MicroBooNE results and BSM program and a review of short-baseline anomalies

Giuseppe Cerati (Fermilab) on behalf of the MicroBooNE Collaboration **FPCP Conference** May 25, 2022



LUCIO FONTANA, Concetto spaziale, Attese [Spatial Concept, Expectations] Waterpaint on canvas, 1968





Neutrinos in the Standard Model



Image courtesy of Symmetry magazine, a joint Fermilab/SLAC publication. Artwork by Sandbox Studio, Chicago

- Neutrinos fit nicely in the SM
 - massless, neutral leptons
 - 3 flavors, mirroring the charged leptons





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- Neutrinos fit nicely in the SM
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 - 3 flavors, mirroring the charged leptons

 Except that we know they have mass! - are there also >3 flavors?







3-flavor neutrino oscillations

normal hierarchy (NH)

- interaction state
- $m^2 \bigstar$ ν_3 ~2·10⁻³ eV² ν_2 $\sim 7 \cdot 10^{-5} \, eV^2$ ν_1 ν_e $\nu_{\mu} \ \nu_{\tau}$



 We know neutrinos have mass because of neutrino oscillations: they have a mass state that differs from the

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \mathsf{PMNS} \\ \mathsf{matrix} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

 Two main oscillations patterns between 2 mass states: solar and atmospheric

$$P(\nu_{\alpha} \to \nu_{\beta}) = \sin^2(2\theta)\sin^2(1.27\Delta m^2 \frac{L}{E})$$

- mixing angle $sin^2(2\theta)$ determines amplitude (size of effect) - L/E determines oscillation frequency, allowing to extract Δm^2

 PMNS matrix formalism works extremely well - except for short-baseline anomalies!

More about neutrino oscillations in R. Patterson's talk later in this session







Short-baseline Neutrino Anomalies



2022/05/25 5 For more complete reviews see e.g. Snowmass whitepaper (arXiv:2203.07323) and Giunti/Lasserre (arXiv:1901.08330)

credit: G. Karagiorgi

Radioactive anomalies - E~1 MeV

- SAGE/GALLEX experiments used ⁵¹Cr and ³⁷Ar radioactive sources (producing v_e) for calibration of their Gallium detectors
 - results show deficit compared to prediction at $\sim 3\sigma$
 - cross-section uncertainties are being re-evaluated, likely to reduce but not remove significance

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 - cross-section uncertainties are being re-evaluated, likely to reduce but not remove significance
- BEST experiment recently confirmed deficit
 - Gallium detector with a two-target (inner/outer) setup
 - larger (>4 σ) deficit compared to prediction
 - no significant rate difference between targets
 - fitted as $\Delta m^2 = 3.3 \text{ eV}^2$ and $\sin^2(2\theta) = 0.42$

Reactor anomalies - E~few MeV

- $\bar{\nu}_e$ produced by reactors are measured by experiments via inverse beta decays:
 - large range of technologies used, results are overall consistent: ~6% deficit compared to flux predictions from Huber and Mueller et al. (HM)
 - no L/E dependence observed; leading interpretation is now that deficit is likely from deficiencies of HM flux prediction

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 - no L/E dependence observed; leading interpretation is now that deficit is likely from deficiencies of HM flux prediction
- Neutrino4 experiment recently reported an L/E dependence, with $>3\sigma$ significance
 - movable detector with L ranging 6-12 m
 - best fit point $\Delta m^2 \sim 7 \text{ eV}^2$ and $\sin^2(2\theta) = 0.36$
 - strong tension with results from PROSPECT

Meson decay anomalies: E~30-600 MeV

- LSND looked for oscillations of anti- v_{μ} produced by decay chain of pions at rest at L=30 m
 - relative beam content of anti- v_e is < 10⁻³
 - excess of anti- v_e events over prediction at 3.8 σ significance
 - interpreted as oscillation signal with $\Delta m^2 \sim 1 \text{ eV}^2$ $(\sin^2 2\theta, \Delta m^2)_{best-fit} = (0.003, 1.2 \text{eV}^2)^0$
 - KARMEN (lower sensitivity) did not see an excess

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 - interpreted as oscillation signal with $\Delta m^2 \sim 1 \text{ eV}^2$ $(\sin^2 2\theta, \Delta m^2)_{best-fit} = (0.003, 1.2 \text{eV}^2)^{10}$
 - KARMEN (lower sensitivity) did not see an excess
- MiniBooNE investigated the LSND result using a beam of 99.5% v_{μ} from meson decays in flight (BNB)
 - L=540 m, same L/E as LSND to test oscillation hypothesis
 - tested both neutrino and antineutrino beam
 - E_v reconstruction based on E_{lep} assuming CCQE process
 - excess of v_e -like events at 4.8 σ significance, mostly low-energy

Figure 31. A schematic drawing of the MiniBooNE detector.

- Cherenkov detector cannot distinguish rings from e vs single γ
 - nor it cannot observe the proton activity, mostly below Cherenkov threshold
- Two alternative mainstream hypotheses have been proposed as reason for the excess: v_e events (oscillations) or $\Delta \rightarrow N\gamma$ (unconstrained background)
 - other possible interpretations are however emerging (we'll discuss them later)

Enter MicroBooNE

- Designed to test the MiniBooNE low-energy excess (LEE)
 - same neutrino beam (BNB)
 - similar distance from source
 - analyzed about 1/2 data
- But experimental program goes well beyond the LEE
 - nu-Ar cross sections
 - BSM physics searches
 - LArTPC development
 - also off-axis to NuMI beam
- Part of broader SBN program to test short-baseline oscillations

MicroBooNE Simulation

MicroBooNE detector

- Charged particles produced in neutrino interactions ionize the argon, ionization electrons drift in electric field towards anode planes
- Sense wires detect the incoming charge, producing beautiful detector data images

MicroBooNE detector

- Charged particles produced in neutrino interactions ionize the argon, ionization electrons drift in electric field towards anode planes
- Sense wires detect the incoming charge, producing beautiful detector data images
- Full detail of neutrino interaction with O(mm) spatial resolution and calorimetric information
- Fast scintillation light detected by PMT system for triggering and cosmic rejection

Superior detector capabilities

Superior detector capabilities: e/g separation

Superior detector capabilities: e/g separation

Superior detector capabilities: proton detection

Searching for the MB excess in uB

- Our first tests of the MB excess look for the signature of leading interpretations of the excess:
 - v_e events (oscillation)
 - neutral current (NC) $\Delta \rightarrow \gamma$

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 Define empirical signal model in uB based on unfolded MB E_v data excess

Searching for the MB excess in uB

- Our first tests of the MB excess look for the signature of leading interpretations of the excess:
 - v_e events (oscillation)
 - neutral current (NC) $\Delta \rightarrow \gamma$

- Define empirical signal model in uB based on unfolded MB E_{ν} data excess
 - Prediction in uB is then obtained from a scale factor to the nominal prediction for that process
- Signal in uB can be tested as
 - simple hypothesis H0 vs H1
 - signal strength fit (normalization scaling parameter)

MICROBOONE-NOTE-1043

Search for Neutral Current Delta Radiative

- Assume excess is due to $\sim 3x$ larger NC $\Delta \rightarrow N\gamma$
- Apply to MicroBooNE to benchmark the analysis wrt excess
- $\Delta \rightarrow N\gamma$ search utilizes $1\gamma 1p$ and $1\gamma 0p$ events which are fit simultaneously to maximize signal statistics
- Major challenge is understanding and rejecting NC π^0 backgrounds
 - In situ measurement used to constrain the background

PRL 128 (2022) 11, 111801

1γ 0p candidate

NC Delta Radiative - Results

- Observe data-MC agreement within error in both channels
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PRL 128 (2022) 11, 111801
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- Control regions used to update prediction, improving agreement with data
 - no excess consistent with NC $\Delta \rightarrow \gamma$ is observed
- Limits can be interpreted in terms of:
 - LEE signal strength < 3 at 90% CL
 - Branching ratio for Delta radiative decay < 1.8%
 - Cross section $< 20 \cdot 10^{-42}$ cm⁻²/nucleon

Three complementary analyses testing different final states and using different reconstruction tools

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CCQE: 1e1p

Dominant interaction at low energies.

Leverage image-based reconstruction,

including DeepLearning tools

PRD 103, 052012 (2021)

Three complementary analyses testing different final states and using different reconstruction tools

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arXiv:2110.13978

Inclusive: 1eX

Largest statistics and sensitivity.

Use tomographic reconstruction,

native 3D, based on Wire-Cell

JINST 17 (2022) P01037

Common Analysis Flow

- PRD 105 (2022) 7, 072001 (1)
- JINST 14, P04004 (2019) (2)
- EPJC 79, 673 (2019) (3)
- PRApp. 15, 064071 (2021) (4)
- JINST 16 (2021) 12, T12017 (5)
- JINST 15, P02007 (2020) (6)
- PRD 103, 092003 (2021) (7)
- JHEP 12 (2021) 153 (8)
- EPJC 82 (2022) 5, 454 (9)
- (10) JINST 16, P09025 (2021)
- (11) JINST 15, P12037 (2020)

Common Analysis Flow

Blind analysis

Compatibility with H0 model

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Common Strategy: data-driven prediction

- Prediction from nominal modeling is "corrected" using v_{μ} data, based on covariance matrix formalism

• Correlations between v_{μ} and v_{e} events due to flux parentage and interaction

Common Strategy: data-driven prediction

- Prediction from nominal modeling is "corrected" using v_{μ} data, based on covariance matrix formalism
- - systematics are significantly reduced thanks to the in-situ measurement

• Correlations between v_{μ} and v_{e} events due to flux parentage and interaction

The prediction is updated in two aspects: the central value and the systematics

Common Strategy: Blind Analysis and Sideband Validations

- Signal region was kept blind until the last stage of the analysis to avoid bias Reconstruction and analysis developed on small-size (<1/10) open dataset Validated in high statistics control regions
 - e.g. π⁰ mass peak for shower energy scale
- Unique validation dataset given by v_e events from NuMI beam (off-axis)

arXiv:2110.14065

Results: nominal (no excess) v_e model prediction arXiv:2110.14065 arXiv:2110.14080 $1e0p0\pi v_e$ selection $pred = 0.86 \pm 0.06 (sys) \pm 0.19 (stat)$ Data (25) MicroBooNE 6.86 ×10²⁰ POT Unblinded data is fitted in CC ve (25.8) MICroBooNE 6.67 ×10²⁰ POT 17.5 ∧ 17.5 ₩ 15.0 00 12.5 eLEE model (x = 1)Fitted Background (3.2) Constrained prediction ≥ 15.0 Constrained uncertainty Pre-Constraint pre-defined intervals, is Prediction Unconstrained prediction Ō ----- LEE (11.6) 12.5 ~ 10.0 Dirt (Outside TPC) Events found to be in agreement 7.5 o 10.0 5.0Entries 7.5 with v_e model prediction 5.0 。2.5 Constrained 11 2.5 Uncertainties <u>۳</u>2.0 - frequentist p-values are 0.0 0.5 1.0 1.5 $\Box 0.5$ extracted for the energy Reconstructed E_{ν} [GeV] 0.0 800 1000 400 600 1200 E_{v} [MeV] spectrum: $1eNp0\pi v_e$ selection $\Sigma DATA/\Sigma(MC+EXT)=0.88\pm0.05(data err)\pm0.0$ Data POT: 6.369e+20 $\chi^{2/ndf=1}$ 100 MeV MicroBooNE 6.86 ×10²⁰ POT BNB data, 338.0 Pred. und • CCQE 1e1p: 0.014 EXT, 2.9 out FV. 5 eLEE model (x = 1)BNB Data $v_{\mu} CC \pi^{0}$ 20 v_{μ}^{μ} CC in FV, 8.5 LEE(x=1), 37.0 counts / Constrained prediction v_e CC v. CC in FV. 333. Constrained uncertainty v other GeV 1eNp0π, 1e0p0π: 0.18, 0.13 MicroBooNE Unconstrained prediction v with π^0 Dirt (Outside TPC) FC, constrained 15 Cosmics 0.14 • 1eX: 0.85 Entries - also looked at kinematic variables, confirming good Pred total uncertainty ┝╋╍╋╍┰╴ ┝╋╍╋╍╹╷┿╋╍╋╍╶┝╋╍╋╍ Dat agreement 1.5 1.0 2.0 $\frac{100 \quad 2000}{\text{Reconstructed E}_{v} (\text{MeV})}$ 0.5 500 1500 1000 Reconstructed E_{ν} [GeV] arXiv:2110.14065 arXiv:2110.13978

Results: limits on the eLEE model

- We reject our eLEE model (H1) at >97% CL for all high purity selections
 - including both exclusive (1e1p CCQE, 1eNp 0π) and inclusive (1eX) event classes
 - signal strength fit with Feldman Cousins procedure consistent with $\mu=0$

Evolution of LEE analyses

- No evidence for excesses relative to prediction for both processes

• Investigated if the MiniBooNE excess originates from v_e or NC $\Delta \rightarrow \gamma$ events

Evolution of LEE analyses

- No evidence for excesses relative to prediction for both processes
- Further steps include:
 - New interpretations of results in terms of oscillations or cross section
 - Search for different BSM explanations for the MB LEE

• Investigated if the MiniBooNE excess originates from v_e or NC $\Delta \rightarrow \gamma$ events

First results at Nu22 next week!

Topologies compatible with the MiniBooNE excess

Rich set of BSM models can explain the MiniBooNE excess

Decay of O(keV) Sterile Neutrinos to active neutrinos

- [13] Dentler, Esteban, Kopp, Machado Phys. Rev. D 101, 115013 (2020)
- [14] de Gouvêa, Peres, Prakash, Stenico JHEP 07 (2020) 141 Ο

New resonance matter effects

- [5] Asaadi, Church, Guenette, Jones, Szelc, PRD 97, 075021 (2018) Ο
- [16] Alves, Louis, deNiverville, [hep-ph]2201.00876 (2022) Ο
- Mixed O(1eV) sterile oscillations and O(100 MeV) sterile decay
 - [7] Vergani, Kamp, Diaz, Arguelles, Conrad, Shaevitz, Uchida, arXiv:2105.06470 Ο

Decay of heavy sterile neutrinos produced in beam

- [4] Gninenko, Phys.Rev.D83:015015,2011 Ο
- [12] Alvarez-Ruso, Saul-Sala, Phys. Rev. D 101, 075045 (2020) 0
- [15] Magill, Plestid, Pospelov, Tsai Phys. Rev. D 98, 115015 (2018) Ο
- [11] Fischer, Hernandez-Cabezudo, Schwetz, PRD 101, 075045 (2020) Ο
- [17] Dutta, Kim, Thompson, Thornton, Van de Water [hep-ph]2110.11944 Ο
- Decay of upscattered heavy sterile neutrinos or new scalars mediated by Z' or more complex higgs sectors
 - [1] Bertuzzo, Jana, Machado, Zukanovich Funchal, PRL 121, 241801 (2018) Ο
 - [2] Abdullahi, Hostert, Pascoli, Phys.Lett.B 820 (2021) 136531 Ο
 - [3] Ballett, Pascoli, Ross-Lonergan, PRD 99, 071701 (2019) 0
 - [10] Dutta, Ghosh, Li, PRD 102, 055017 (2020) 0
 - [6] Abdallah, Gandhi, Roy, Phys. Rev. D 104, 055028 (2021) Ο
- Decay of axion-like particles
 - [8] Chang, Chen, Ho, Tseng, Phys. Rev. D 104, 015030 (2021)

A model-independent approach to any new particle

[9] Brdar, Fischer, Smirnov, PRD 103, 075008 (2021)

A couple of examples

Dark neutrino portal

- with dark Z' decay
- could explain MiniBooNE: if e+e- resolved as single shower

- Dark matter produced in beamline
 - scattering off argon

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Ν

BSM Searches in uB: e+e-

- MicroBooNE has been producing a rich set of BSM searches, including
 - heavy neutral leptons, n-nbar oscillations, supernovae v reconstruction, Higgs portal scalars

PRD 101, 052001 (2020)

MICROBOONE-NOTE-1113-PUB

- The Higgs portal scalar search in particular pioneered e⁺e⁻ searches in uB

 - "Portal" to the dark sector, via a dark scalar mixing with the Higgs - Kaons decaying to scalar particle in beam, then scalar decays to fermion pair in detector
 - Our first search uses kaons decaying at rest in the NuMI beam dump
 - Limits are competitive, extend exclusion in mixing vs mass parameter space
- Analyses in BNB data including unresolved e⁺e⁻ pairs will be key to test further BSM interpretations of the MB excess

Conclusions

- Era of precision neutrino physics with LArTPC experiments has started
- Puzzle of short baseline neutrino anomalies is not yet complete - MicroBooNE has addressed the leading explanations of the MiniBooNE anomaly
- MicroBooNE is still looking for cuts in the canvas of the SM picture

Thank you!

MicroBooNE, EXT data, 2019

Backup slides

HNL Search in MicroBooNE

PRD 101, 052001 (2020)

Test of Null Hypothesis

(a) Electron angle relative to beam direction.

(a) All selected events.

arXiv:2110.14065

-	Channel	χ^2	$\chi^2/{ m dof}$	<i>p</i> -
-	$1e{ m N}p0\pi$	15.2	1.52	0
	$1e0p0\pi$	16.3	1.63	0
	$1eNp0\pi + 1e0p0\pi$	31.50	1.58	0

Fermilab

(b) Low energy selected events from 0.15–0.43 GeV.

value 0.182).126 0.098

Simple Hypothesis Test

		obs.	$\Delta \chi^2 < \text{obs.}$	$\Delta \chi^2 < \text{obs.}$	Sensitivity	Sensitivity
	Channel	$\Delta\chi^2$	$p ext{-value}(H_0)$	$p ext{-value}(H_1)$	$p ext{-value}(H_0) \mid H_1$	p -value $(H_1) \mid H_0$
	$1e\mathrm{N}p0\pi$	-3.89	0.285	0.021	0.957	0.061
arXiv:2110.14065	$1e0p0\pi$	3.11	0.984	0.928	0.759	0.249
	$1eNp0\pi + 1e0p0\pi$	-0.58	0.748	0.145	0.968	0.049

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Signal Strength Fit

Channel $1eNp0\pi$ $1e0p0\pi$ $1eNp0\pi + 1e0p0\pi$ 0

arXiv:2110.14065

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ata	Data	Sensitivity
$l_{ m BF}$	90% CL interval on μ	90% upper limit on μ
.00	[0.00, 0.82]	1.16
.00	$[1.13 \;,\; 15.01]$	3.41
.36	$[0.00 \;,\; 1.57]$	1.07

Reactor experiments

\overline{a}	Experiment	f^{a}_{235}	f^{a}_{238}	f^{a}_{239}	f^{a}_{241}	$\sigma_{f,a}^{exp}$	$R_{a,HM}^{exp}$	$R_{a,EF}^{exp}$	$R_{a,HKSS}^{exp}$	$R_{a,KI}^{exp}$	δ^{exp}_a [%]	$\delta_a^{ m cor}$ [%]	L_a [m]
1	Bugey-4	0.538	0.078	0.328	0.056	5.75	0.927	0.962	0.916	0.962	1.4		15
2	Rovno91	0.614	0.074	0.274	0.038	5.85	0.924	0.965	0.914	0.962	2.8	<u>}</u> 1.4	18
3	Rovno88-1I	0.607	0.074	0.277	0.042	5.70	0.902	0.941	0.892	0.939	6.4		18
4	Rovno88-2I	0.603	0.076	0.276	0.045	5.89	0.931	0.971	0.920	0.969	6.4	^{3.1}	17.96
5	Rovno88-1S	0.606	0.074	0.277	0.043	6.04	0.956	0.997	0.945	0.995	7.3	} 2.2	18.15
6	Rovno88-2S	0.557	0.076	0.313	0.054	5.96	0.956	0.994	0.945	0.993	7.3	3.1	25.17
7	Rovno88-3S	0.606	0.074	0.274	0.046	5.83	0.922	0.962	0.911	0.960	6.8		18.18
8	Bugey-3-15	0.538	0.078	0.328	0.056	5.77	0.930	0.966	0.920	0.966	4.2		15
9	Bugey-3-40	0.538	0.078	0.328	0.056	5.81	0.936	0.972	0.926	0.972	4.3	\ 4.0	40
10	Bugey-3-95	0.538	0.078	0.328	0.056	5.35	0.861	0.895	0.852	0.894	15.2	J	95
11	Gosgen-38	0.619	0.067	0.272	0.042	5.99	0.949	0.992	0.939	0.988	5.4		37.9
12	Gosgen-46	0.584	0.068	0.298	0.050	6.09	0.975	1.016	0.964	1.014	5.4	2.0	45.9
13	Gosgen-65	0.543	0.070	0.329	0.058	5.62	0.909	0.945	0.899	0.944	6.7		64.7
14	ILL	1.000	0.000	0.000	0.000	5.30	0.787	0.843	0.777	0.827	9.1	Ĵ	8.76
15	Krasnoyarsk87-33	1	0	0	0	6.20	0.920	0.986	0.909	0.967	5.2		32.8
16	Krasnoyarsk87-92	1	0	0	0	6.30	0.935	1.002	0.924	0.983	20.5	4.1	92.3
17	Krasnoyarsk94-57	1	0	0	0	6.26	0.929	0.995	0.918	0.977	4.2	Ó	57
18	Krasnoyarsk99-34	1	0	0	0	6.39	0.948	1.016	0.937	0.997	3.0	0	34
19	SRP-18	1	0	0	0	6.29	0.934	1.000	0.923	0.982	2.8	0	18.2
20	SRP-24	1	0	0	0	6.73	0.998	1.070	0.987	1.050	2.9	0	23.8
21	Nucifer	0.926	0.008	0.061	0.005	6.67	1.007	1.074	0.995	1.056	10.8	0	7.2
22	Chooz	0.496	0.087	0.351	0.066	6.12	0.990	1.025	0.979	1.027	3.2	0	≈ 1000
23	Palo Verde	0.600	0.070	0.270	0.060	6.25	0.991	1.033	0.980	1.031	5.4	0	≈ 800
24	Daya Bay	0.564	0.076	0.304	0.056	5.94	0.950	0.988	0.939	0.987	1.5	0	≈ 550
25	RENO	0.571	0.073	0.300	0.056	5.85	0.936	0.974	0.925	0.973	2.1	0	≈ 411
26	Double Chooz	0.520	0.087	0.333	0.060	5.71	0.918	0.952	0.907	0.953	1.1	0	≈ 415
27	STEREO	1	0	0	0	6.34	0.941	1.008	0.930	0.989	2.5	0	9 - 11

group of experiments indicated by the braces; L_a is the source-detector distance.

Table 1: List of the experiments which measured the absolute reactor antineutrino flux [66]. For each experiment numbered with the index a: f^a_{235} , f^a_{238} , f^a_{239} , and f^a_{241} are the effective fission fractions of the four isotopes 235 U, 238 U, 239 Pu, and 241 Pu, respectively; $\sigma_{f,a}^{exp}$ is the experimental IBD yield in units of 10^{-43} cm²/fission; $R_{a,HM}^{exp}$, $R_{a,EF}^{exp}$, $R_{a,HKSS}^{exp}$, and $R_{a,KI}^{exp}$, are the ratios of measured and predicted rates for the IBD yields of the models in Tab. 6; δ_a^{exp} is the total relative experimental statistical plus systematic uncertainty, δ_a^{cor} is the part of the relative experimental systematic uncertainty which is correlated in each

arXiv:2203.07323

Alternative reactor flux models

Model	Rat	tes	Evolu	ution	Rates + Evol		
	$\overline{oldsymbol{R}}_{\sf mod}$	RAA	$\overline{oldsymbol{R}}_{mod}$	RAA	$\overline{oldsymbol{R}}_{mod}$		
НМ	$0.936\substack{+0.024\\-0.023}$	2.5σ	$0.933\substack{+0.025\\-0.024}$	2.6σ	$0.930\substack{+0.024\\-0.023}$		
EF	$0.960\substack{+0.033\\-0.031}$	1.2σ	$0.975\substack{+0.032\\-0.030}$	0.8σ	$0.975\substack{+0.032\\-0.030}$		
HKSS	$0.925\substack{+0.025\\-0.023}$	2.9σ	$0.925\substack{+0.026\\-0.024}$	2.8σ	$0.922\substack{+0.024\\-0.023}$		
KI	$0.975\substack{+0.022\\-0.021}$	1.1σ	$0.973\substack{+0.023\\-0.022}$	1.2σ	0.970 ± 0.021		

Table 7: Average ratio \overline{R}_{mod} obtained in Ref. [66] from the least-squares analysis of the reactor rates in Tab. 1 and of the Daya Bay [251] and RENO [252] evolution data for the IBD yields of the models in Tab. 6. The RAA columns give the corresponding statistical significance of the reactor antineutrino anomaly. The descriptions of all the models are given in text.

arXiv:2203.07323

lution RAA $2.8\,\sigma$ $0.8\,\sigma$ $3.0\,\sigma$ $1.4\,\sigma$

Figure 38: Contours of the 2σ allowed regions in the $(\sin^2 2\vartheta_{ee}, \Delta m_{41}^2)$ plane obtained from the combined neutrino oscillation fit of the reactor rates in Tab. 7 and the Daya Bay [251] and RENO [252] evolution data. The blue, red, green, and magenta curves correspond, respectively, to the HM, EF, HKSS, and KI models in Tab. 6. Also shown are the contour of the 2σ allowed regions of the Gallium anomaly obtained in Ref. [137] from the combined analysis of the GALLEX, SAGE and BEST data (orange curve), and the 2σ bound obtained from the analysis of solar neutrino data in Ref. [258] (dark red vertical line).

Alternative reactor flux models

PLB 829 (2022) 137054

MiniBooNE 3+1 Oscillation Fit

PRD 103 (2021) 5, 052002

$\Delta m^2 = 0.043 \text{ eV}^2$, and $\sin^2 2\theta = 0.807$,

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Appearance vs Disappearance

Gallium cross sections

	GALLEX ₁	$GALLEX_1$	$SAGE_{\mathrm{Cr}}$	$SAGE_{\mathrm{Ar}}$	Avg.
$R_{\rm B}$	0.95 ± 0.11	$0.81\substack{+0.10 \\ -0.11}$	0.95 ± 0.12	0.79 ± 0.08	0.86 ± 0.05
$R_{ m HK}$	0.85 ± 0.12	0.71 ± 0.11	$0.84\substack{+0.13 \\ -0.12}$	0.71 ± 0.09	0.77 ± 0.08
$R_{ m FF}$	0.93 ± 0.11	$0.79\substack{+0.10 \\ -0.11}$	$0.93\substack{+0.11 \\ -0.12}$	$0.77\substack{+0.09 \\ -0.07}$	0.84 ± 0.05
$R_{ m HF}$	$0.83\substack{+0.13 \\ -0.11}$	0.71 ± 0.11	$0.83\substack{+0.13 \\ -0.12}$	$0.69\substack{+0.10 \\ -0.09}$	$0.75\substack{+0.09 \\ -0.07}$
$R_{ m JUN45}$	0.97 ± 0.11	0.83 ± 0.11	0.97 ± 0.12	0.81 ± 0.08	0.88 ± 0.05

arXiv:2203.07323

KARMEN limits

100 $\Delta m^2 [eV^2]$

10

PRD 65 (2002) 112001

0.1

External 3+1 Fit of uB data

arXiv:2111.10359

arXiv:2109.11482

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Figure 66: BEST detector configuration [137].

Figure 67: Allowed regions obtained in the $\sin^2 2\theta - \Delta m_{41}^2$ parameter space from the analysis of the BEST results only (left) and from the BEST results combined with results from GALLEX and SAGE (right) [294]. Note that here $\sin^2 2\theta = 4 |U_{e4}|^2 (1 - |U_{e4}|^2).$

