constrained by various sources*

Solar (MSW) }  $\Delta m_{21}^2$   $\theta_{12}$   
Reactor (~60 km) }

Reactor (~2 km) →  $|\Delta m_{32}^2|$   $\theta_{13}$

Atmospheric →  $|\Delta m_{32}^2|$   $\theta_{23}$  [ $\nu$ MO]

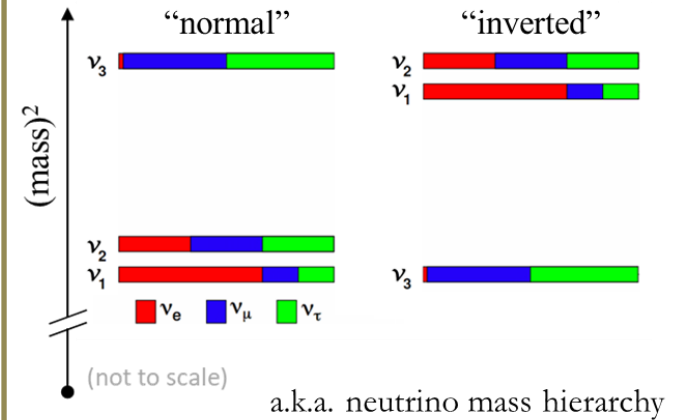
Accelerator (LBL) →  $|\Delta m_{32}^2|$   $\theta_{23}$   $\nu$ MO  $\delta_{CP}$  [ $\theta_{13}$ ]

These classes of experiment expected to have significant new oscillation results in the near term

*Will spend some time here before looking to the future...*

Next “structural” questions

Neutrino mass ordering ( $\nu$ MO)?



Leptonic  $CP$  violation?

$$\sin \delta_{CP} \neq 0$$

Maximal mixing?

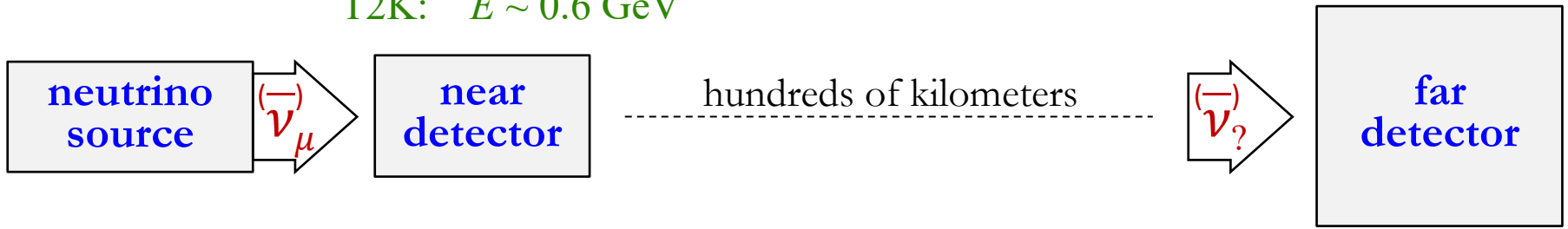
$$\theta_{23} = 45^\circ$$
$$(|U_{\mu 3}| = |U_{\tau 3}|)$$

# Accelerator-based long-baseline experiments

NO<sub>v</sub>A, T2K → DUNE, T2HK/Hyper-K  
(current) (next-gen)

# Sketch

★ NOvA:  $E \sim 2$  GeV  
T2K:  $E \sim 0.6$  GeV

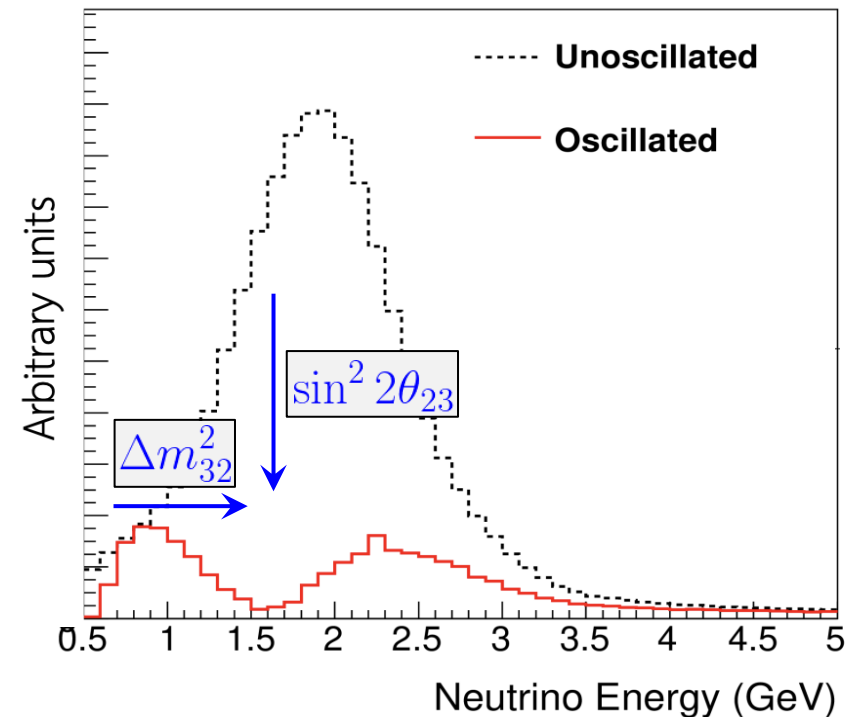


$\nu_{\mu}$  disappearance

$$P(\nu_{\mu} \rightarrow \nu_{e,\tau}) \approx \underbrace{\sin^2 2\theta_{23}}_{\text{experimental data are consistent with unity (i.e., maximal mixing)}} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right)$$

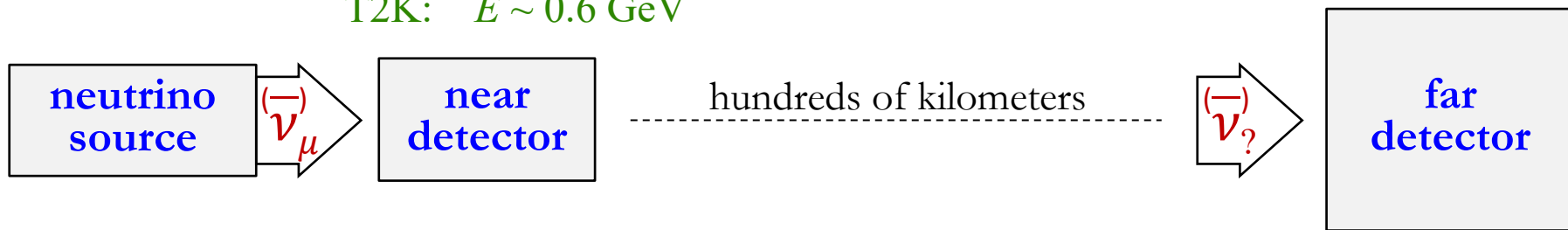
experimental data are **consistent with unity**  
(i.e., maximal mixing)

$\nu_{\mu}$  events in Far Detector  
(NOvA-like example)



# Sketch

★ NOvA:  $E \sim 2$  GeV  
T2K:  $E \sim 0.6$  GeV



## $\nu_\mu$ disappearance

$$P(\nu_\mu \rightarrow \nu_{e,\tau}) \approx \underbrace{\sin^2 2\theta_{23}}_{\text{experimental data are consistent with unity (i.e., maximal mixing)}} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right)$$

experimental data are **consistent with unity**  
(i.e., maximal mixing)

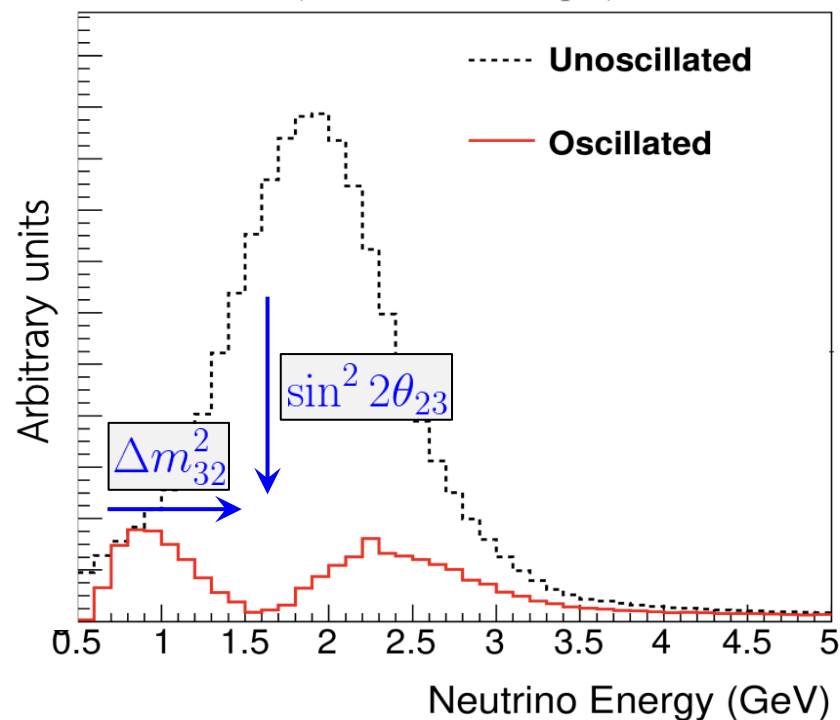
## $\nu_e$ appearance

$$P(\nu_\mu \rightarrow \nu_e) = \dots$$

...a much more involved function of the parameters of interest.

Significant connections to  $\theta_{13}$ ,  $\theta_{23}$ ,  $\delta_{CP}$ , and the **neutrino mass ordering\***

## $\nu_\mu$ events in Far Detector (NOvA-like example)

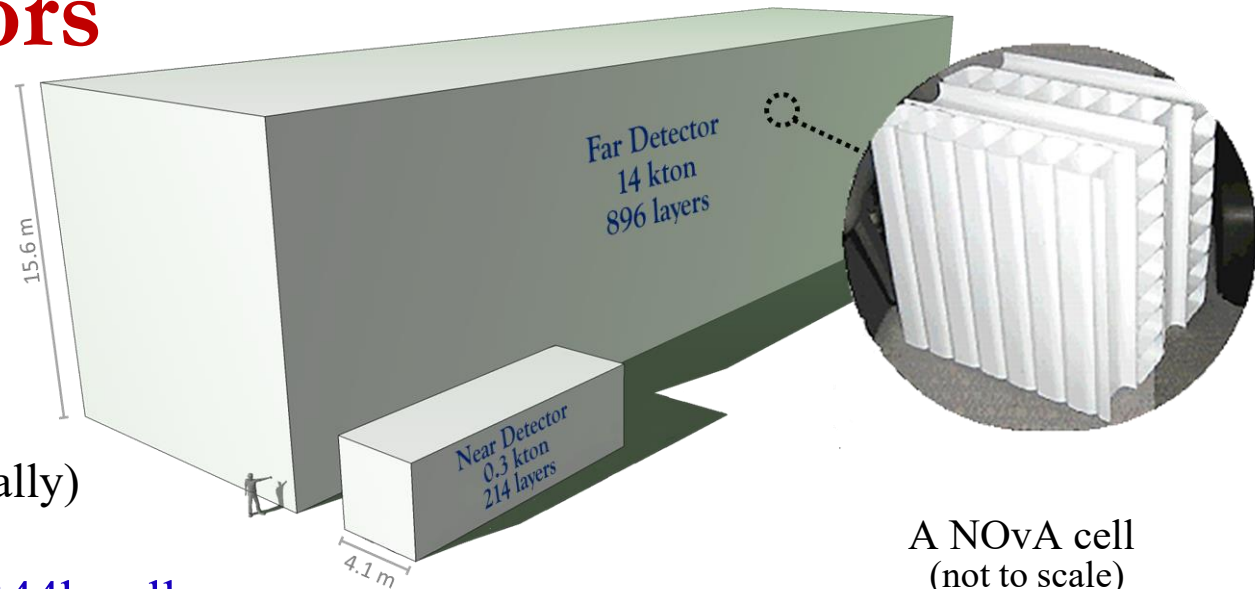


\*due to the neutrinos passing through matter



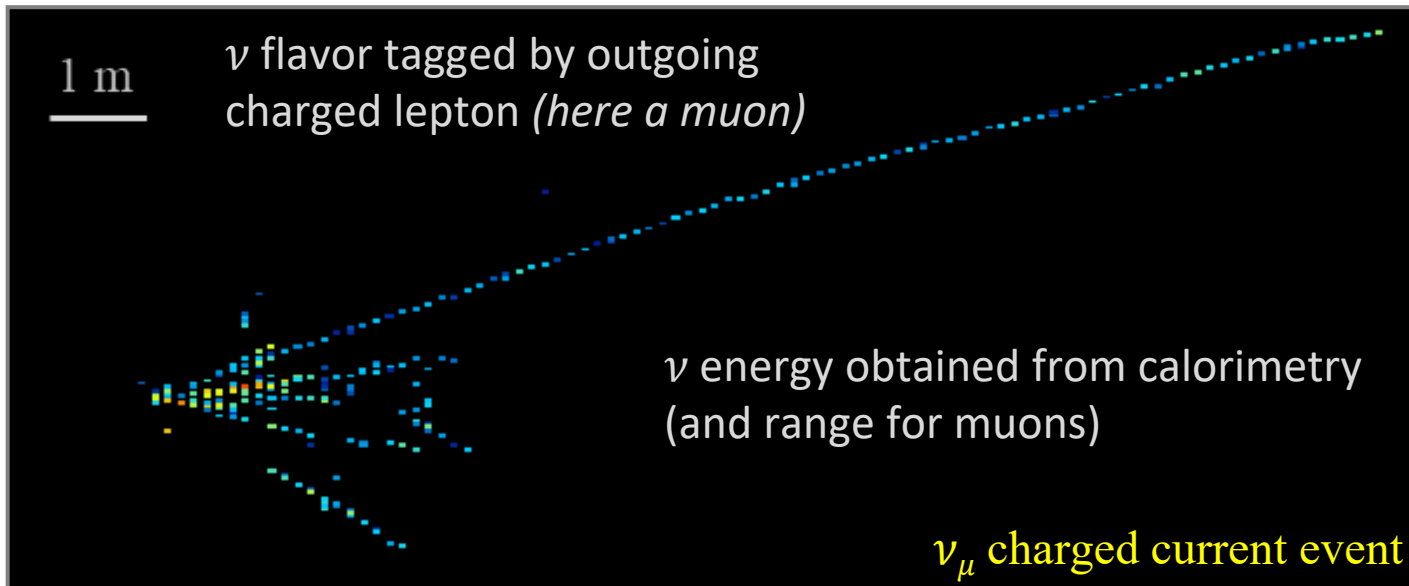
# NOvA detectors

- Liquid scintillator cells
- Fiber-to-APD readout
- **FD and ND functionally identical by design**  
→ Differ only in size (essentially)



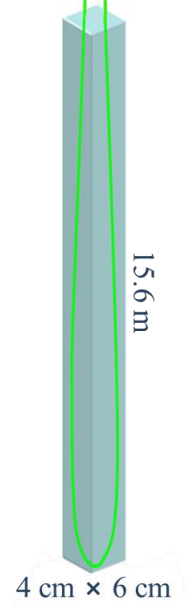
**Far detector:** 14 kton, 344k cells

**Near detector:** 0.3 kton, 20k cells



A NOvA cell  
(not to scale)

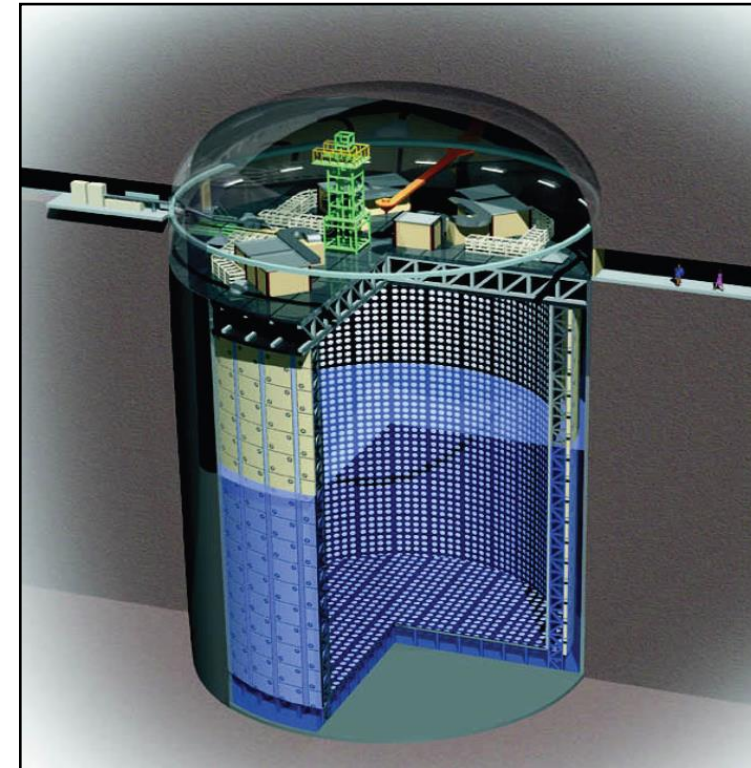
To APD



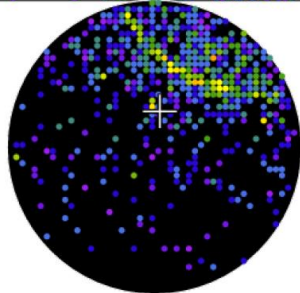
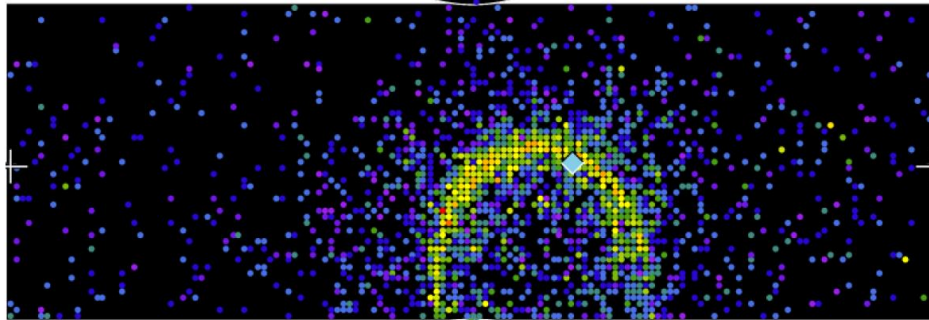
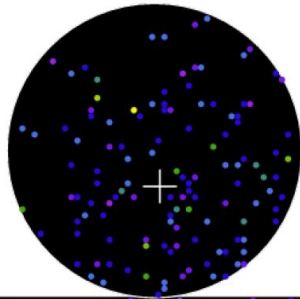
# T2K (far) detector

- The long-running **Super-Kamiokande** detector
- 50-kton water Cherenkov detector
- 11,000 20" PMTs on inner surface (40% coverage)

Super-Kamiokande



Super-K  
 $\nu_e$  CC event

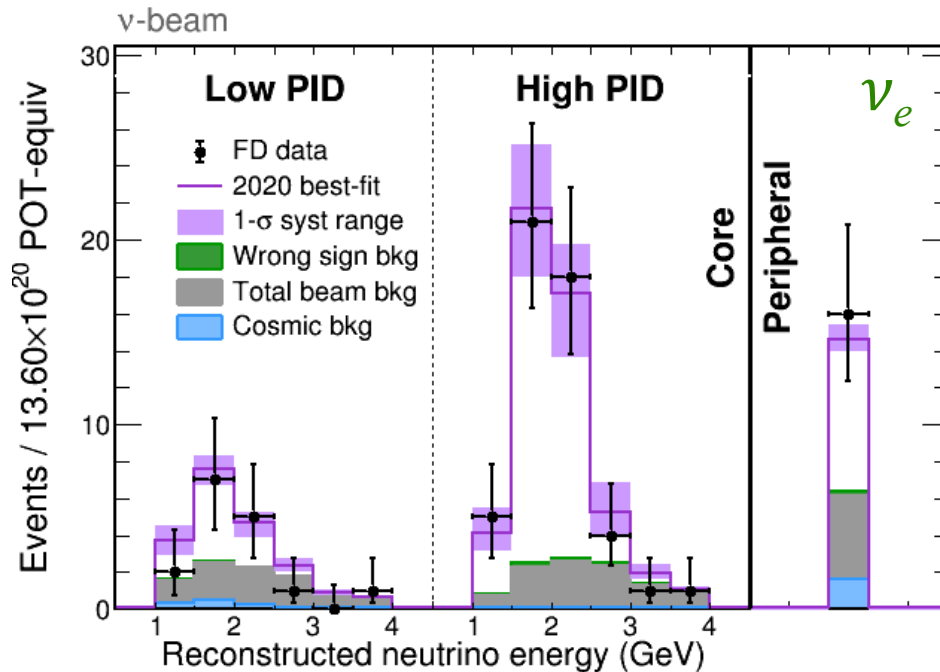
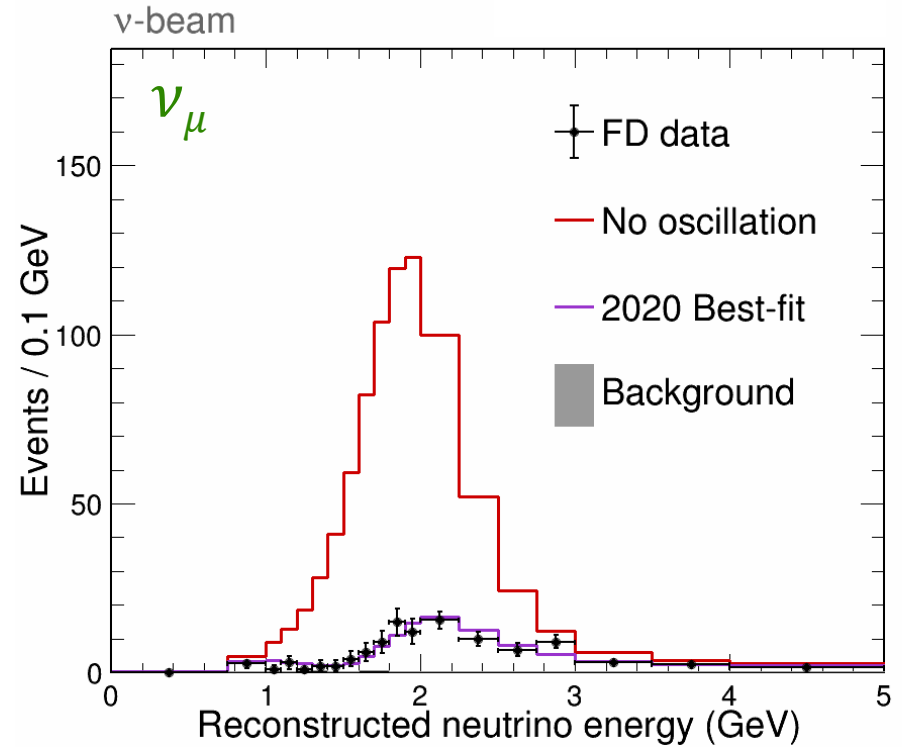


- **$e/\mu$  discrimination** from ring characteristics
- **$E_\nu \sim 0.6$  GeV**
  - $\Rightarrow$  relatively simple final states  
(*e.g.*,  $\mu + p$  or  $\mu + p + \pi$ )
  - $\Rightarrow$  protons below Cherenkov threshold
- **$E_\nu$  from reconstructed lepton momentum** (and kinematics)

# NOvA FD data

**At right:**  $\nu_\mu$  CC candidate events (“disappearance”)

**Below:**  $\nu_e$  CC candidate events (“appearance”)



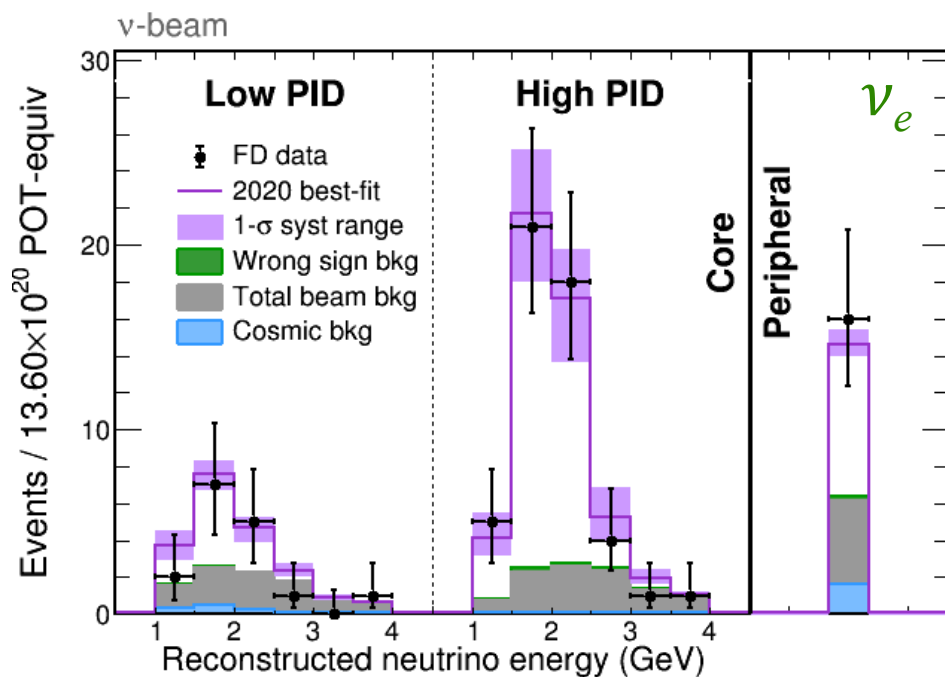
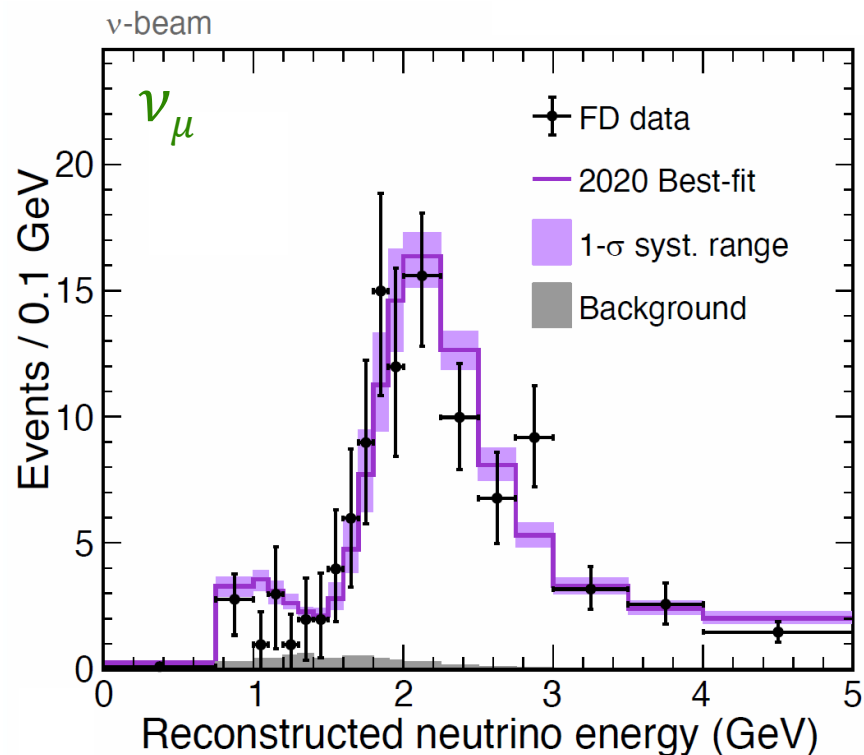
Not shown, but important pieces of oscillation fit:

- $\bar{\nu}_\mu$  and  $\bar{\nu}_e$  samples!
- Breakdown of  $\nu_\mu/\bar{\nu}_\mu$  samples by hadronic energy fraction

# NOvA FD data

**At right:**  $\nu_\mu$  CC candidate events (“disappearance”)

**Below:**  $\nu_e$  CC candidate events (“appearance”)



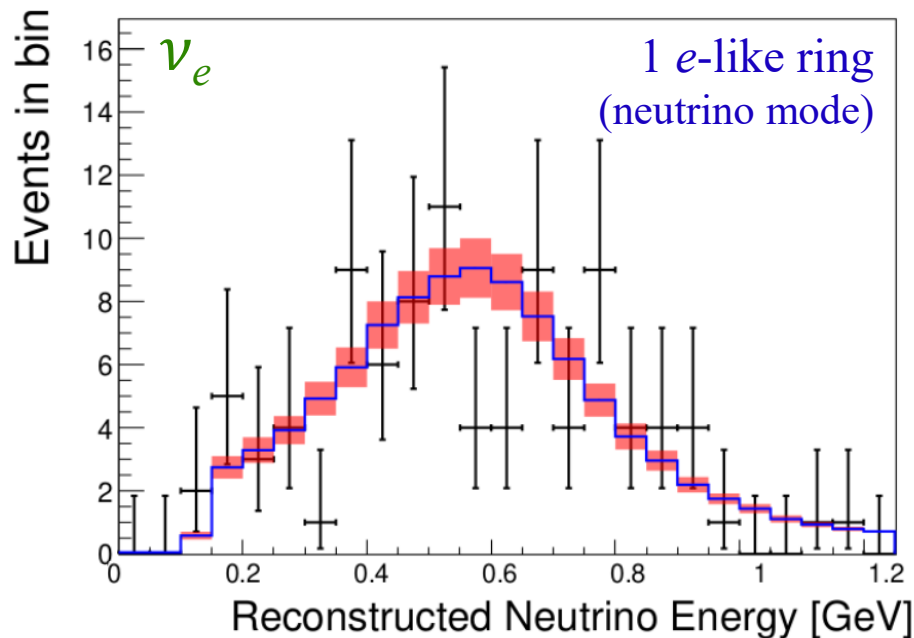
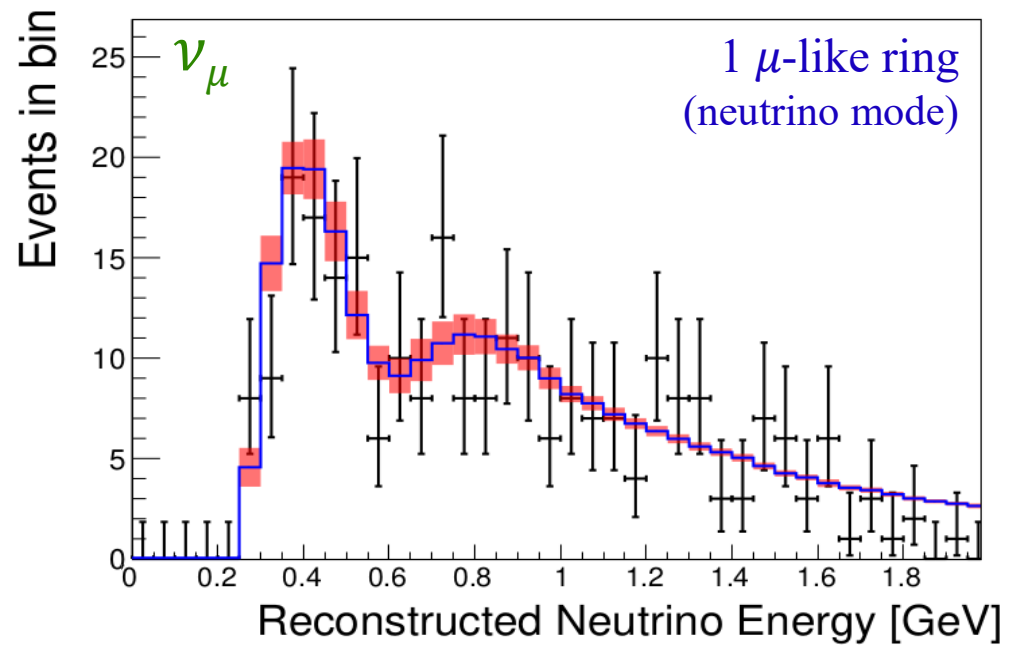
Not shown, but important pieces of oscillation fit:

- $\bar{\nu}_\mu$  and  $\bar{\nu}_e$  samples!
- Breakdown of  $\nu_\mu/\bar{\nu}_\mu$  samples by hadronic energy fraction

# T2K FD data

**At right:**  $\nu_\mu$  CC quasi-elastic candidate events

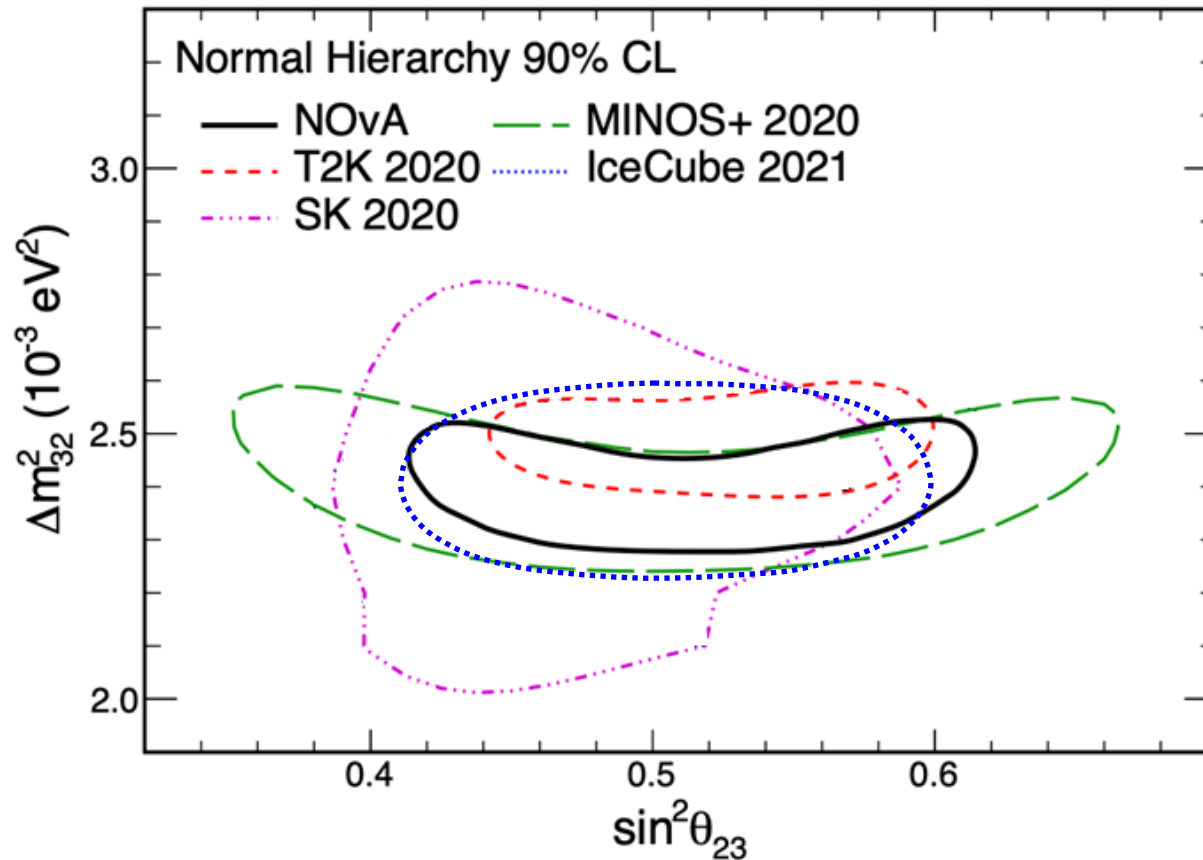
**Below:**  $\nu_e$  CC quasi-elastic candidate events



Not shown, but important pieces of oscillation fit:

- $\bar{\nu}_\mu$  and  $\bar{\nu}_e$  samples!
- Additional “1  $e$  with Michel” sample ( $e^- + \pi^+$  final state)

# $\Delta m_{32}^2$ and $\theta_{23}$



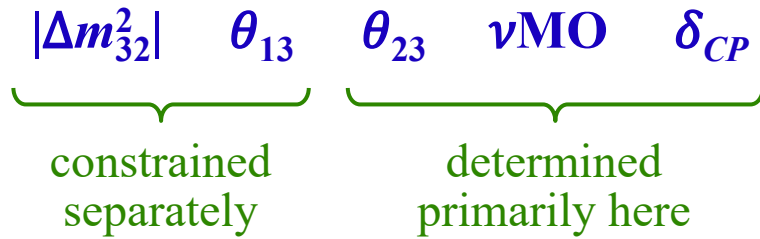
Will come back to  
the atmospheric  $\nu$   
experiments.

Daya Bay reactor  
experiment also  
impactful in  $|\Delta m_{32}^2|$ .

- $|\Delta m_{32}^2|$  **precision:** 1.1% on global average given a mass ordering
- Still no evidence of deviation from **maximal mixing** ( $\sin^2 \theta_{23} = 0.5$ )

# LBL $\nu_e$ appearance

$P(\nu_\mu \rightarrow \nu_e)$  and  $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$  depend on:



**Measurement ambiguities possible**

Complementarity between T2K and NOvA

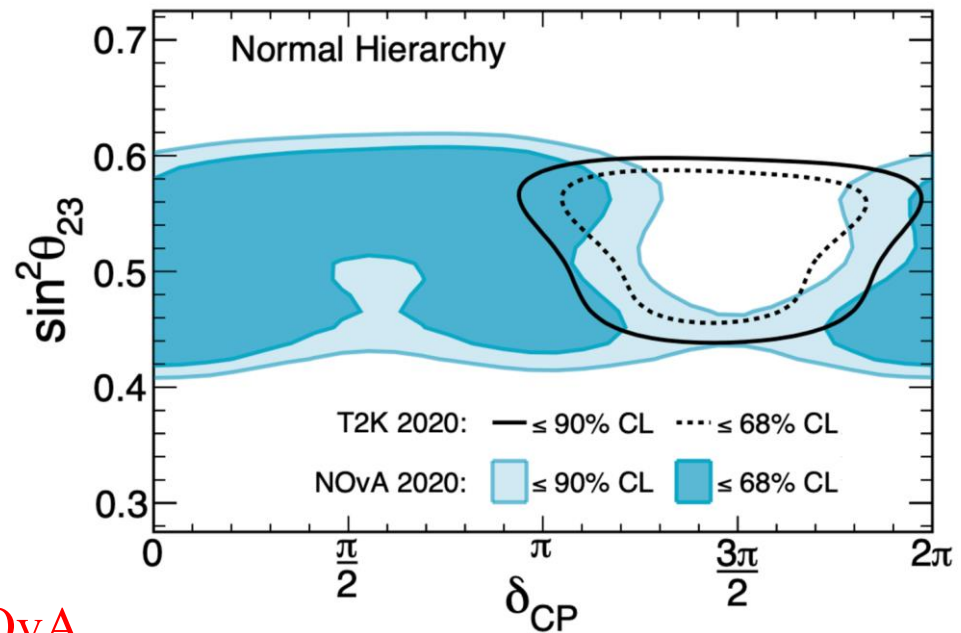
**T2K:** minimal  $\nu_{MO}$  sensitivity

**NOvA:** good  $\nu_{MO}$  sensitivity

Inter-collaboration joint fits underway:

NOvA + T2K and also T2K + SuperK\*

Results expected later this year



Current best fits land in the **normal ordering**, but only at **polar opposite  $\delta_{CP}$  values**

In slightly less favored inverted ordering, **good agreement in  $\delta_{CP}$**

**>2.5× increase in LBL exposures anticipated (vs. currently analyzed) through *c.* 2027**

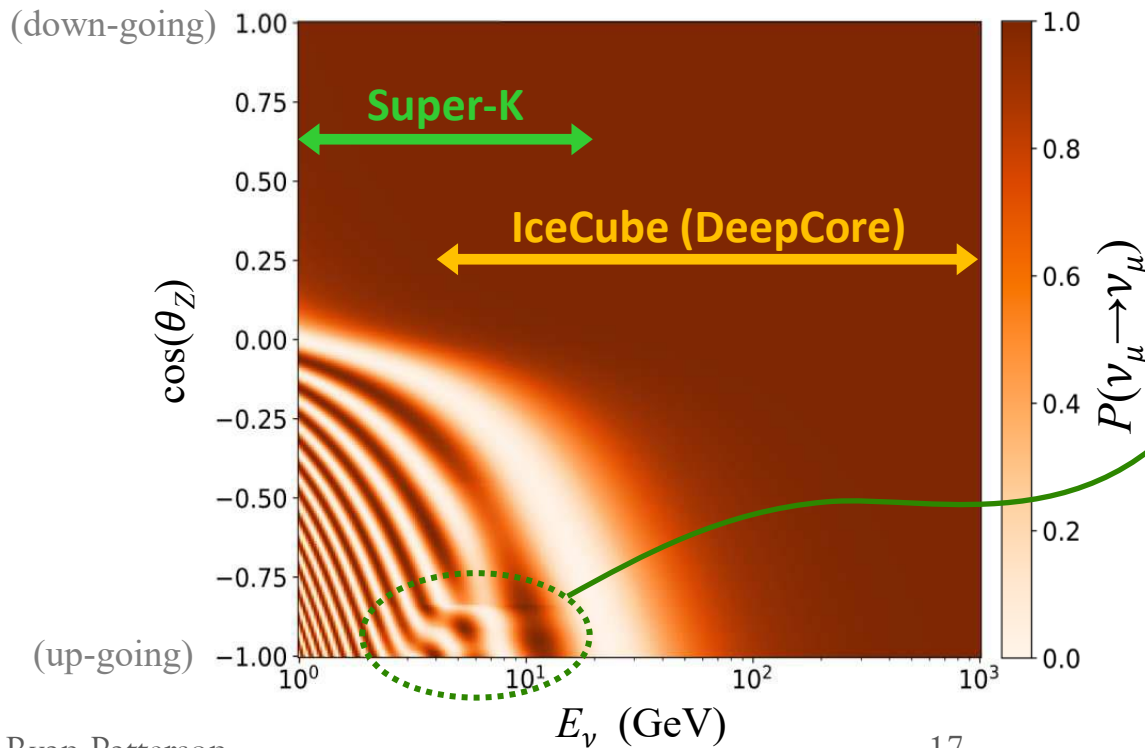
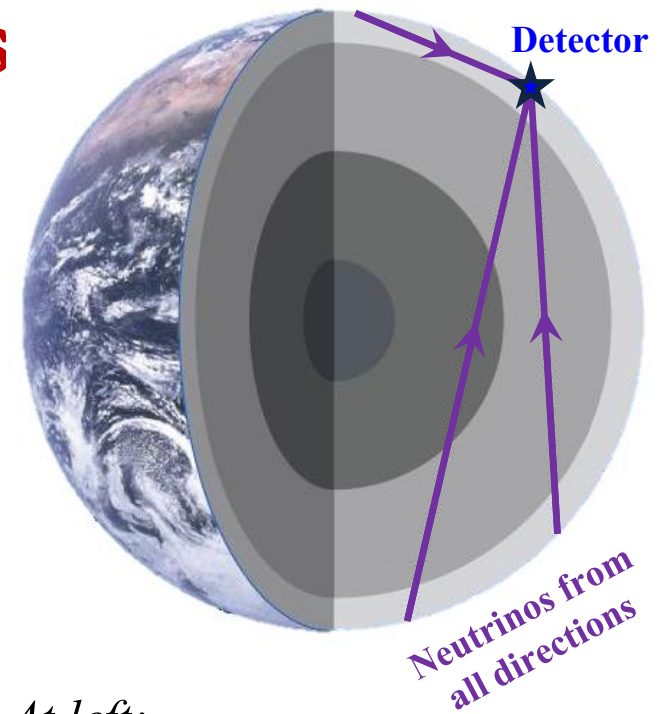
\* SuperK atmospheric  $\nu$  data:  $1.3\sigma$  preference for normal ordering

# Oscillations with atmospheric neutrinos



# Atmospheric $\nu$ oscillations

- $\nu$  from cosmic ray interactions in atmosphere  
Relevant energies:  $\sim 1 - 1000$  GeV
- **Broad spectrum**, mixture of initial flavors
- Baseline varies with  $\nu$  incident angle  
 $L \sim 10^1$  to  $10^4$  km
- $\nu_\mu \rightarrow \nu_\mu$  and  $\nu_\mu \rightarrow \nu_e$  and  $\nu_\mu \rightarrow \nu_\tau$  (above  $\tau$  prod. threshold)



At left:

**Regular oscillation pattern**

*plus*

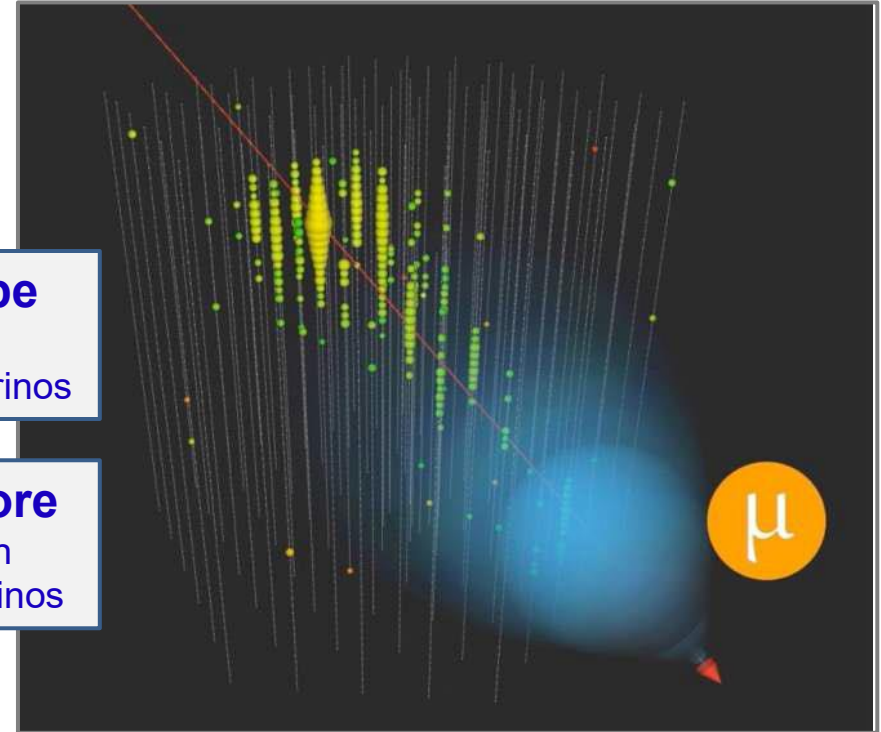
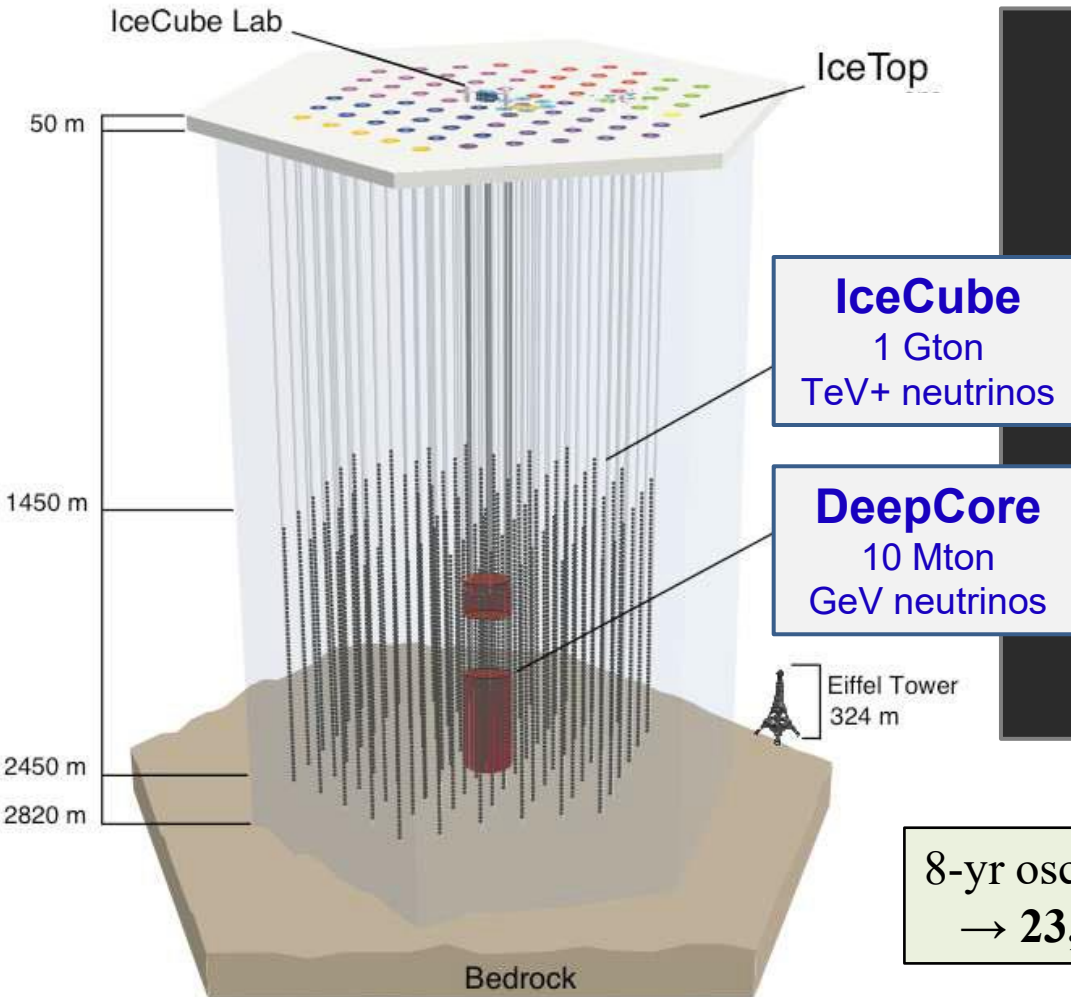
**Localized distortion** due to resonant matter effects  
 $\Rightarrow \nu$  mass ordering

$E_\nu$ ,  $\theta_z$  smearing is **significant**  
(not shown in this cartoon!)

# Super-K detector described already

- *Underway*: **Gd doping** → improved neutron tagging ( $\nu/\bar{\nu}$  separation)

## IceCube / DeepCore



8-yr oscillation data set (*results shown earlier*)  
→ **23,000** selected events!

# DeepCore analysis improvements (c. 2022)

- **>200k neutrinos**, all flavors. Competitive oscillation sensitivity w/ LBL expts.
- **$\nu_\tau$  appearance**: 10k events, expecting 11% precision on normalization (PMNS unitarity)

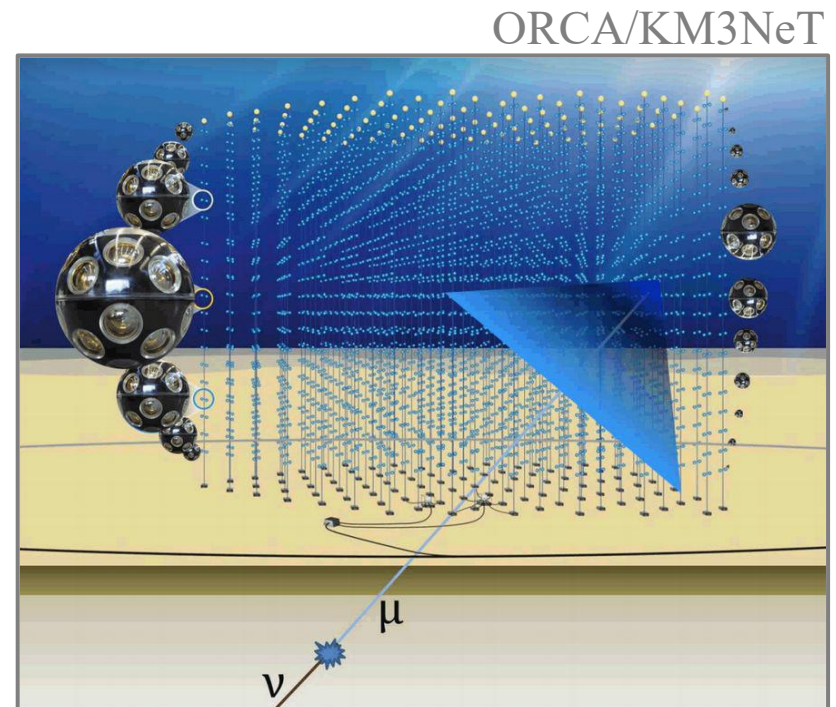
*A bit later on the timeline...*

## IceCube Upgrade

- **2 Mton sub-array** in 10 Mton DeepCore (and other improvements)
- **Drop threshold** to 1 GeV

## ORCA/KM3NeT

- **7 Mton PMT array** in Mediterranean
- Optimized for  $\sim 1 - 100$  GeV
- *Currently*: Preliminary operations with 5% deployment



Competitive  $\nu$  mass ordering reach ultimately (few to several  $\sigma$ )

# Synthesizing $\sim 5$ yr projections (so, *c.* 2027)

**Precision on  $\theta_{12}$ ,  $\theta_{13}$ ,  $\Delta m_{21}^2$**

→ Minimal changes until next-gen experiments (*e.g.*, *JUNO*)

**Precision on  $\theta_{23}$ ,  $|\Delta m_{32}^2|$**

→ Some gains to come in current generation. Large gains in next-gen.

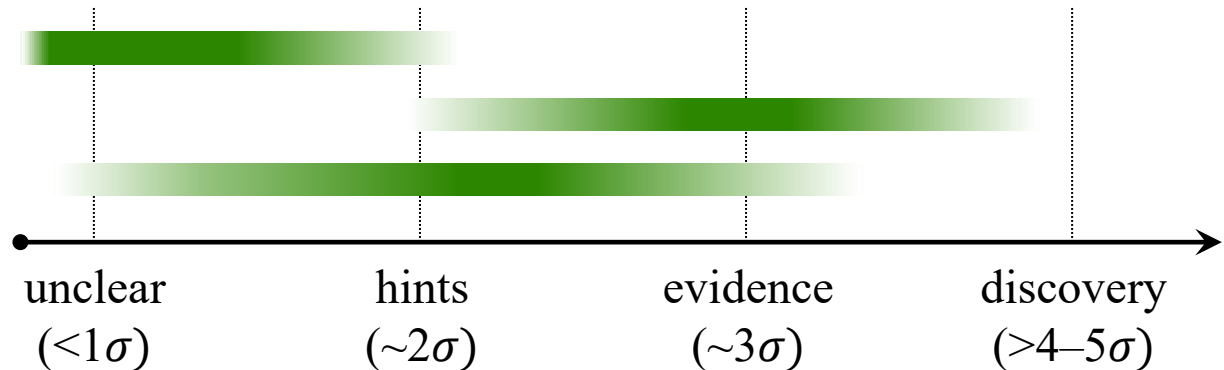
## ★ 3-flavor “structural” questions

→ Reach heavily depends on (*still unknown!*) actual answers

$\theta_{23}$  octant / max. mixing?

$\nu$  mass ordering?

$\nu$   $CP\nu$ ?



(A qualitative sketch.  
Don't try to read precise  
numbers off this diagram!)

# Synthesizing ~5 yr projections (so, c. 2027)

Precision on  $\theta_{12}, \theta_{13}, \Delta m_{21}^2$

**All possible outcomes have natural next steps**

if ambiguities → **aim for clarity**

if evidence → **aim for discovery**

if discovery → **aim for characterization**

...and ultimately toward the precision era in leptonic flavor



3-fla

$\theta_{23}$  octant

$\nu$  mass ordering?

$\nu$  CPv?

(A qualitative sketch.  
Don't try to read precise  
numbers off this diagram!)

unclear  
( $<1\sigma$ )

hints  
( $\sim 2\sigma$ )

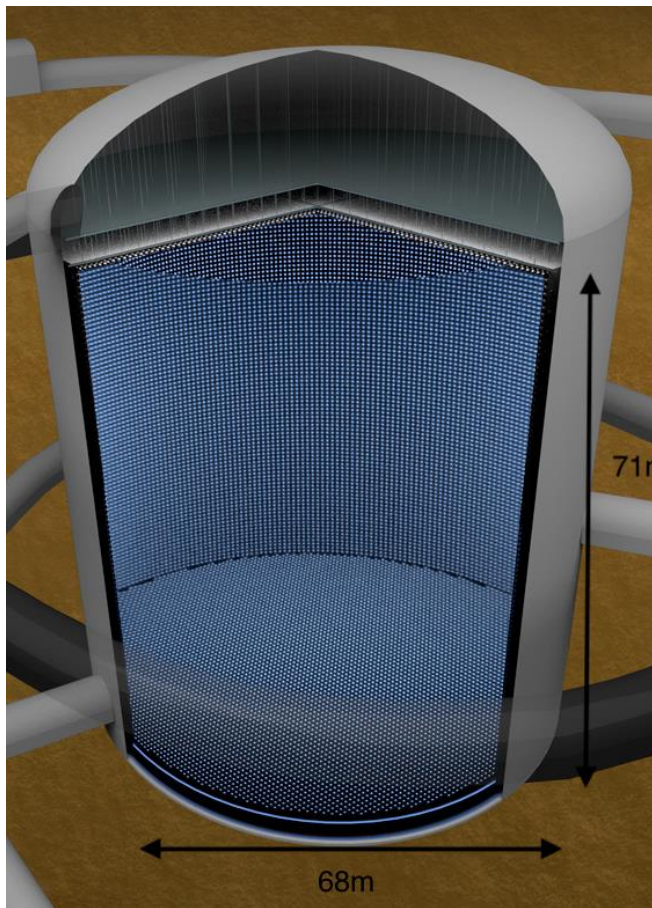
evidence  
( $\sim 3\sigma$ )

discovery  
( $>4-5\sigma$ )

**The next steps...**

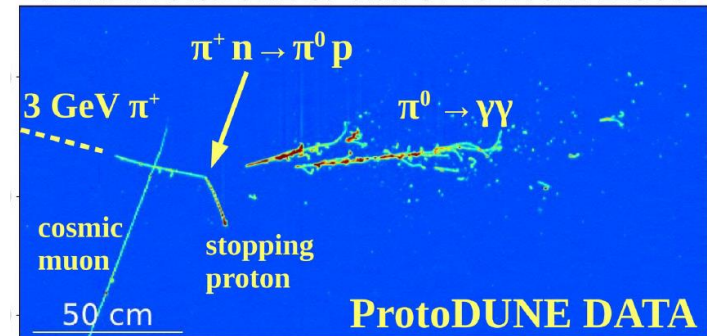
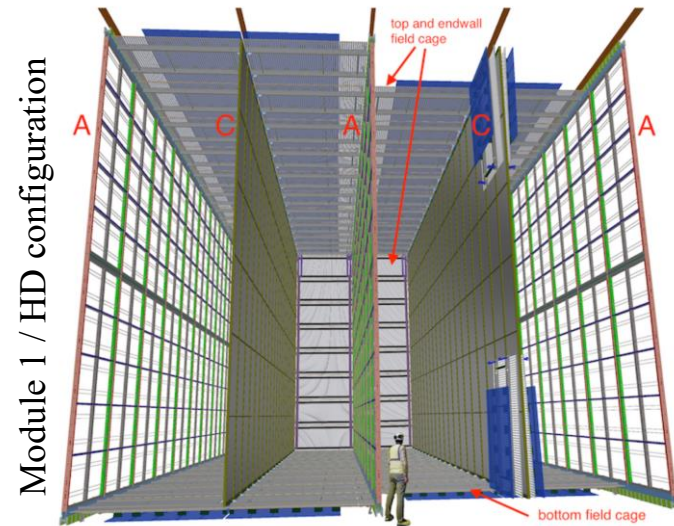
# Next-generation LBL experiments

## Hyper-K / T2HK



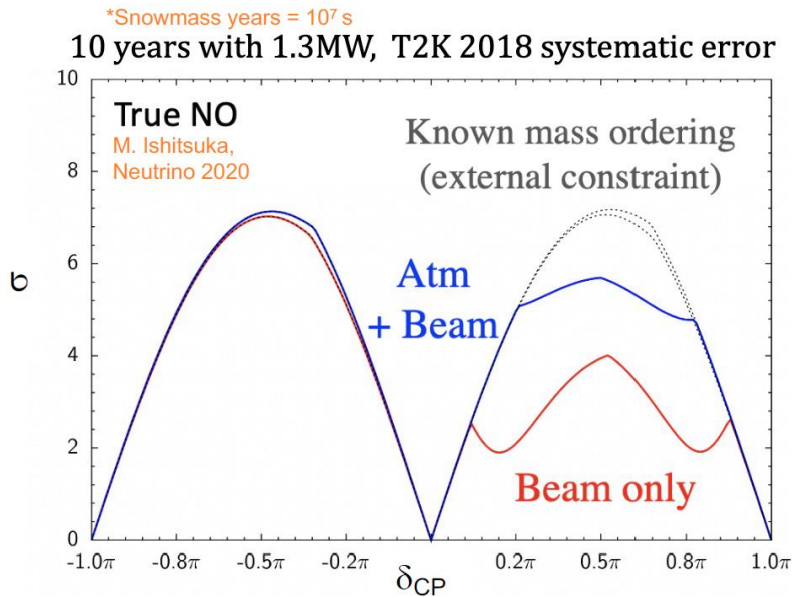
- 187-kt fid. volume water Cherenkov detector
- 1.3 MW off-axis beam from J-PARC
- 295 km oscillation baseline

## DUNE



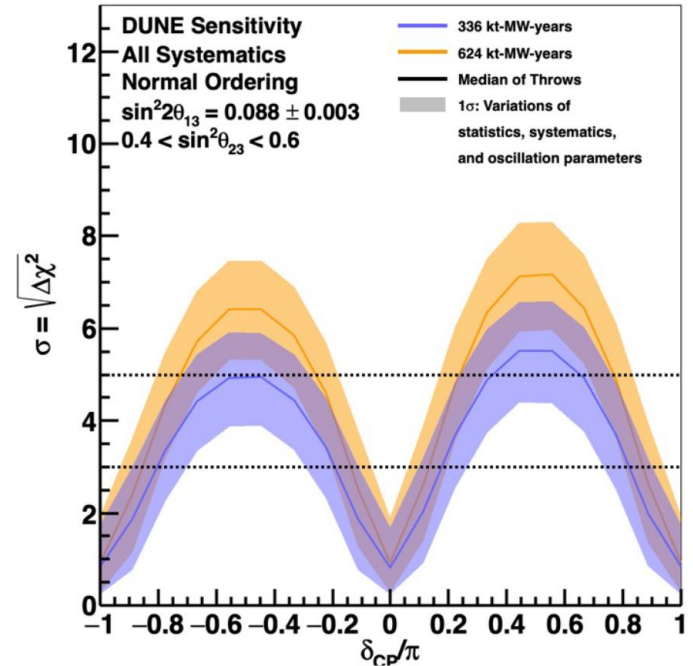
- >40-kt fiducial volume liquid argon TPC
- 1.2 – 2.4 MW broad-band beam from FNAL
- 1300 km oscillation baseline

# Hyper-K / T2HK



- Same baseline as T2K. Can use atmospheric data to help with mass ordering.

# DUNE



- Long baseline → **Definitive ( $>5\sigma$ ) mass ordering determination quickly** for all parameter space

**CPv reach**

*looking at both...*

$5\sigma$  CPv reach for  $>50\%$  of parameter space. Precision PMNS measurements.

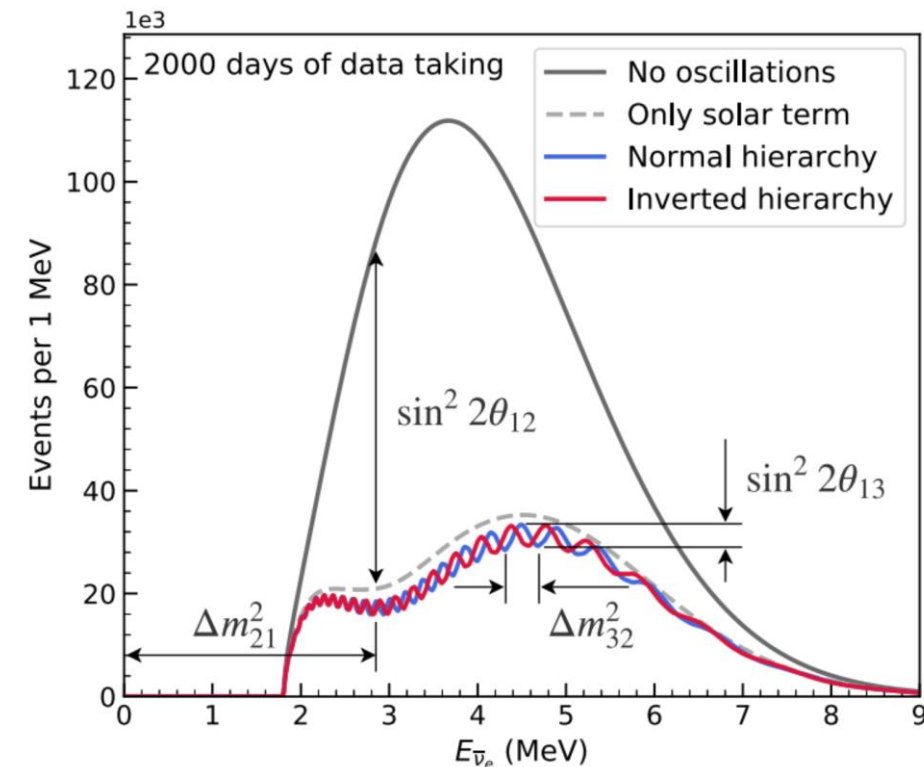
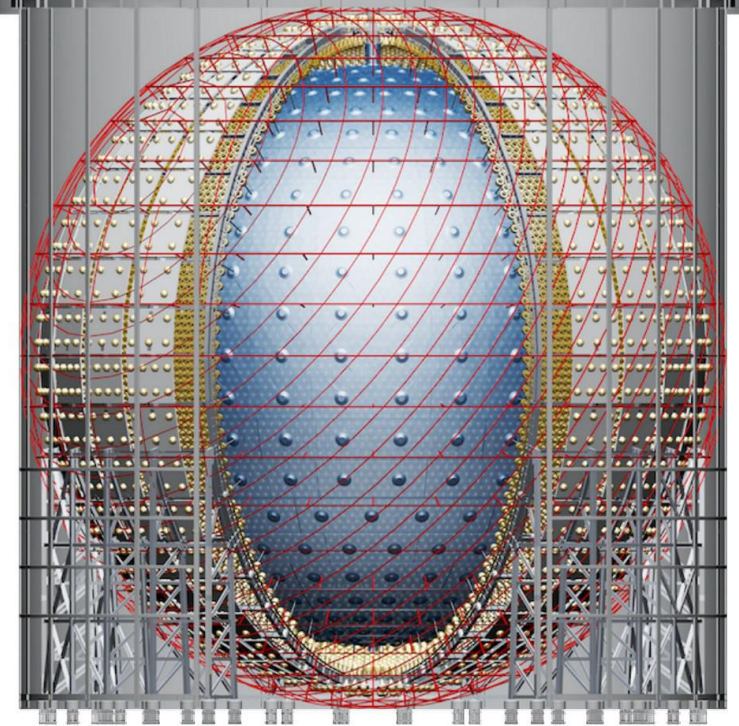
Civil construction underway. **First physics data toward end of this decade.**

**Highly complementary** in detector design, systematics mitigation, and broader science program



# JUNO

- 20 kton open-volume **liquid scintillator detector**
- **Oscillations at  $L \approx 50$  km with reactor  $\bar{\nu}_e$**   
(Plus atmospheric, solar, geo, supernova  $\nu$ )
- Designed for **highly efficient light collection**  
*strong requirements on energy resolution (3%)  
and linearity/scale uncertainties (<1%)*



- **Sub-percent measurements of  $\theta_{12}$ ,  $|\Delta m_{ij}^2|$**
- Novel approach to **mass ordering**:  
→ Direct spectral measurement of all  $\Delta m_{ij}^2$ -driven oscillation contributions
- **Construction well underway**. Start of operations planned for 2023  
→  $3\sigma$   $\nu MO$  by end of decade

# Closing

- Several experiments with a range of time scales are poised to provide **the next answers in the neutrino sector**.
- In this short talk, I have **only scratched the surface** of the broad programs of the experiments discussed here.  
(*e.g.*, measurements related to astrophysics, baryon number violation, sterile neutrinos, dark matter, and other BSM physics)

