

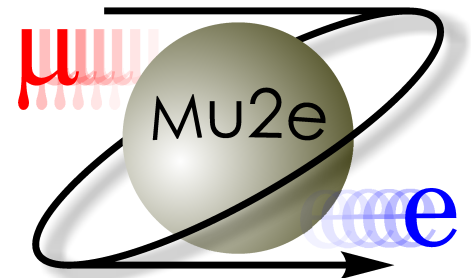


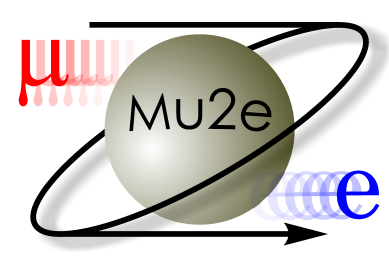
## A Search for New Physics in the Lepton Sector: Charged Lepton Flavor Violation and the Mu2e Experiment

Mete Yucel on behalf of the Mu2e Collaboration

FPCP 2022

5/25/2022

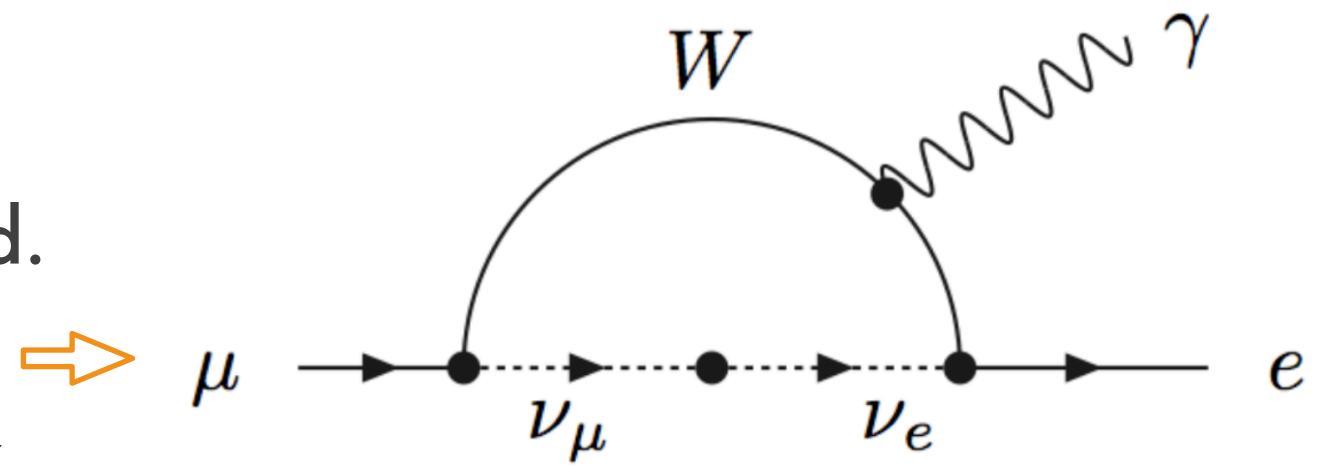




# Motivation - What is CLFV ?

## CLFV (Charged Lepton Flavor Violation)

- Quarks mix, neutrinos mix, why don't we observe charged leptons mixing?
- Charged lepton flavor is not conserved.
  - As neutrino masses indicate.
- Let's look at SM process;  $\mu^\pm \rightarrow e^\pm + \gamma$   
 Let's look at BR result for this process



$$\mathcal{B}(\mu \rightarrow e\gamma) \sim \mathcal{O} 10^{-54}$$

Heavily suppressed in SM, perfect for searching for new physics !!!

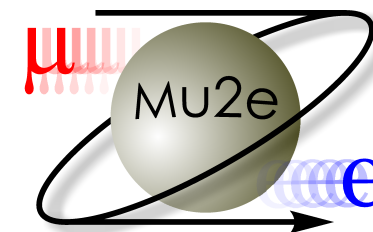
- Searching for CLFV in the muon sector;

Experiment	Institute	Process
MEG II	PSI	$\mu^\pm \rightarrow e^\pm + \gamma$
<b>Mu2e</b>	<b>FNAL</b>	<b><math>\mu^- + N \rightarrow e^- + N</math></b>
COMET	JPARC	$\mu^- + N \rightarrow e^- + N$
Mu3e	PSI	$\mu^\pm \rightarrow e^\pm + e^+ + e^-$

Mu2e focuses on the neutrino-less conversion of the muon in the presence of Al nucleus.



# Motivation - What mass scale( $\Lambda$ ) Mu2e probes ?



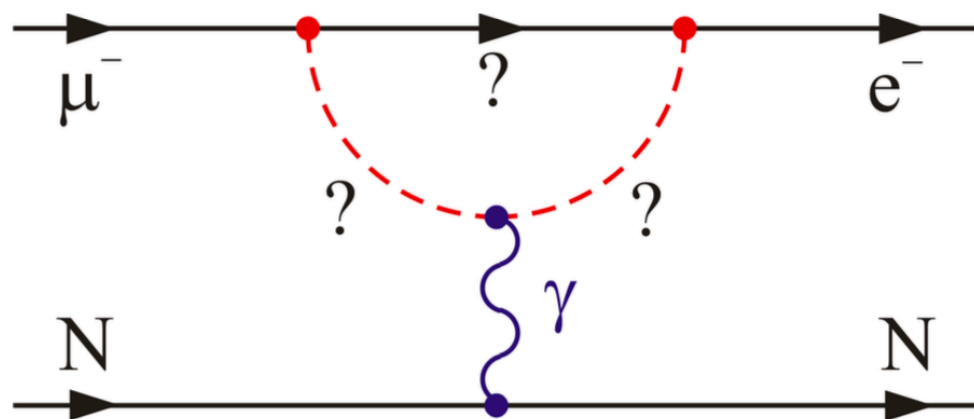
$$\mathcal{L}_{CLFV} = \frac{m_\mu}{(\kappa + 1)\Lambda^2} \bar{\mu}_R \sigma_{\mu\nu} e_L F^{\mu\nu} + \frac{\kappa}{(\kappa + 1)\Lambda^2} \bar{\mu}_L \gamma_\mu e_L (\bar{e} \gamma^\mu e)$$

A. De Gouvea and P. Vogel; [arXiv:1303.4097](https://arxiv.org/abs/1303.4097)

Lower  $\kappa$  is sensitive to the **loop** contributions to the  $\mathcal{L}_{CLFV}$

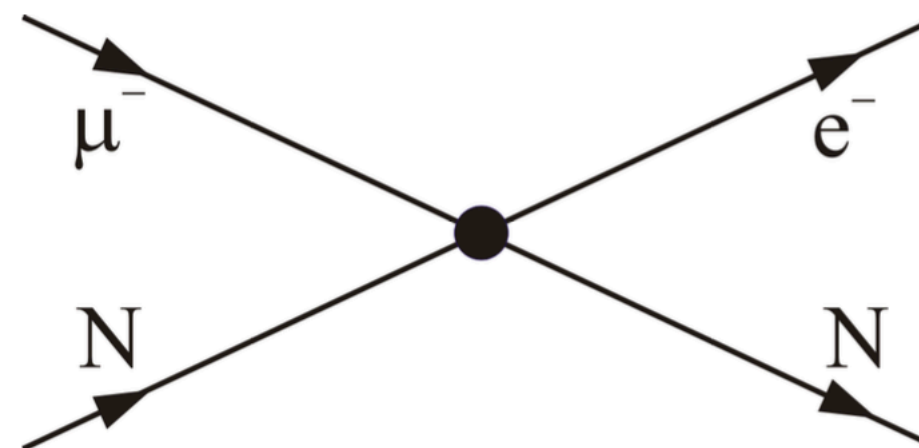
Higher  $\kappa$  is sensitive to the **contact term** of the  $\mathcal{L}_{CLFV}$

$\kappa = 0$  ←



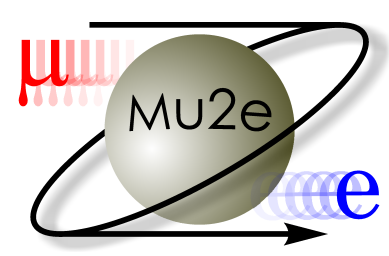
Supersymmetry & Heavy neutrinos  
 $\mu \rightarrow e\gamma$  contribution

→  $\kappa = \infty$

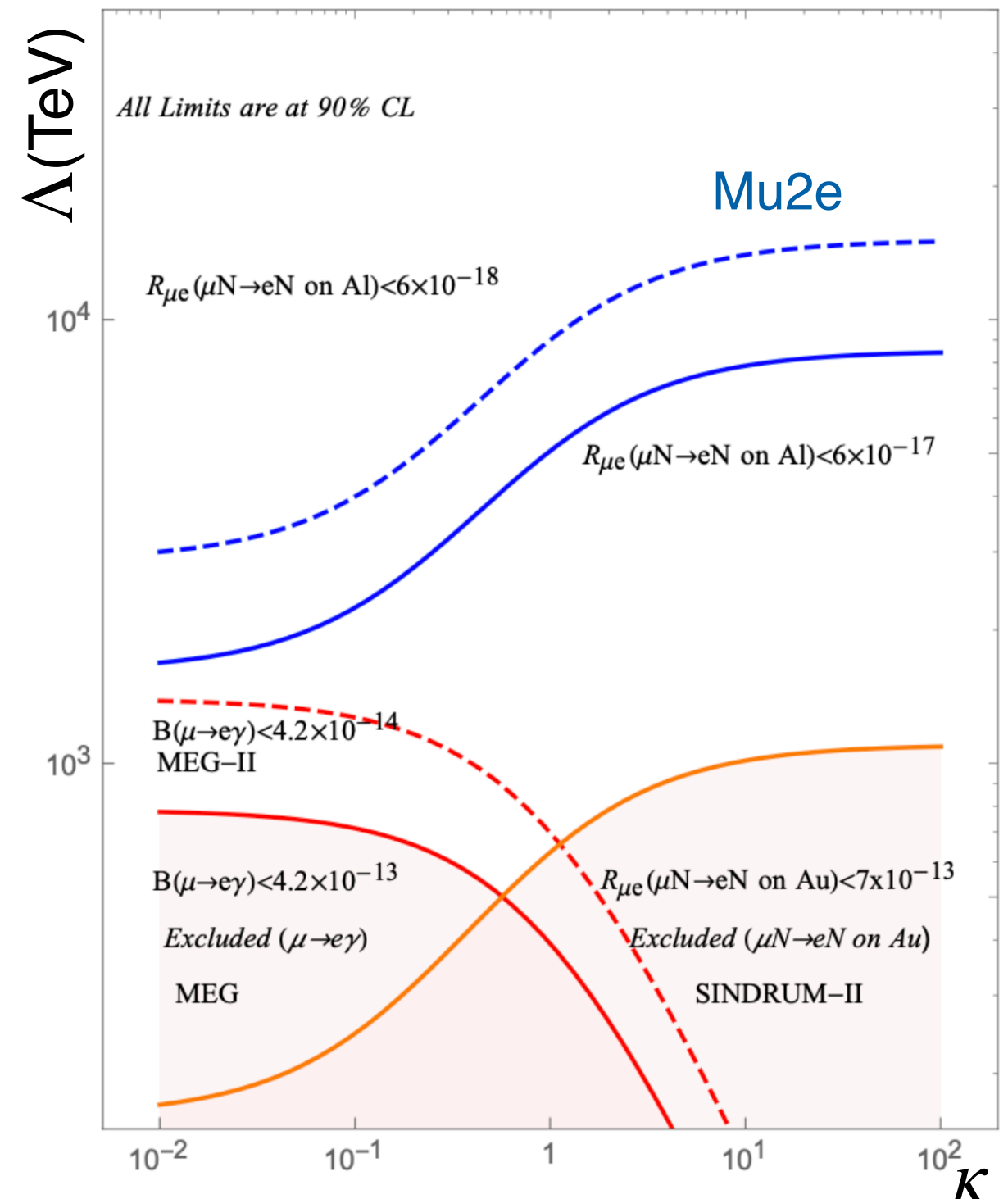


Leptoquarks, heavy Z ...  
No contribution from  $\mu \rightarrow e\gamma$

# Motivation - Why is Mu2e unique?



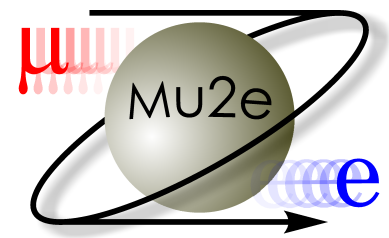
- Mu2e probes  $\Lambda$ (mass scale) up to  $10^4$  TeV
- Advantage over collider experiments on probing rare process;
  - Free of SM backgrounds
  - Intense muon beams for high statistics.
  - High sensitivity to couplings.



A. De Gouvea and P. Vogel; [arXiv:1303.4097](https://arxiv.org/abs/1303.4097)



# Mu2e signal



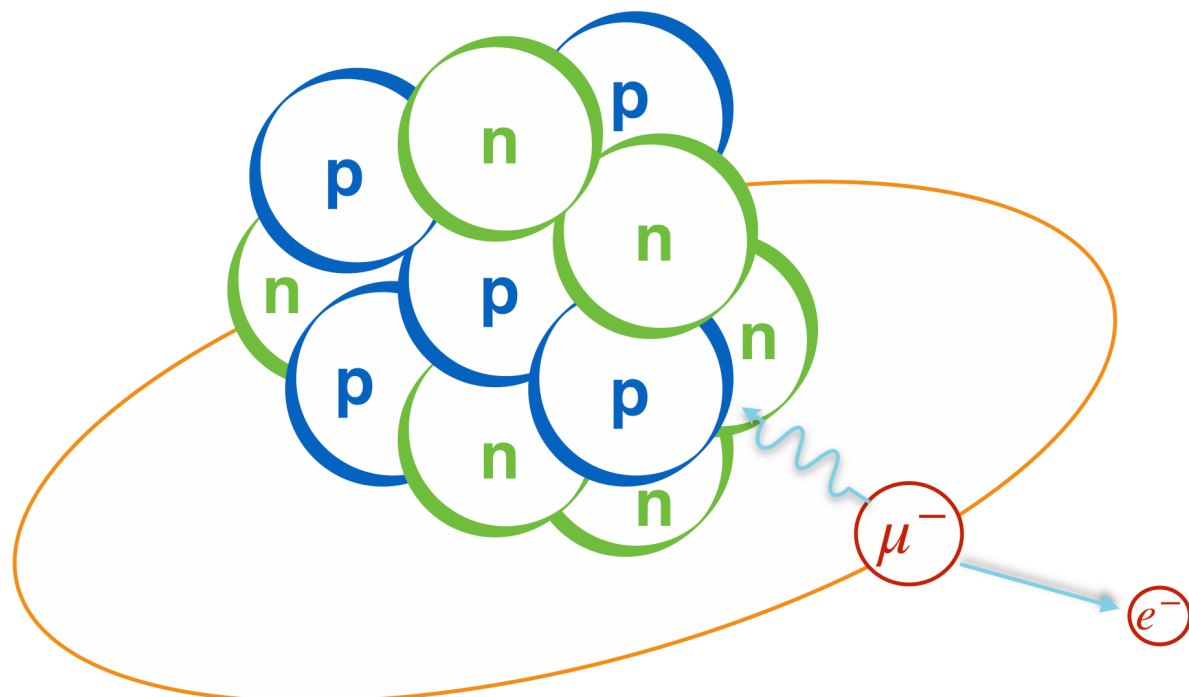
- Physics signal properties;
  - Coherent electron conversion
  - Little energy lost to
    - $\mu$  atomic binding energy  $E_b = 0.48$  MeV
    - Nuclear recoil  $E_R = 0.21$  MeV
  - **No neutrinos** are produced
  - **Monoenergetic**  $e^-$  is **104.97 MeV**

## Coherent electron conversion with Al



$$E_{e^-} = M_{\mu^-} - E_b - E_{recoil} = 104.97 \text{ MeV}$$

### Aluminum nuclei



Muonic Al lifetime = 864 ns

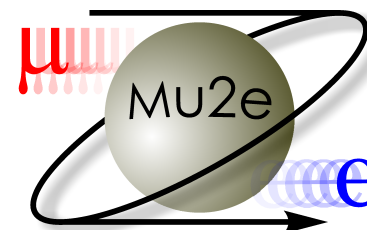
## Conversion rate

$$R_{\mu e} = \frac{\Gamma(\mu^- N \rightarrow e^- N)}{\Gamma(\mu^- N \rightarrow \text{All captures})}$$

Mu2e goal =  $3 \times 10^{-17}$  SES

$\times 10^4$  improvement over SINDRUM-II

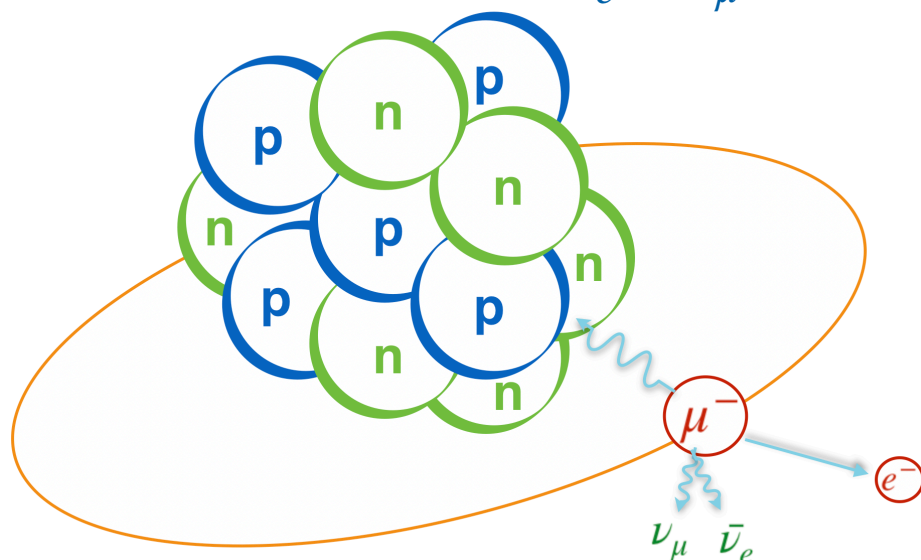
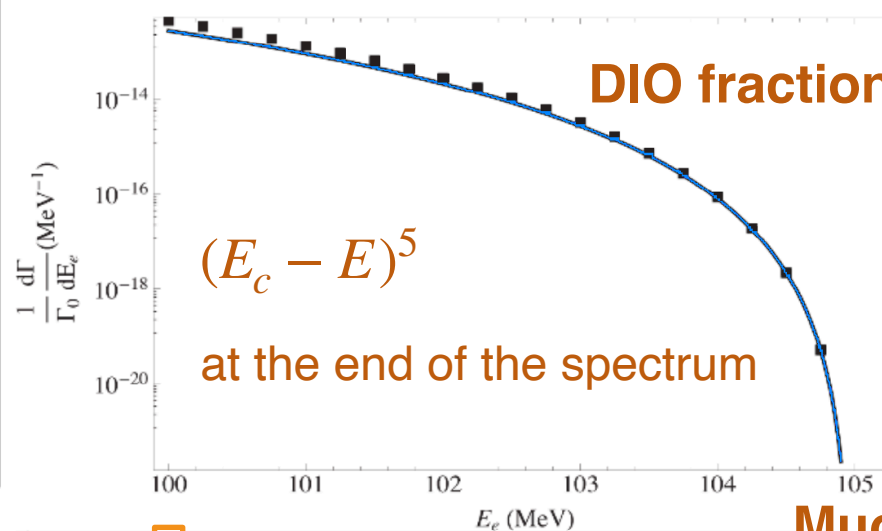
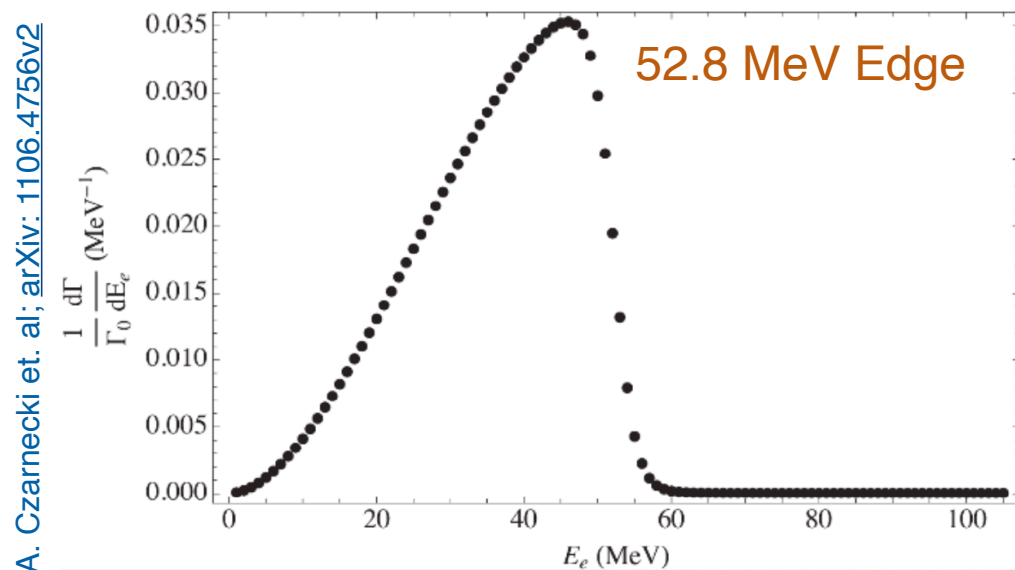




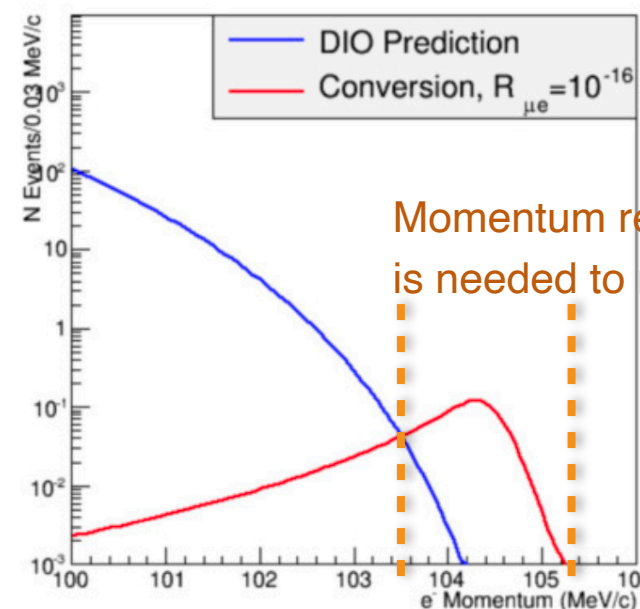
# Backgrounds - DIO

- Decay In Orbit(DIO)

Michel spectrum ( $\mu$  decay in orbit)

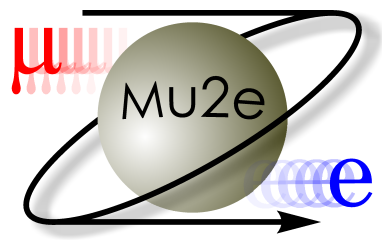


Plot conversion  $e^-$  on top  
After Reco Acceptance+  $\Delta E$ +Resolution



Momentum res. < 180 keV/c  
is needed to make the separation

# Design principles



How to get  $10^4$  improvement over SINDRUM II & SES  $3 \times 10^{-17}$

1. High intensity pulsed muon beam.

- High statistics.
- Introduces beam related backgrounds !!!

2. High resolution on the momentum

- Separation of 105 MeV/c conversion  $e^-$  from DIO  $e^-$  tail.
  - Low mass straw drift detector with 180 keV/c resolution for 105 MeV/c  $e^-$ .
  - Couple with EM calorimeter to complement tracker.
3. Background suppression of  $<1$  event for the experiment.
- Blind to low momentum particles.
  - $10^{-10}$  extinction factor for out-of-time protons.
  - Event window separation with pulsed muon beam.
  - Cosmic ray veto.

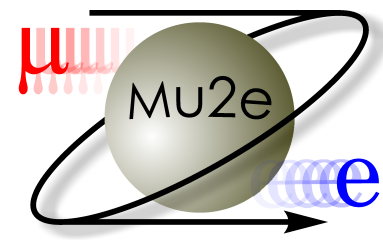
## Beam related bg;

- $\pi/\mu$  decay in flight
- Radiative pion capture
- Antiproton annihilation

## Cosmics;

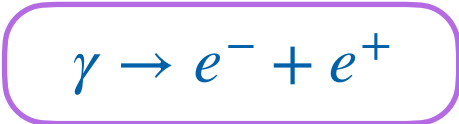
- 1 conversion like  $e^-$  per day.
  - $\mu$  misidentified as  $e^-$
  - Decay in flight
  - Interaction with detector material
- 99.99% veto efficiency is needed !





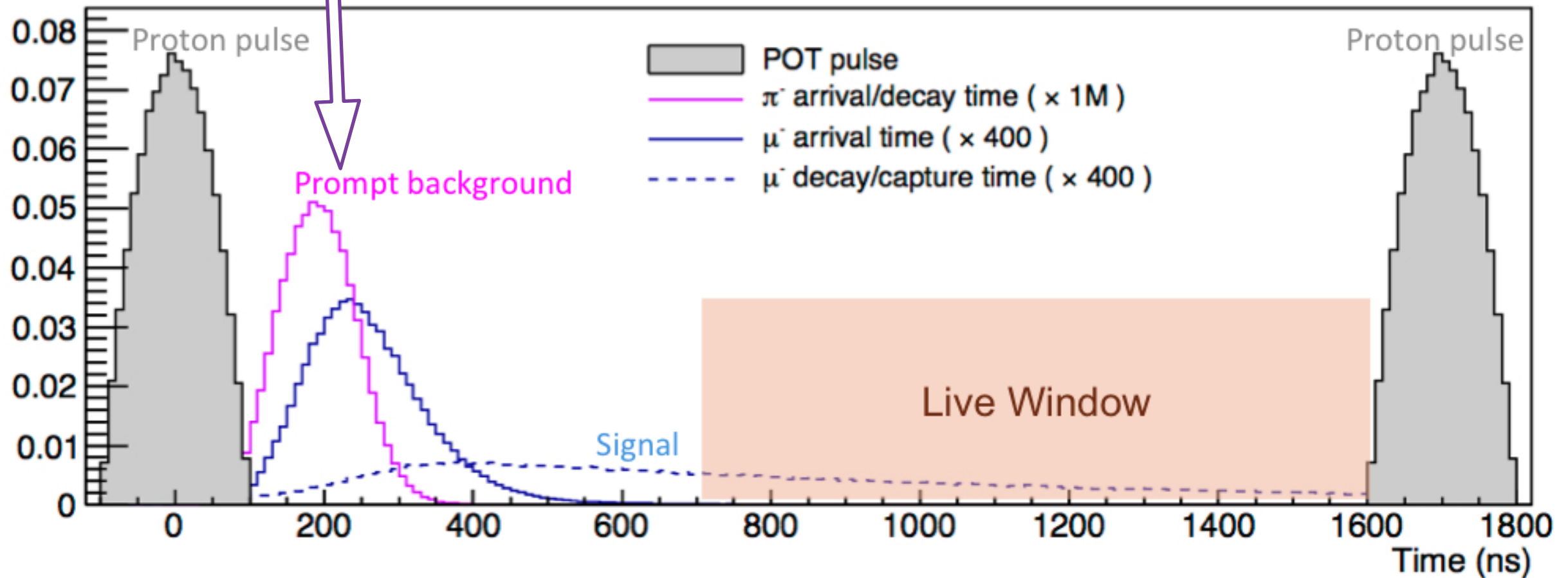
# Live Event Window

Background suppression: Radiative Pion Capture + other beam related bg



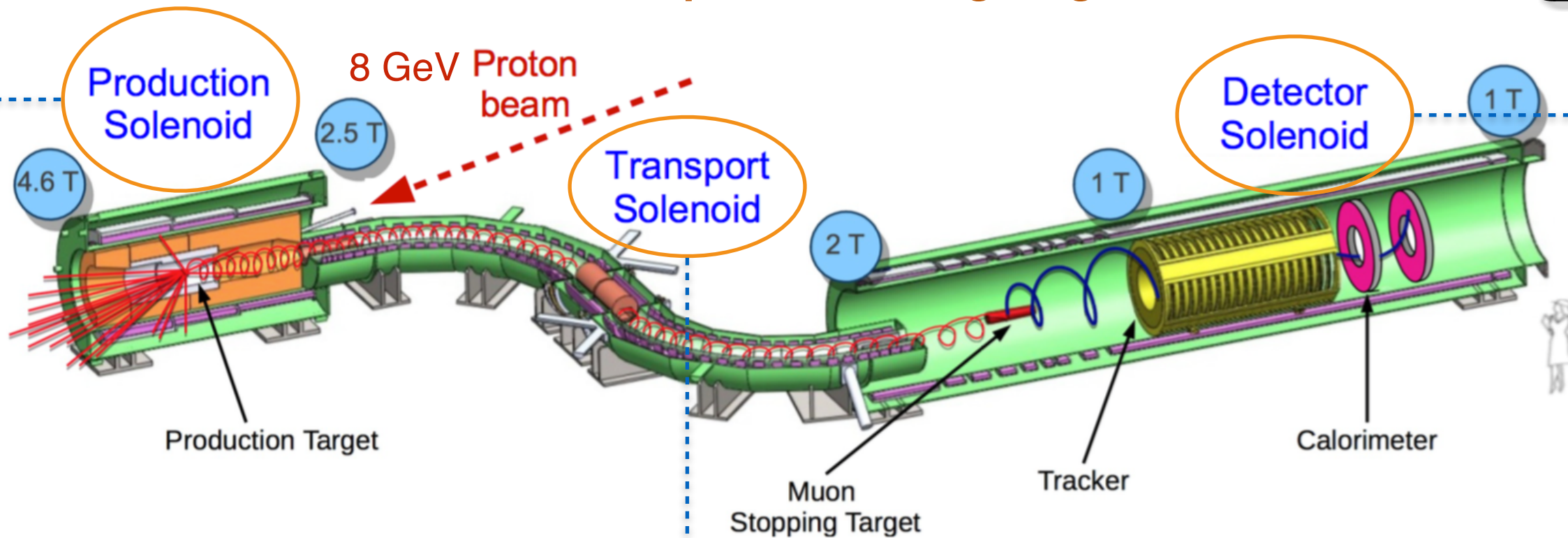
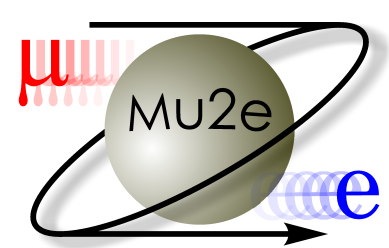
$e^-$  with enough momentum can fake conversion events

- 8 GeV pulsed proton beam @ 1695 ns intervals.
- We wait 700 ns before taking C.E data to avoid most of the **prompt** background
  - Muonic Al lifetime = 864 ns.
- Out of time protons/ beam  $< 10^{-10}$



# Solenoids

## NbTi Superconducting Magnets



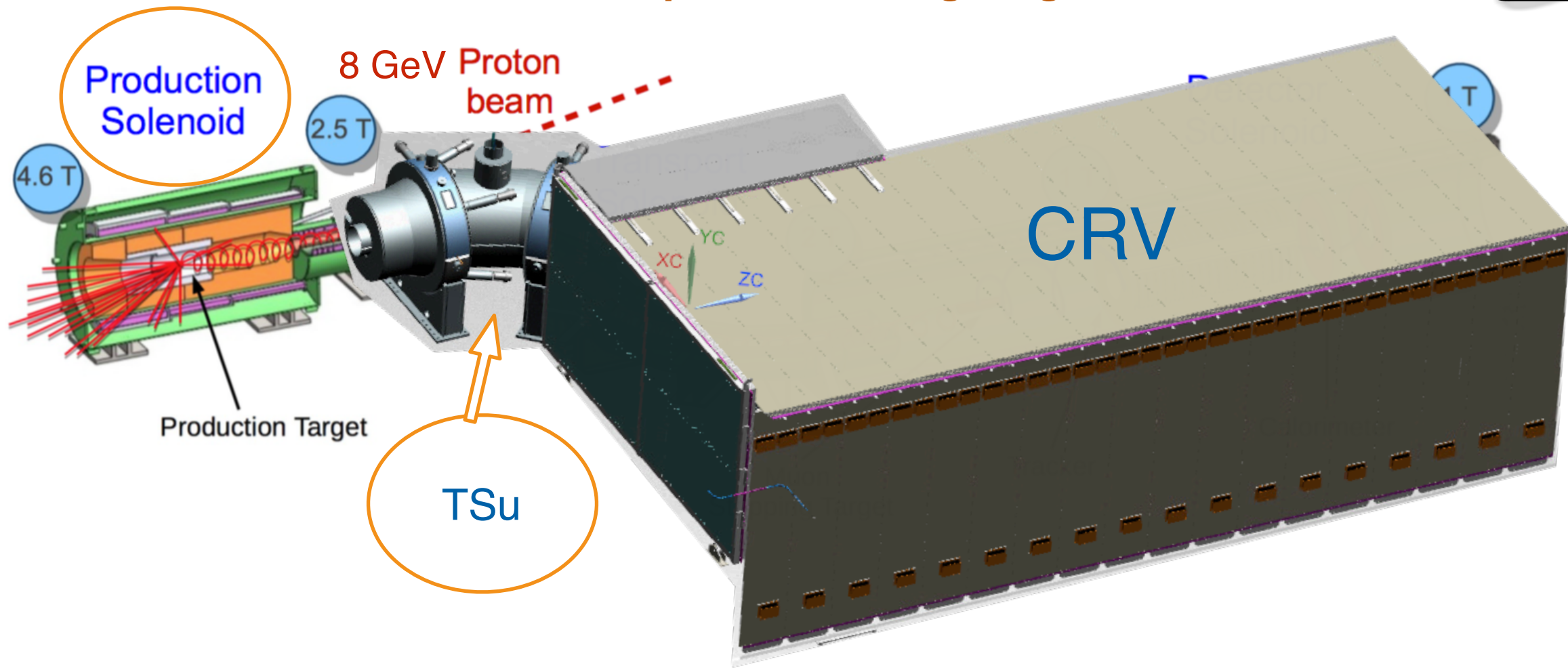
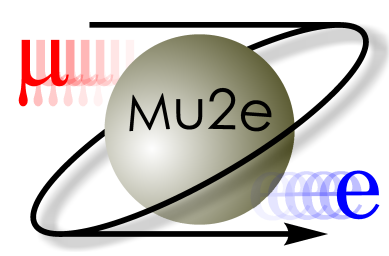
- Direct low momentum pions/muons to transport solenoid.

- S-shaped geometry with collimators select low momentum and negatively charged particles.

- Houses muon stopping target, tracker & calorimeter.

# Solenoids + Cosmic Ray Veto(CRV) + Shielding

## NbTi Superconducting Magnets

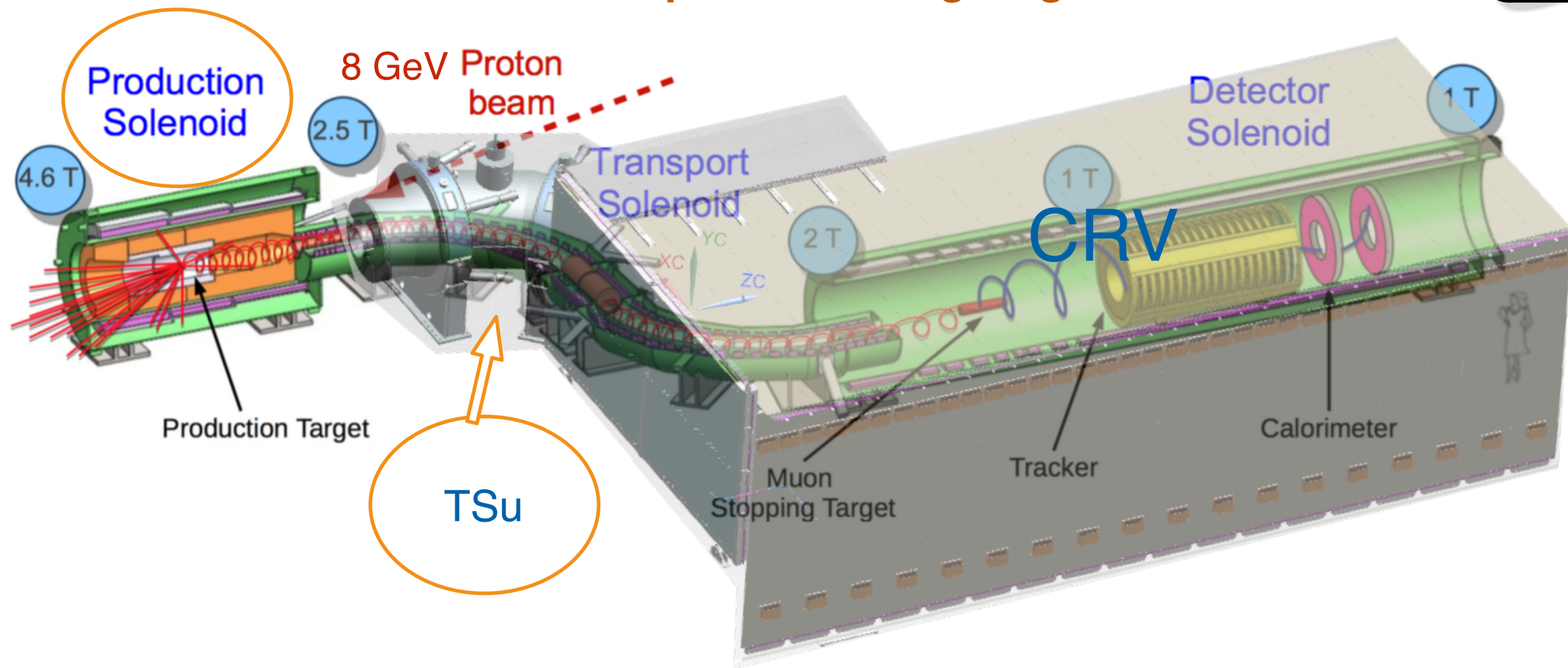
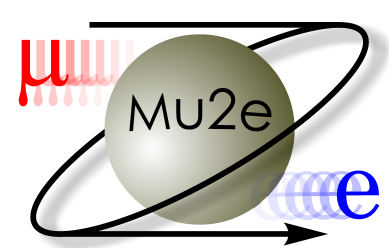


Cosmic Ray Veto covers all of DS and some of TS



# Solenoids + Cosmic Ray Veto(CRV) + Shielding

## NbTi Superconducting Magnets

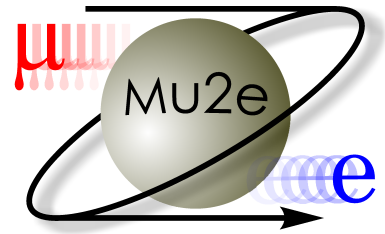


Cosmic Ray Veto covers all of DS and some of TS



# Tracker

Background suppression: DIO

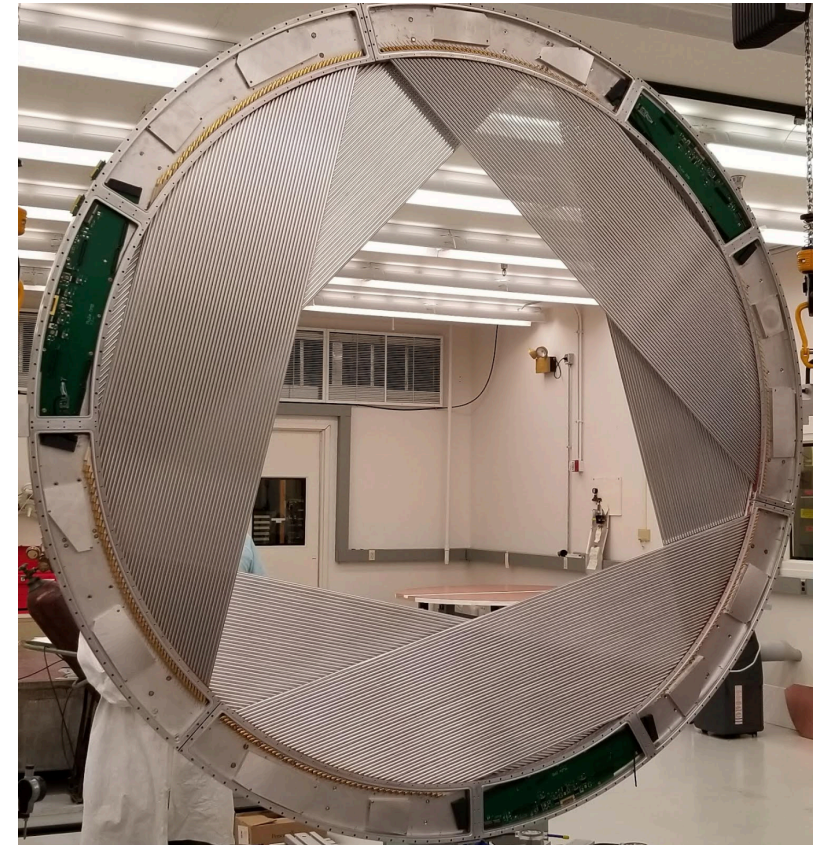


- Main detection element of Mu2e.
- Low mass tracker using straw drift tubes running ArCO<sub>2</sub>(80/20).
- 25  $\mu\text{m}$  Tungsten wire as the anode.
- 21600 x 5 mm OD metalized 15  $\mu\text{m}$  thick walled Mylar straws;
  - Inner coat provides cathode
  - Outer coat provides shielding and reduces leaks.
- Highly segmented -> 36 planes -> each made from 6 panels.
- Momentum resolution < 180 keV/c.

Tracker panel



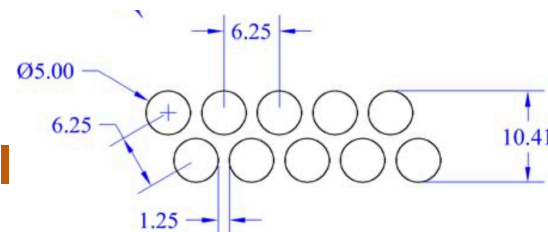
Tracker plane



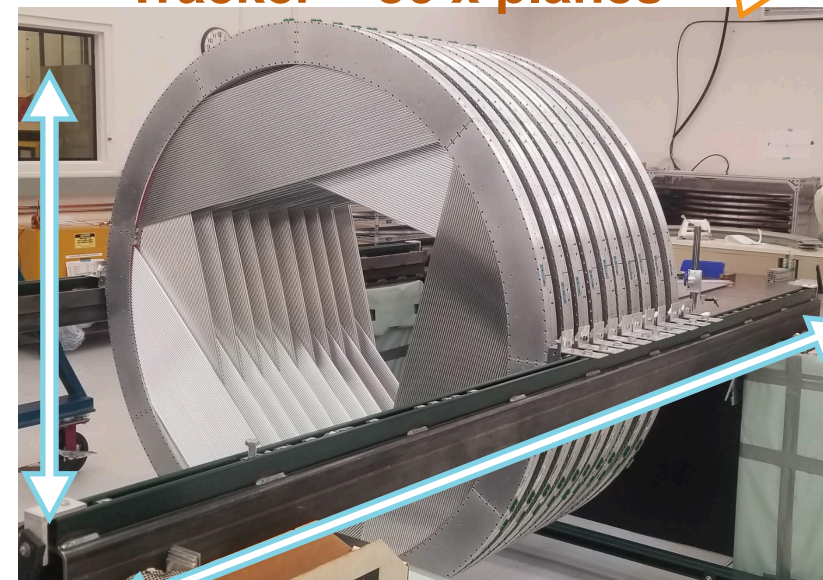
21600 x straws



Double layered  
96 per tracker panel



Tracker = 36 x planes

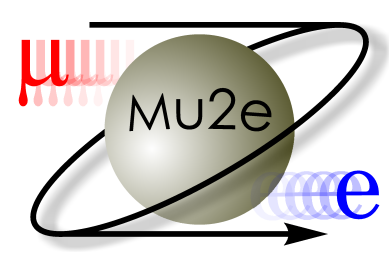


1.6m

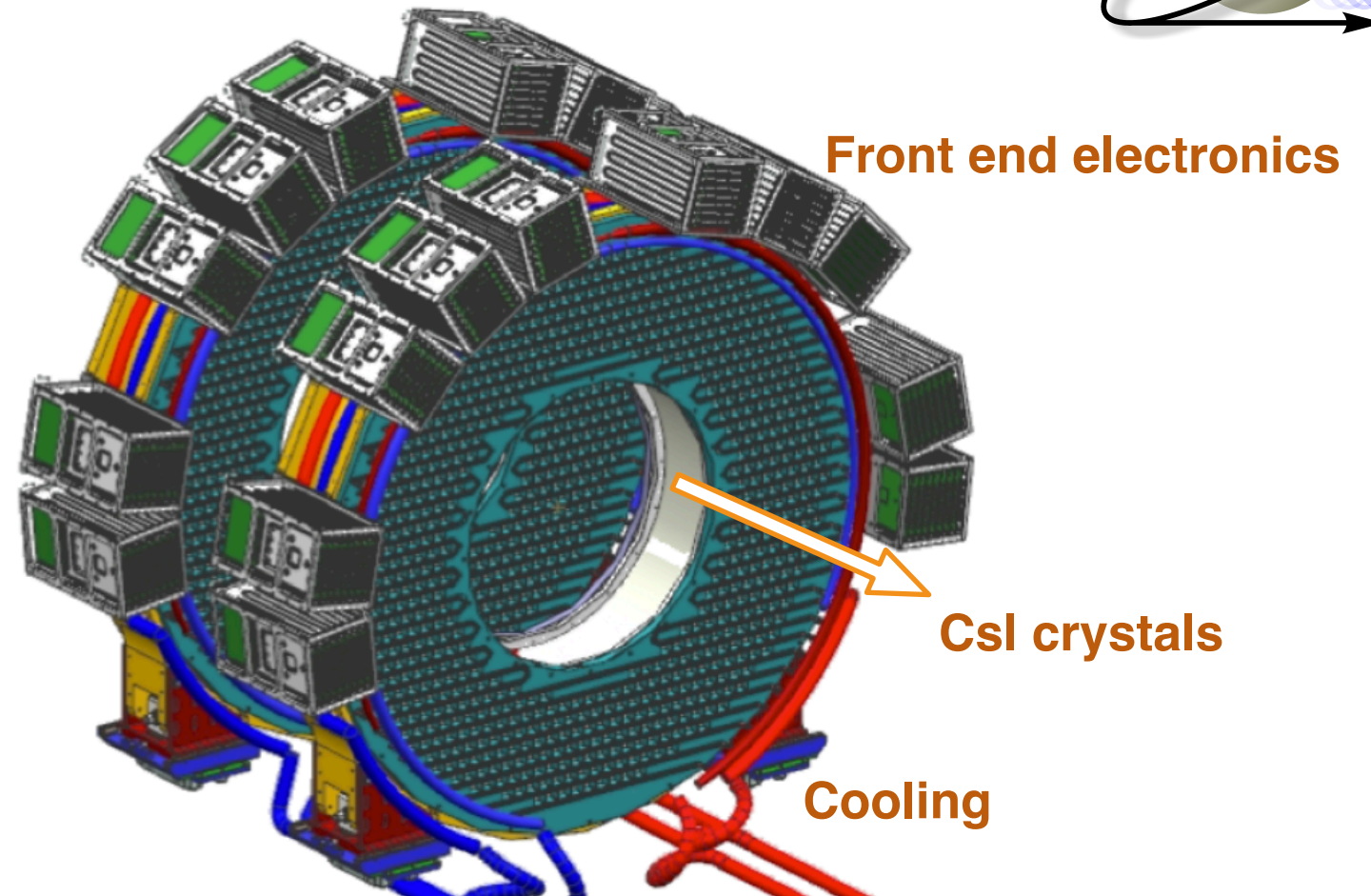
3 m



# EM Calorimeter

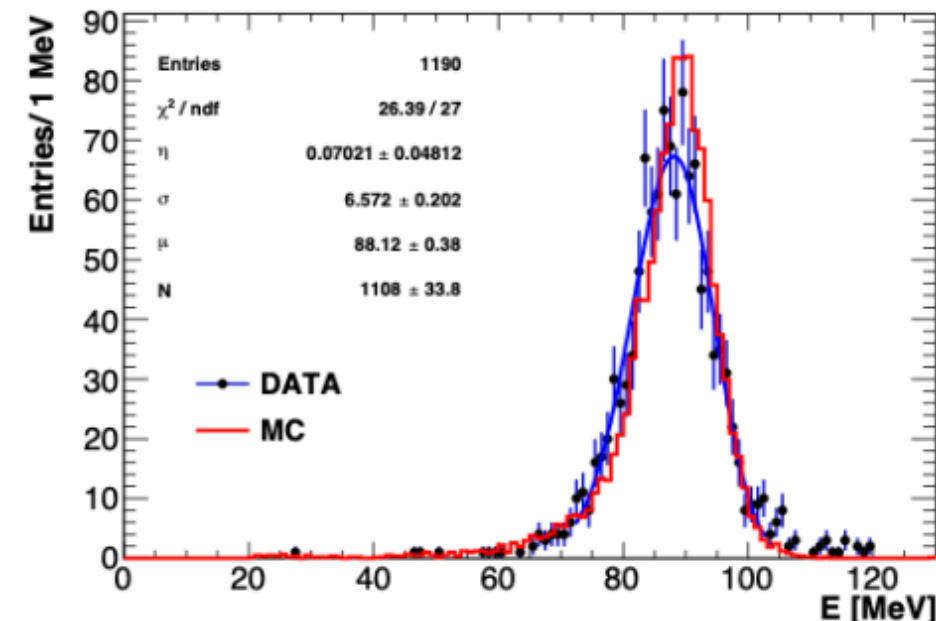
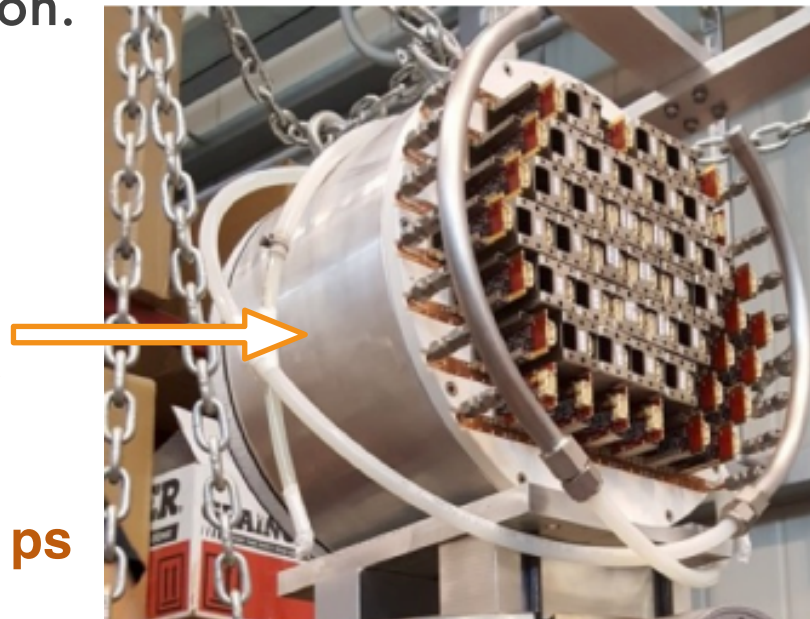


2 disk annular design with hole



- 1348 CsI crystals;
  - 3.4x3.4 cm surface area
  - 20 cm in length.
- Readout by SiPMs.
- Annular design like tracker with hole in the middle.
- Distance between two disks = 70 cm
  - half wavelength of electron's path
- Provides
  - Seed to complement tracking.
  - 0.5 ns time resolution.
  - particle ID, 10% energy resolution.
  - Position, 1 cm spatial resolution.

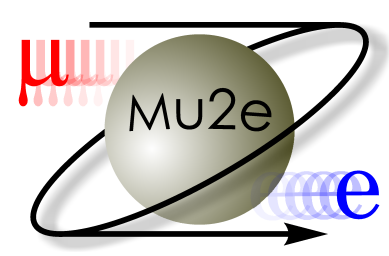
- **Prototype using 51 CsI crystals & 102 SiPMs**
- **5.4% at 100 MeV energy resolution**
- **Timing resolution < 150 ps**



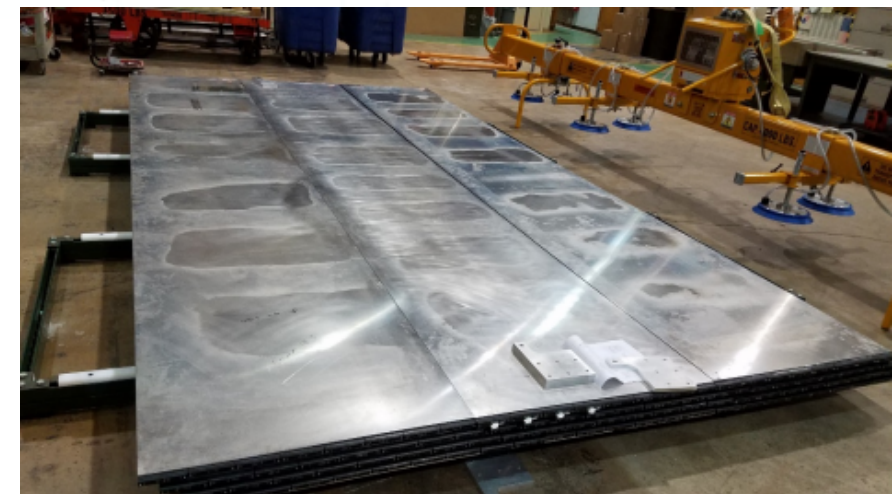
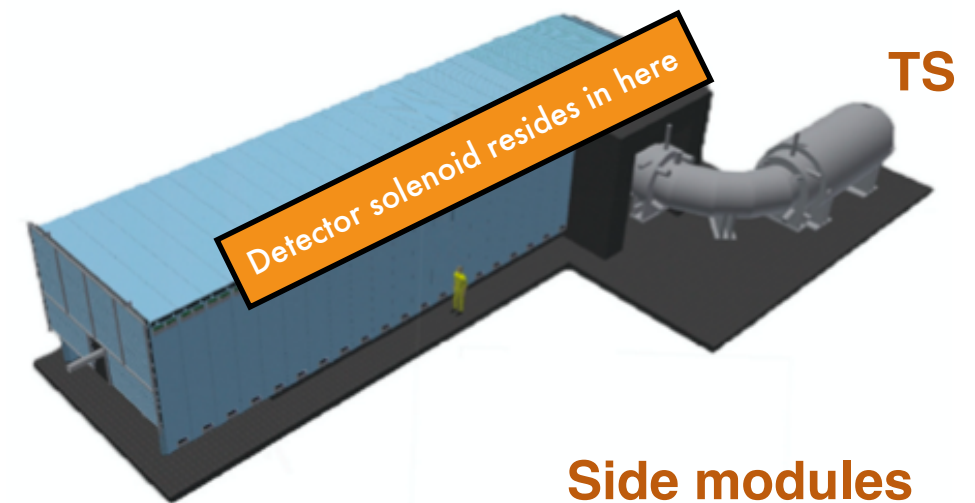


# Cosmic Ray Veto

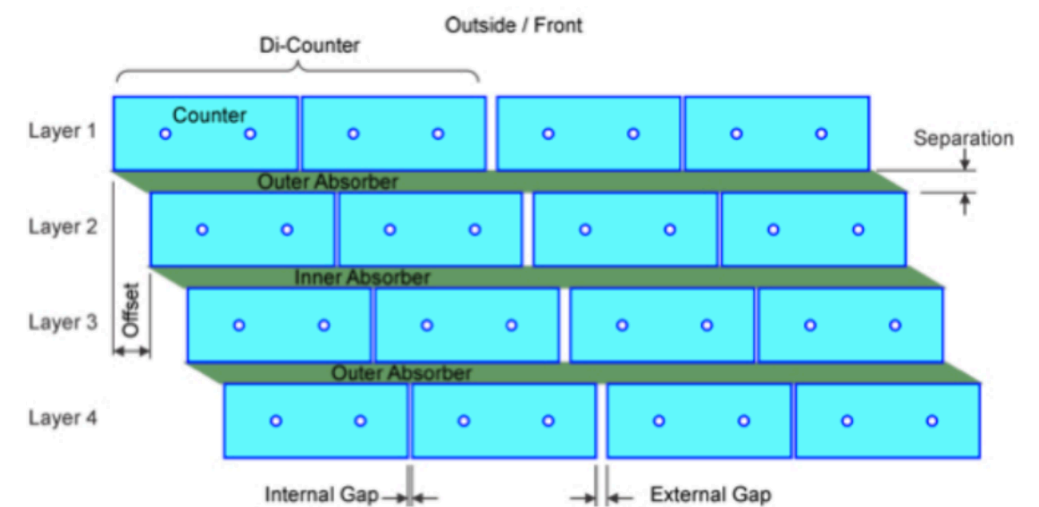
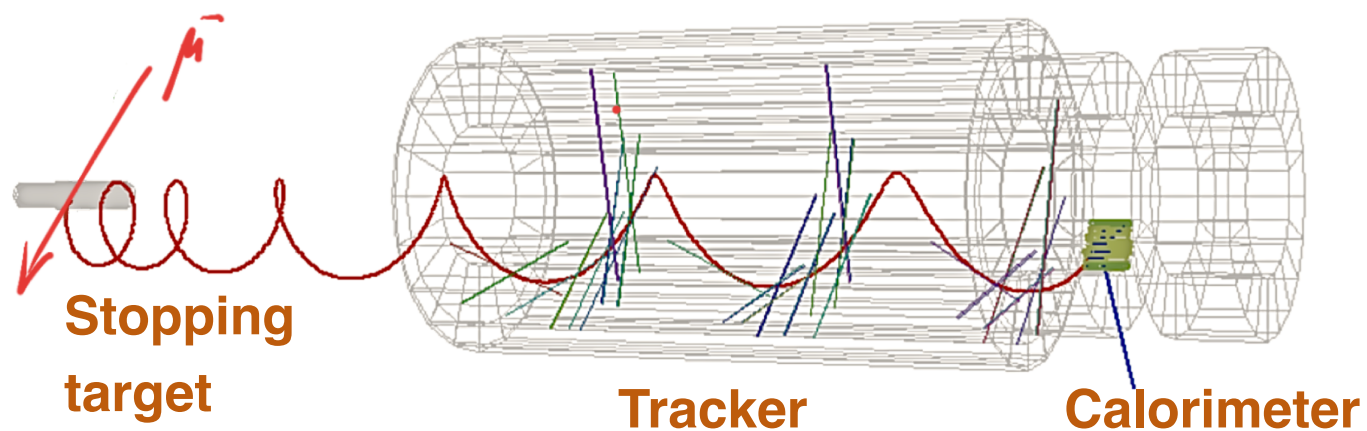
Background suppression: Cosmics



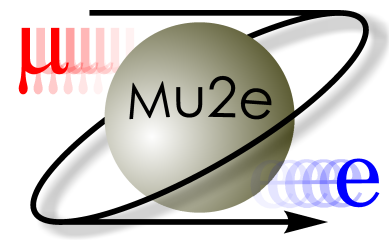
- Unvetoed cosmic background  $\approx$  1 bg event per day.
- Covers all DS and part of TS.
- 337 m<sup>2</sup> surface area.
- Polystyrene scintillators coated with TiO<sub>2</sub> sandwiched between Al absorbers.
- 4 overlapping layers of scintillators.
  - 3 layer coincidence veto
- Readout through WLS fibers & 2x2 mm<sup>2</sup> SiPMs on both ends.



## Cosmic muon

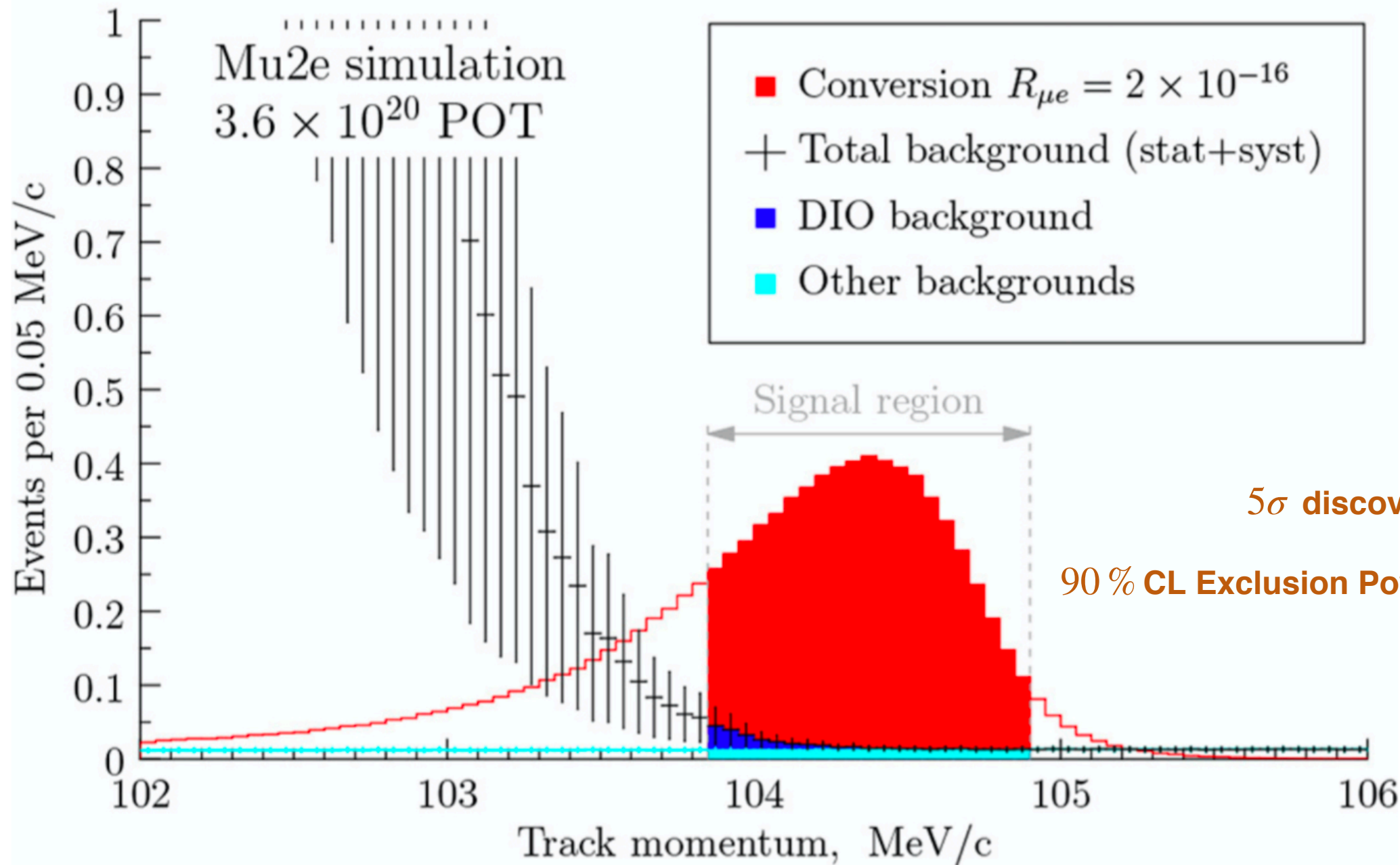
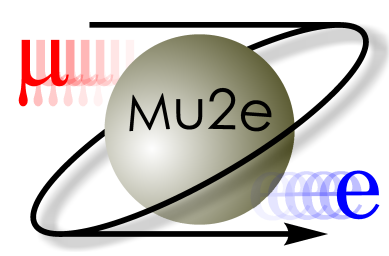


# Background summary



	<i>Process</i>	<i>Estimated yield(events)</i>
Intrinsic	<i>Muon DIO</i>	$0.144 \pm 0.028(\text{stat}) \pm 0.11(\text{syst})$
	<i>RMC</i>	$0.000^{+0.004}_{-0.000}$
Beam related prompt	<i>RPC</i>	$0.021 \pm 0.001(\text{stat}) \pm 0.002(\text{syst})$
	<i>Muon DIF</i>	$< 0.003$
	<i>Pion DIF</i>	$0.001 \pm < 0.001$
	<i>Beam electrons</i>	$(2.1 \pm 1.0) \times 10^{-4}$
Other	<i>Antiproton induced</i>	$0.040 \pm 0.001(\text{stat}) \pm 0.020(\text{syst})$
	<i>Cosmic ray induced</i>	$0.209 \pm 0.022(\text{stat}) \pm 0.055(\text{syst})$
	<b><i>TOTAL</i></b>	<b><math>0.41 \pm 0.13(\text{stat+syst})</math></b>

# Sensitivity



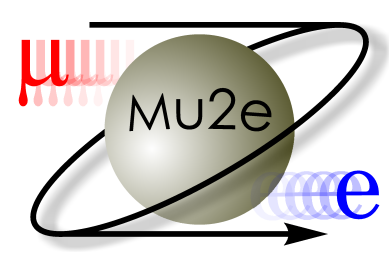
$5\sigma$  discovery :  $R_{\mu e} \geq 2 \times 10^{-16}$

90% CL Exclusion Power :  $R_{\mu e} \geq 8 \times 10^{-17}$

7 events are needed for  $5\sigma$



# Mu2e-II with PIP-II beam line

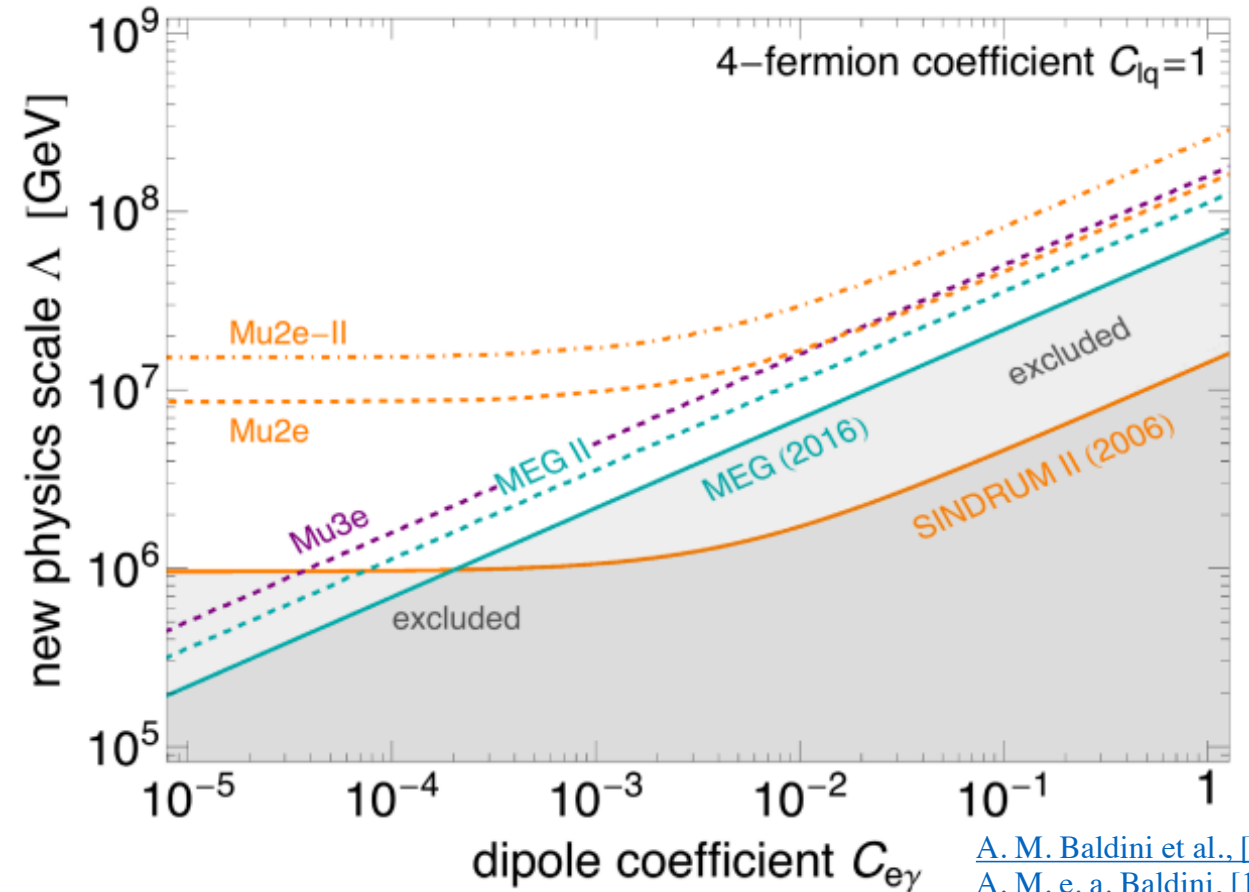


Increasing Mu2e capability

- Improve sensitivity.
- Probe higher mass scale.

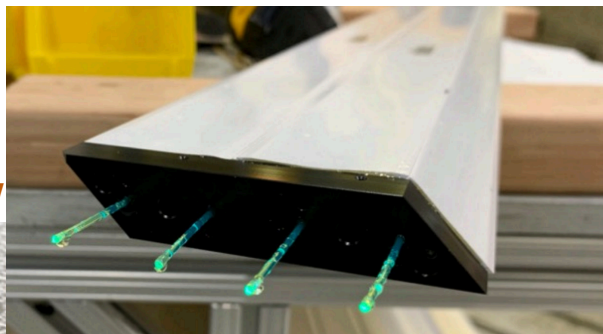
Expanding upon Mu2e goals

- Change targets.
- Focus on excluding/including models.
- $\mu^- + N \rightarrow e^+ + N'$
- $\mu \rightarrow eX$



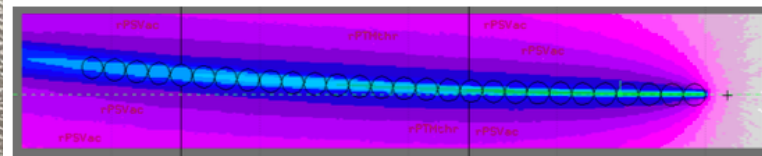
[A. M. Baldini et al., \[1605.05081\]](#)  
[A. M. e. a. Baldini, \[1801.04688\]](#)  
[A. Blondel et al., \[1301.6113\]](#)

Parameter	Mu2e	Mu2e-II
Proton source	Slow extraction from DR	PIP-II Linac
Proton kinetic energy	8 GeV	0.8 GeV
Beam Power for expt.	8 kW	100 kW
Protons/s	$6.25 \times 10^{12}$	$7.8 \times 10^{14}$
Pulse Cycle Length	1.693 $\mu$ s	1.693 $\mu$ s
Proton rms emittance	2.7	0.25
Proton geometric emittance	0.29	0.16
Proton Energy Spread ( $\sigma_E$ )	20 MeV	0.275 MeV
$\delta p/p$	$2.25 \times 10^{-3}$	$2.2 \times 10^{-4}$
Stopped $\mu$ per proton	$1.59 \times 10^{-3}$	$9.1 \times 10^{-5}$
Stopped $\mu$ per cycle		$1.2 \times 10^5$



Triangular  
CRV  
scintillators

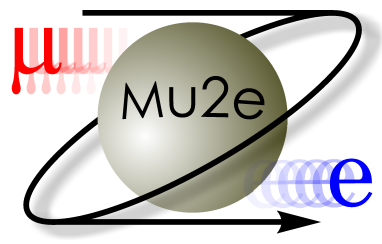
8  $\mu$ m Tracker straw



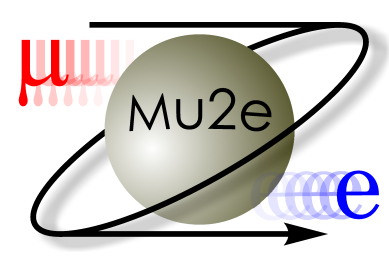
Carbide production target

[hep-ex > arXiv: 2203.07569](https://arxiv.org/abs/2203.07569)

# Summary

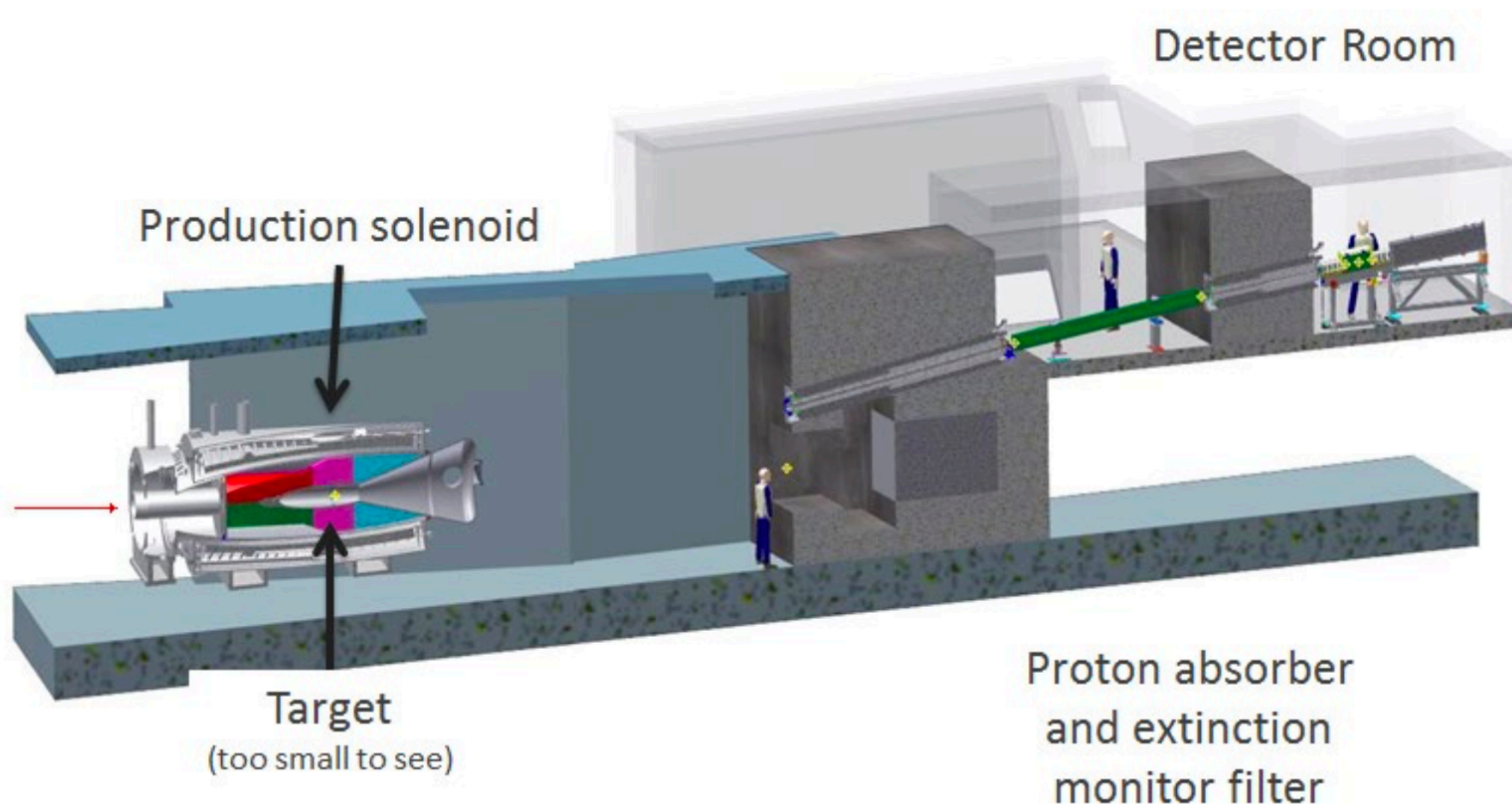
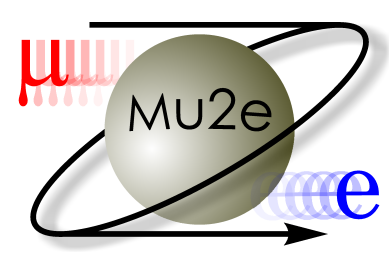


- Mu2e will improve current limit on conversion rate by  $10^4$  @ SES =  $3 \times 10^{-17}$ .
- Will probe mass scales up to  $10^4$  TeV.
- Current schedule;
  - Installation and commissioning starting in 2023.
  - Beam commissioning and physics data taking 2026.
  - $\times 1000$  improvement over current limit by 2027.
  - LBNF/PIP-II shutdown.
  - $\times 10000$  improvement over current limit by the end of the decade.
- Next 2 years will see a big effort on building and commissioning the detector.



# BACKUP

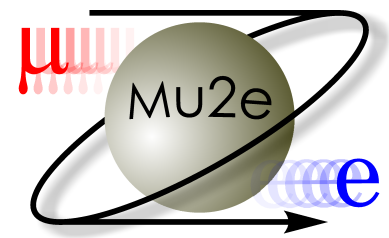
# Extinction monitor



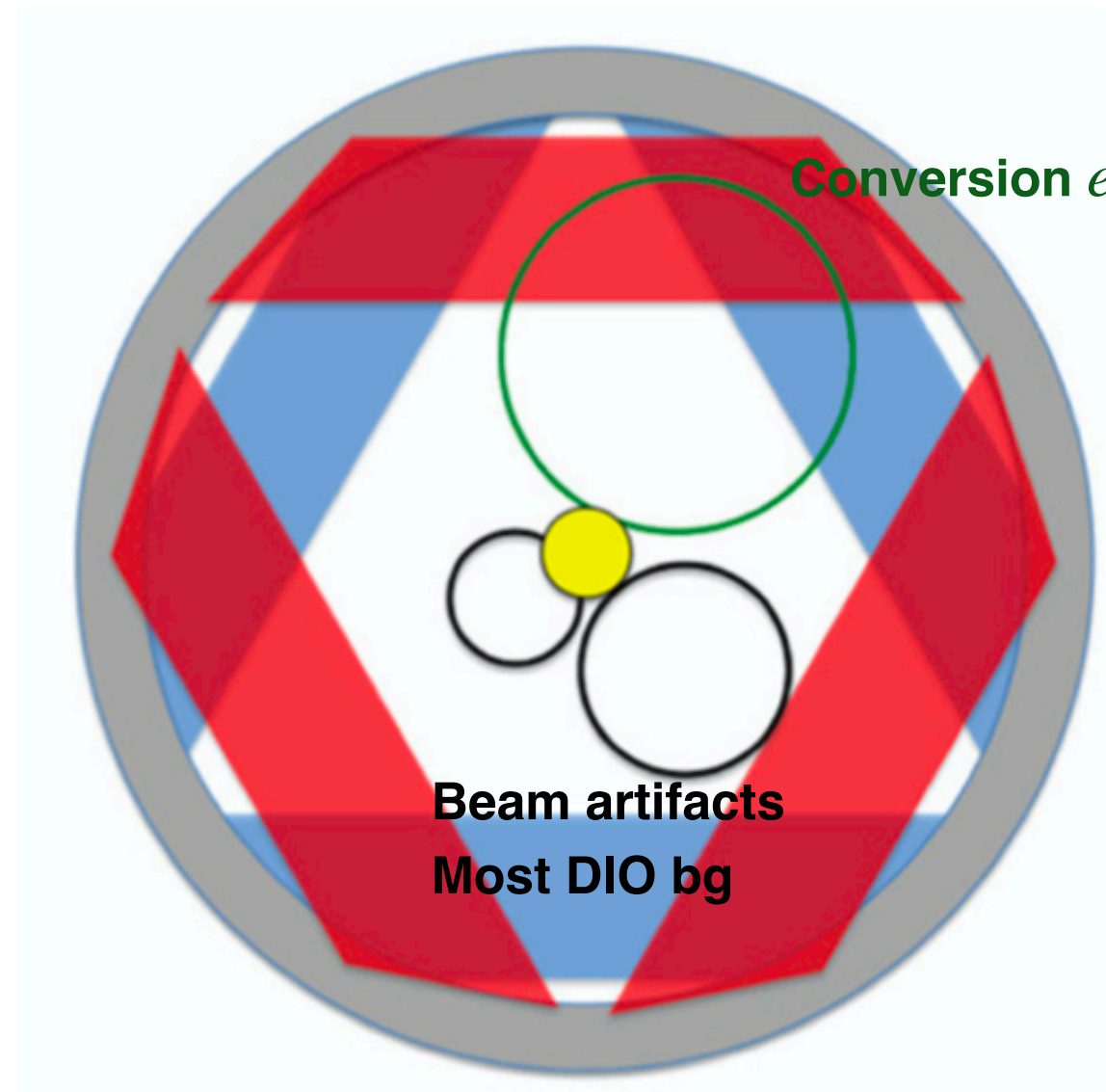
Mu2e Extinction monitor

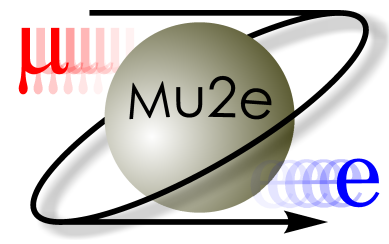


# Tracker hole in the middle design



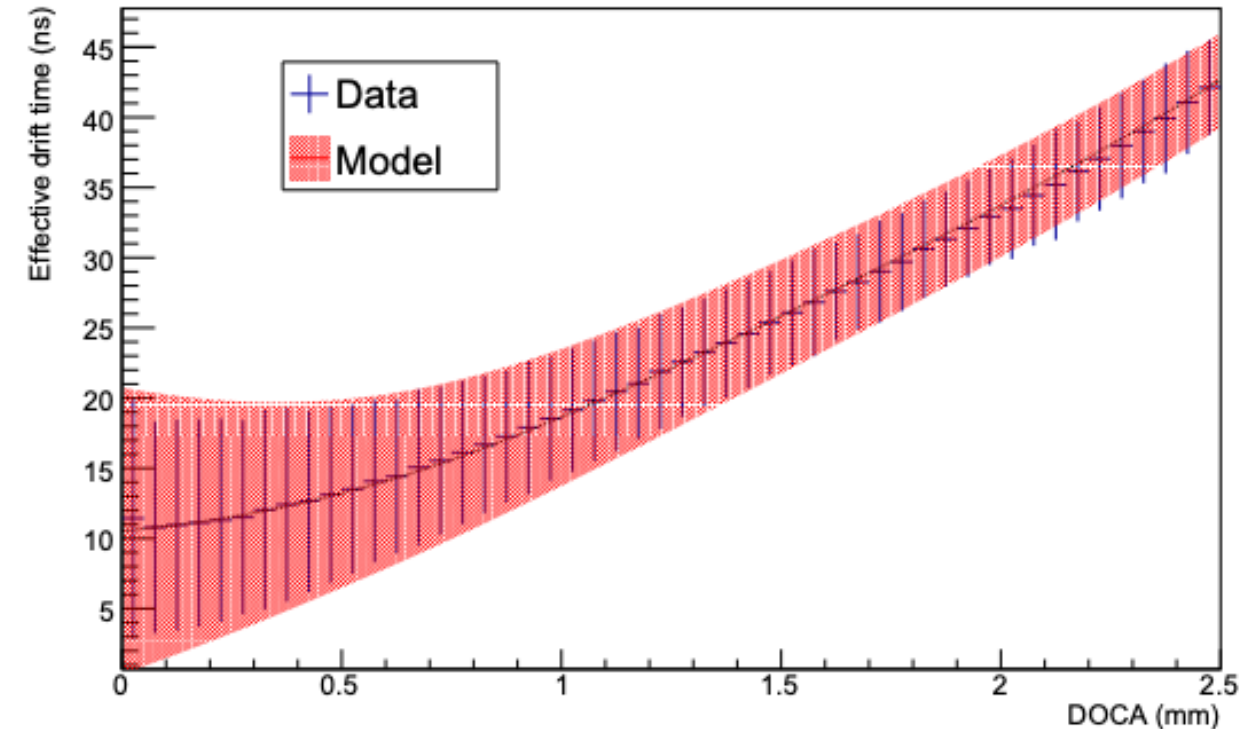
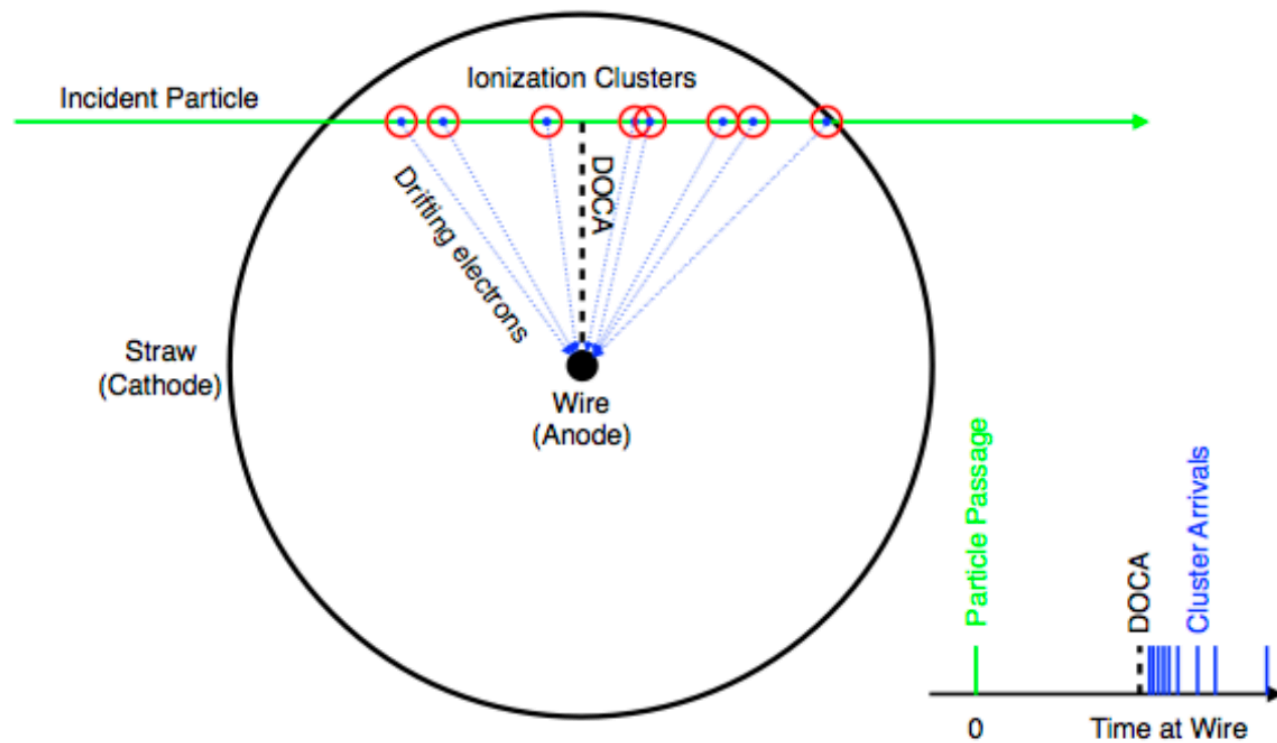
- Center (<400 mm) is empty to make the detector blind against most DIO electrons, beam artifacts.

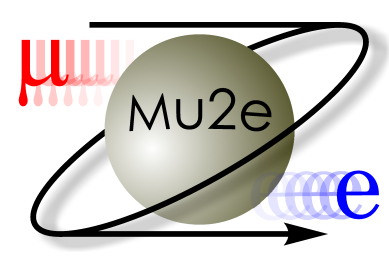




# Cosmic run with tracker panel

- Data was taken May 2020 with production tracker panel.
- DOCA (distance of closest approach) is determined to compute drift time.





# Targets

- Production target
  - Tungsten
  - Suspended on spokes
  - Minimize scattering &  $\pi$  absorption
  - 1400 msec beam cycles
  - 630 W power absorption
  - 2000 K temperature
  - Operate 1 year

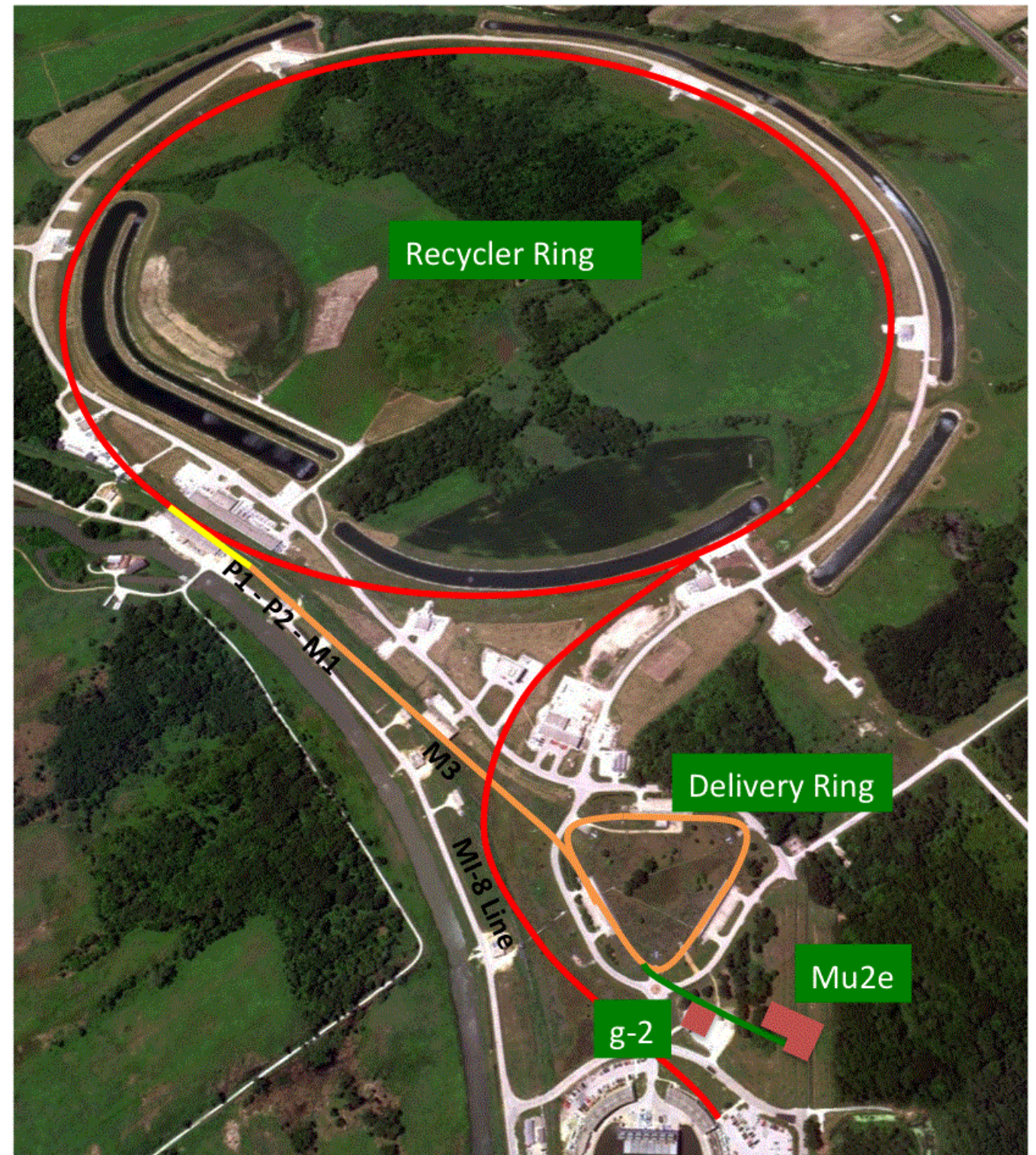
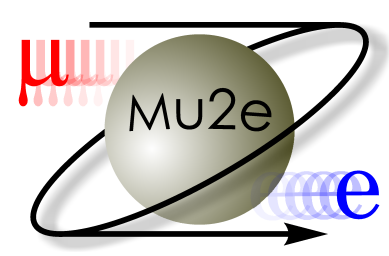


- Stopping target
  - 37 high purity Al disks
  - Each 100  $\mu m$  thick, 150 mm OD, 40 mm ID.
  - 740 mm in length.
  - Suspended with 76  $\mu m$  diameter gold plated W wires.



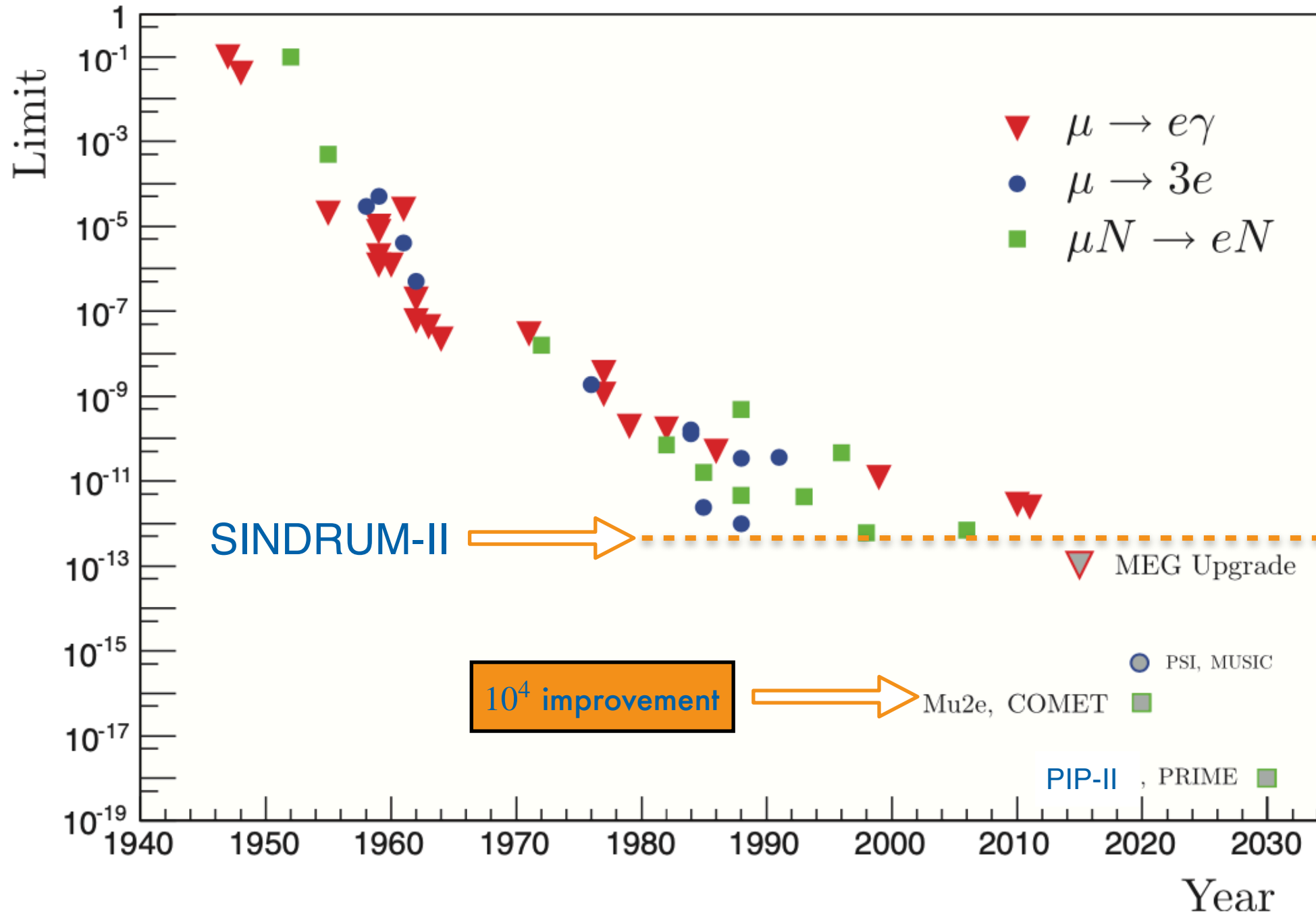
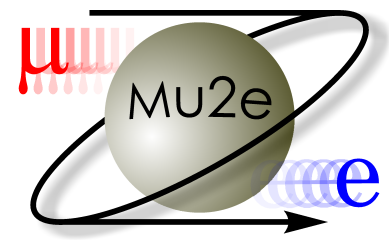
# Beam delivery

- 8 GeV protons are transferred to DR(delivery ring) from recycler.
- 2.5 MHz bunches.
- Protons are extracted from DR and sent to *Mu2e* in 1695 ns intervals.
- $3.9 \times 10^7$  POT per bunch.
- $3.6 \times 10^{20}$  POT total



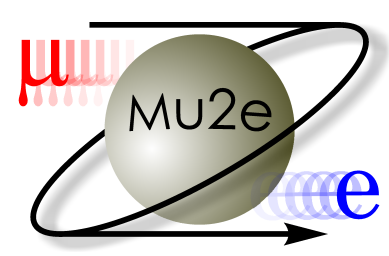


# Muon searches history



R. Bernstein, P. Cooper, arXiv:1307.5787

# CLFV processes sensitivity to BSM



W. Altmannshofer, A.J.Buras, S.Gori, P.Paradisi, D.M.Straub

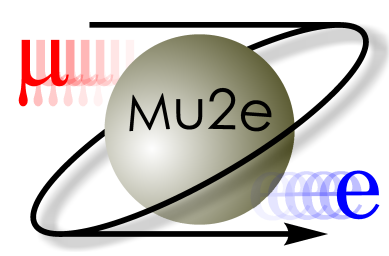
★★★★ = Discovery Sensitivity

	AC	RVV2	AKM	$\delta LL$	FBMSSM	LHT	RS
$D^0 - \bar{D}^0$	★★★★	★	★	★	★	★★★★	?
$\epsilon_K$	★	★★★★	★★★★	★	★	★★	★★★★
$S_{\psi\phi}$	★★★★	★★★★	★★★★	★	★	★★★★	★★★★
$S_{\phi K_S}$	★★★★	★★	★	★★★★	★★★★	★	?
$A_{CP}(B \rightarrow X_s \gamma)$	★	★	★	★★★★	★★★★	★	?
$A_{7,8}(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★★★★	★★★★	★★	?
$A_9(B \rightarrow K^* \mu^+ \mu^-)$	★	★	★	★	★	★	?
$B \rightarrow K^{(*)} \nu \bar{\nu}$	★	★	★	★	★	★	★
$B_s \rightarrow \mu^+ \mu^-$	★★★★	★★★★	★★★★	★★★★	★★★★	★	★
$K^+ \rightarrow \pi^+ \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★
$K_L \rightarrow \pi^0 \nu \bar{\nu}$	★	★	★	★	★	★★★★	★★★★
$\mu \rightarrow e \gamma$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
$\tau \rightarrow \mu \gamma$	★★★★	★★★★	★	★★★★	★★★★	★★★★	★★★★
$\mu + N \rightarrow e + N$	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★	★★★★
$d_n$	★★★★	★★★★	★★★★	★★	★★★★	★	★★★★
$d_e$	★★★★	★★★★	★★	★	★★★★	★	★★★★
$(g-2)_\mu$	★★★★	★★★★	★★	★★★★	★★★★	★	?

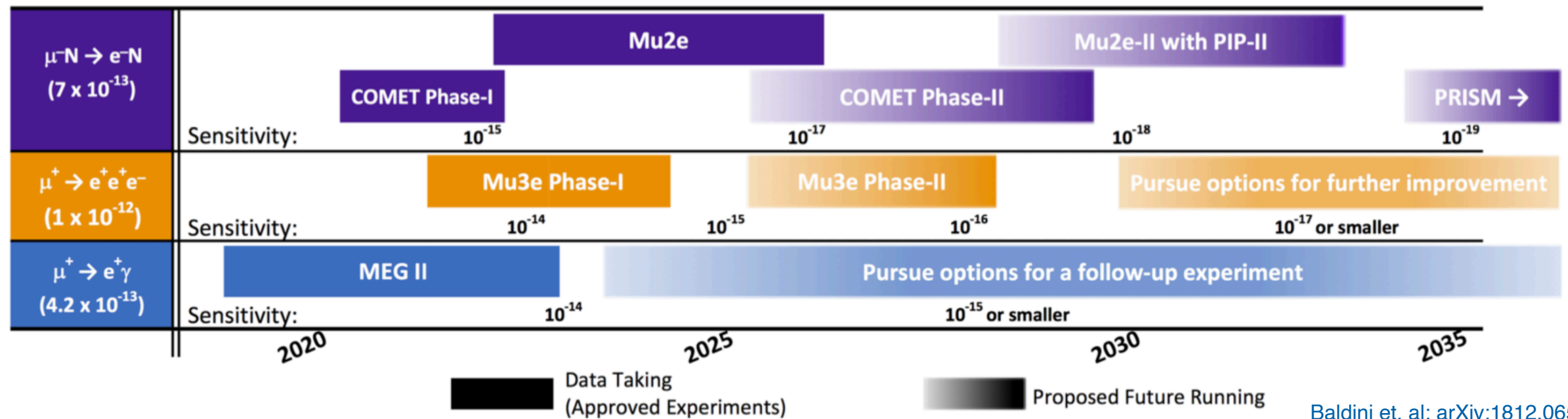
arXiv:0909.1333[hep-ph]

Table 8: “DNA” of flavour physics effects for the most interesting observables in a selection of SUSY and non-SUSY models ★★★★★ signals large effects, ★★ visible but small effects and ★ implies that the given model does not predict sizable effects in that observable.

# Looking forward in muon searches

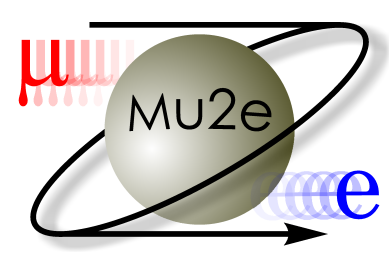


Searches for Charged-Lepton Flavor Violation in Experiments using Intense Muon Beams



Baldini et. al; [arXiv:1812.06540v1](https://arxiv.org/abs/1812.06540v1)

# CLFV experimental limits

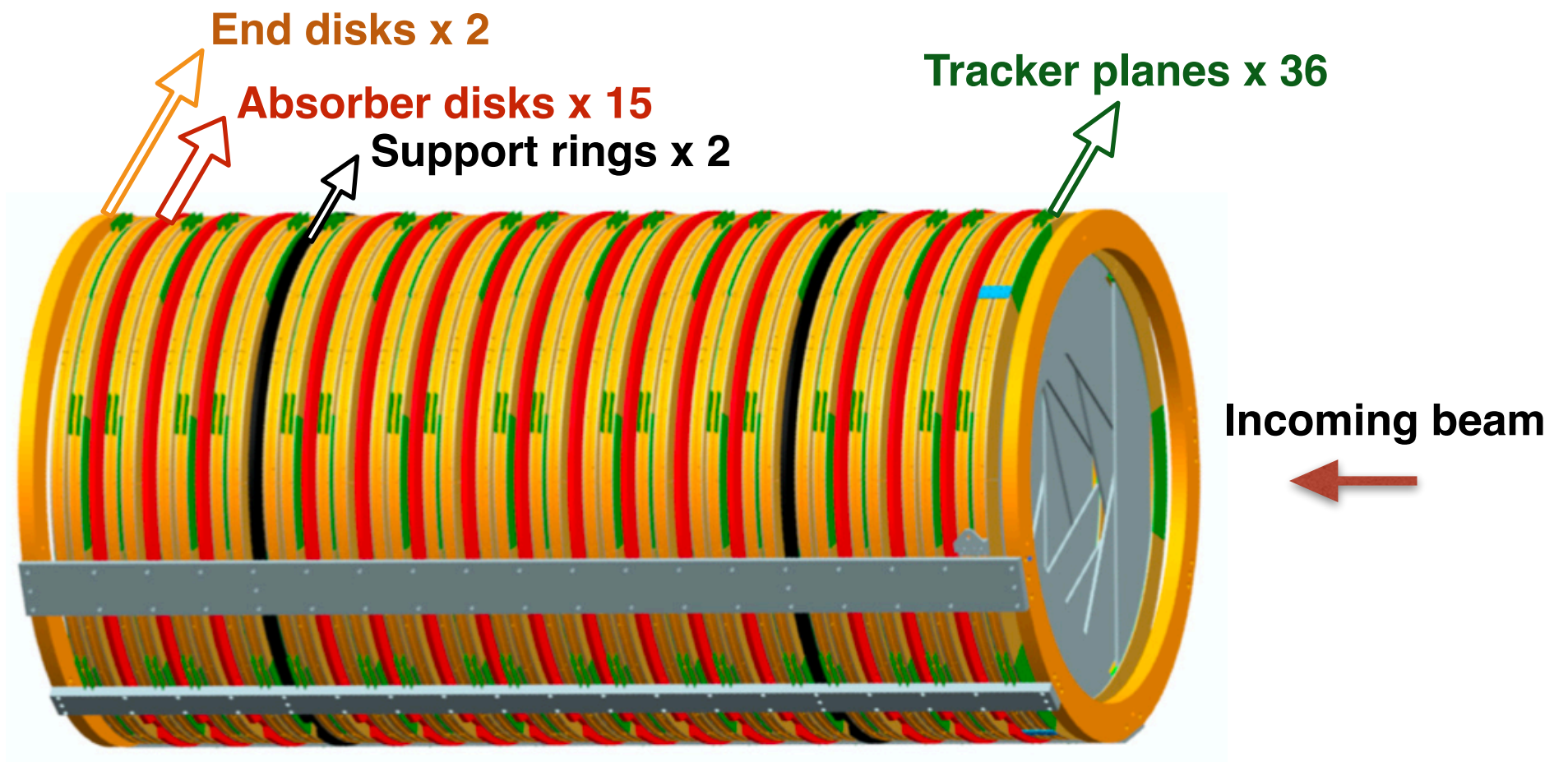
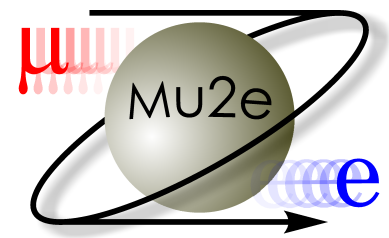


Reaction	Present limit	C.L.	Experiment	Year
$\mu^+ \rightarrow e^+ \gamma$	$< 4.2 \times 10^{-13}$	90%	MEG at PSI	2016
$\mu^+ \rightarrow e^+ e^- e^+$	$< 1.0 \times 10^{-12}$	90%	SINDRUM	1988
$\mu^- \text{Ti} \rightarrow e^- \text{Ti}^\dagger$	$< 6.1 \times 10^{-13}$	90%	SINDRUM II	1998
$\mu^- \text{Pb} \rightarrow e^- \text{Pb}^\dagger$	$< 4.6 \times 10^{-11}$	90%	SINDRUM II	1996
$\mu^- \text{Au} \rightarrow e^- \text{Au}^\dagger$	$< 7.0 \times 10^{-13}$	90%	SINDRUM II	2006
$\mu^- \text{Ti} \rightarrow e^+ \text{Ca}^* \dagger$	$< 3.6 \times 10^{-11}$	90%	SINDRUM II	1998
$\mu^+ e^- \rightarrow \mu^- e^+$	$< 8.3 \times 10^{-11}$	90%	SINDRUM	1999
$\tau \rightarrow e \gamma$	$< 3.3 \times 10^{-8}$	90%	BaBar	2010
$\tau \rightarrow \mu \gamma$	$< 4.4 \times 10^{-8}$	90%	BaBar	2010
$\tau \rightarrow e e e$	$< 2.7 \times 10^{-8}$	90%	Belle	2010
$\tau \rightarrow \mu \mu \mu$	$< 2.1 \times 10^{-8}$	90%	Belle	2010
$\tau \rightarrow \pi^0 e$	$< 8.0 \times 10^{-8}$	90%	Belle	2007
$\tau \rightarrow \pi^0 \mu$	$< 1.1 \times 10^{-7}$	90%	BaBar	2007
$\tau \rightarrow \rho^0 e$	$< 1.8 \times 10^{-8}$	90%	Belle	2011
$\tau \rightarrow \rho^0 \mu$	$< 1.2 \times 10^{-8}$	90%	Belle	2011
$\pi^0 \rightarrow \mu e$	$< 3.6 \times 10^{-10}$	90%	KTeV	2008
$K_L^0 \rightarrow \mu e$	$< 4.7 \times 10^{-12}$	90%	BNL E871	1998
$K_L^0 \rightarrow \pi^0 \mu^+ e^-$	$< 7.6 \times 10^{-11}$	90%	KTeV	2008
$K^+ \rightarrow \pi^+ \mu^+ e^-$	$< 1.3 \times 10^{-11}$	90%	BNL E865	2005
$J/\psi \rightarrow \mu e$	$< 1.5 \times 10^{-7}$	90%	BESIII	2013
$J/\psi \rightarrow \tau e$	$< 8.3 \times 10^{-6}$	90%	BESII	2004
$J/\psi \rightarrow \tau \mu$	$< 2.0 \times 10^{-6}$	90%	BESII	2004
$B^0 \rightarrow \mu e$	$< 2.8 \times 10^{-9}$	90%	LHCb	2013
$B^0 \rightarrow \tau e$	$< 2.8 \times 10^{-5}$	90%	BaBar	2008
$B^0 \rightarrow \tau \mu$	$< 2.2 \times 10^{-5}$	90%	BaBar	2008
$B \rightarrow K \mu e^\dagger$	$< 3.8 \times 10^{-8}$	90%	BaBar	2006
$B \rightarrow K^* \mu e^\dagger$	$< 5.1 \times 10^{-7}$	90%	BaBar	2006
$B^+ \rightarrow K^+ \tau \mu$	$< 4.8 \times 10^{-5}$	90%	BaBar	2012
$B^+ \rightarrow K^+ \tau e$	$< 3.0 \times 10^{-5}$	90%	BaBar	2012
$B_s^0 \rightarrow \mu e$	$< 1.1 \times 10^{-8}$	90%	LHCb	2013
$\Upsilon(1s) \rightarrow \tau \mu$	$< 6.0 \times 10^{-6}$	95%	CLEO	2008
$Z \rightarrow \mu e$	$< 7.5 \times 10^{-7}$	95%	LHC ATLAS	2014
$Z \rightarrow \tau e$	$< 9.8 \times 10^{-6}$	95%	LEP OPAL	1995
$Z \rightarrow \tau \mu$	$< 1.2 \times 10^{-5}$	95%	LEP DELPHI	1997
$h \rightarrow e \mu$	$< 3.5 \times 10^{-4}$	95%	LHC CMS	2016
$h \rightarrow \tau \mu$	$< 2.5 \times 10^{-3}$	95%	LHC CMS	2017
$h \rightarrow \tau e$	$< 6.1 \times 10^{-3}$	95%	LHC CMS	2017

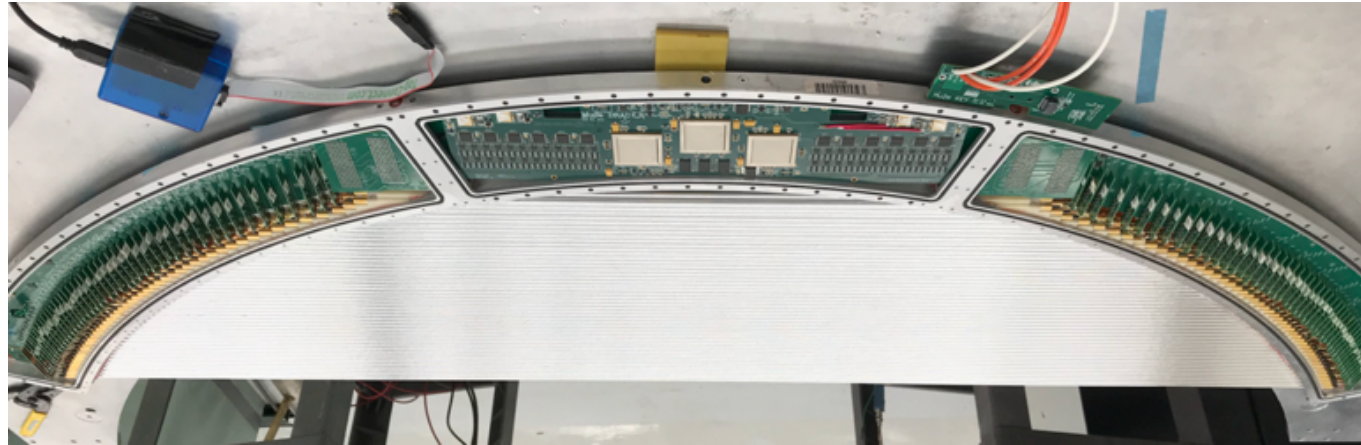
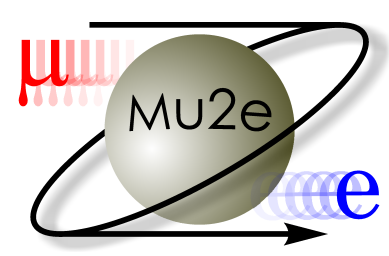
L. Calibbi and G. Signorelli; [arXiv:1709.00294v2](https://arxiv.org/abs/1709.00294v2)



# Tracker Frame



# Cosmic track reconstruction



Panel orientation

