Reevaluating Uncertainties in $\bar{B} \to X_s \gamma$ Decay

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Based on A.G and Gil Paz "Reevaluating Uncertainties in $\bar{B} \to X_s \gamma$ " JHEP 11 (2019) 141

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 - It is suppressed at tree level in SM
 - Can receive contributions from SM extensions.

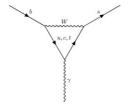


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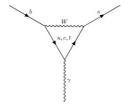


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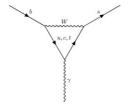


Figure: $b o s \gamma$ flavor changing neural current (FCNC) in SM

- SM extensions modify the $C_{7\gamma}$ Wilson coefficient
- CP violation in $\bar{B} o X_{\rm s} \gamma$ can be enhanced by new physics

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- At leading power: Only $Q_{7\gamma}-Q_{7\gamma}$ contributes to decay rate
- At $1/m_b$: Γ get $Q_1-Q_{7\gamma}$, $Q_{8g}-Q_{8g}$ and $Q_{7\gamma}-Q_{8g}$ contributions

Decay rate

World average for experimental value:

$$\mathcal{B}(B \to X_s \gamma) (E_{\gamma} > 1.6 \text{ GeV}) = (3.32 \pm 0.15) \times 10^{-4}$$

[Y. Amhis et. al. EPJC 77, 895 (2017)]

NNLO prediction

$$\Gamma\left(\overline{B} \to X_q \gamma\right) = \underbrace{\Gamma\left(b \to X_q^p \gamma\right)}_{\text{Perturbatively calculable}} + \underbrace{\delta\Gamma_{\text{nonp}}}_{\mathcal{O}(\frac{\Lambda_{\text{QCD}}}{m_b})}$$

SM prediction (2015) [Misiak et. al. PRL 114, 221801 (2015)]

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- $\delta\Gamma_{\rm nonp} \equiv \text{Non-perturbative contribution}$
 - The largest contribution to the error 5% from $\mathcal{O}(\frac{\Lambda_{QCD}}{m_b})$

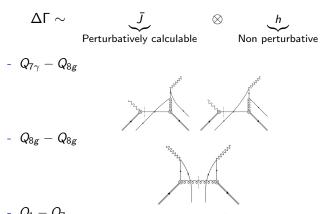
Order $1/m_b$ power corrections to $\Gamma(\bar{B} \to X_s \gamma)$

Non-perturbative effects arise from Resolved Photon Contributions

$$\Delta\Gamma \sim \underbrace{\bar{\textit{J}}}_{\text{Perturbatively calculable}} \otimes \underbrace{\textit{h}}_{\text{Non perturbative}}$$

Order $1/m_b$ power corrections to $\Gamma(\bar{B} \to X_s \gamma)$

Non-perturbative effects arise from Resolved Photon Contributions





- 2010 estimates for non-perturbative contribution to error
 - From $Q_1^c Q_{7\gamma} \in [-1.7, +4.0]\%$
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- Now $Q_1^c Q_{7\gamma}$ is the largest contribution to the error! Can we reduce it?

$Q_1^c - Q_{7\gamma}$ contribution

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where

$$\Lambda_{17} = e_c \operatorname{Re} \int_{-\infty}^{\infty} \frac{d\omega_1}{\omega_1} \left| 1 - \underbrace{F\left(\frac{m_c^2 - i\varepsilon}{m_b\omega_1}\right)}_{\text{perturbative}} + \frac{m_b\omega_1}{12m_c^2} \right| \underbrace{h_{17}\left(\omega_1\right)}_{\text{non-perturbative}}$$

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- Need a new model for h₁₇ to reduce the error
 - New information on moments of h_{17} : constrain **new model**
 - What can we learn from moments?

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$$\begin{split} &h_{17}(\omega_1) = \\ &= \int \frac{dr}{2\pi} e^{-i\omega_1 r} \frac{\langle \bar{B}|(\bar{h}S_{\bar{n}})(0) \not n (1+\gamma_5) i \gamma^{\perp} \bar{n}_{\beta} (S_{\bar{n}} g G^{\alpha\beta} S_{\bar{n}}) (r\bar{n}) (S_{\bar{n}}^{\dagger} h)(0) |\bar{B}\rangle}{2M_B} \end{split}$$

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$$S_n(x) = \mathbf{P} \exp \left(ig \int_{-\infty}^0 du n \cdot A_s(x + un) \right)$$

$$n^{\mu} \equiv (1,0,0,1)$$
 and $\overline{n}^{\mu} \equiv (1,0,0,-1)$

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• Using the (new) identity

$$i\bar{n}\cdot\partial\left(S_{\bar{n}}^{\dagger}(x)O(x)S_{\bar{n}}(x)\right)=S_{\bar{n}}^{\dagger}(x)[i\bar{n}\cdot D,O(x)]S_{\bar{n}}(x)$$

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- New result Moments over ω₁

$$\langle \omega_1^k h_{17} \rangle = (-1)^k \frac{1}{2M_B} \langle \bar{B} | \bar{h} \cdots \underbrace{[i\bar{n} \cdot D, [i\bar{n} \cdot D, \cdots [i\bar{n} \cdot D, \cdots [i\bar{n} \cdot D]]] s^{\lambda} h | \bar{B} \rangle}_{k \text{ times}}$$

Moments of the g_{17}

 Procedure to obtain these HQET matrix elements derived in [A. Gunawardana and G. Paz, JHEP 07(2017)137 [arXiv:1702.08904]]

$$\langle h_{17}
angle = 2\lambda_2 = 2\mu_G^2/3$$
 $\langle \omega_1^2 h_{17}
angle = rac{2}{15} \left(5m_5 + 3m_6 - 2m_9
ight)$ New result

m_i were extracted from data for the first time in 2016
 [P. Gambino, K. J Healey, S. Turczyk PLB 763, 60 (2016)]

$$\mu_G^2 = 0.355 \pm 0.060 \text{ GeV}^2$$
 $m_5 = 0.072 \pm 0.045 \text{ GeV}^4$ $m_6 = 0.060 \pm 0.164 \text{ GeV}^4$ $m_9 = -0.280 \pm 0.352 \text{ GeV}^4$

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 - 2010 models provide $\langle \omega_1^2 h_{17} \rangle \in (-0.31, 0.49) \text{ GeV}^4$.
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 - These older models were constructed before m_i were extracted
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- Expect in future
 - Further improvements on HQET matrix elements
 - Belle II or LQCD data ⇒ Better constrains on moments

Applications

• Properties of h_{17}

- Properties of h₁₇
 - Real and even function over ω_1
 - $\langle \omega_1^k h_{17}(\omega_1) \rangle = 0$ for $k=1,3,5,\cdots$
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• $|h_{17}| < 1$ GeV and no peaks beyond $\omega_1 = 1$ GeV

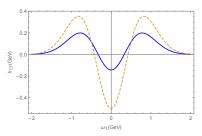


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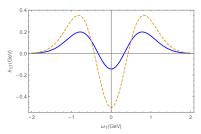


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• Orange dashed line: 2010 model $h_{17}\left(\omega_1,\mu\right)=rac{2\lambda_2}{\sqrt{2\pi}\sigma}rac{\omega_1^2-\Lambda^2}{\sigma^2-\Lambda^2}e^{-rac{\omega_1^2}{2\sigma^2}}$

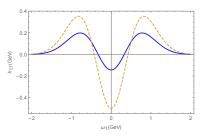


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 - $\sigma=0.5$ GeV, $\Lambda=0.425$ GeV and $\Rightarrow \langle \omega_1^2 h_{17} \rangle=0.49$ GeV⁴

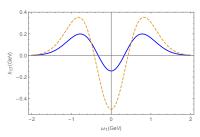


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 - $\sigma = 0.5 \text{ GeV}, \Lambda = 0.425 \text{ GeV} \text{ and } \Rightarrow \langle \omega_1^2 h_{17} \rangle = 0.49 \text{ GeV}^4$
- Blue line; 2019 model: $\sigma=0.5$ GeV and $\langle \omega_1^2 h_{17} \rangle=0.27$ GeV⁴

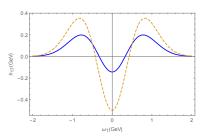


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 - $\sigma = 0.5 \text{ GeV}, \Lambda = 0.425 \text{ GeV}$ and $\Rightarrow \langle \omega_1^2 h_{17} \rangle = 0.49 \text{ GeV}^4$
- Blue line; 2019 model: $\sigma=0.5$ GeV and $\langle \omega_1^2 h_{17} \rangle=0.27$ GeV⁴
- New function is 50% smaller than the 2010

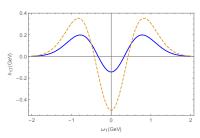


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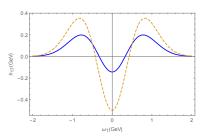


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- Consider also unknown higher moments, up to 6 Hermite polynomials

• New results: Our estimates for nonperturbative parameter Λ_{17}

$$-24\,\text{MeV} < \Lambda_{17} < +5\,\text{MeV}$$

- Compare with the 2010 estimate $-60\,\text{MeV} < \Lambda_{17} < 25\,\text{MeV}$ [Benzke, Lee, Neubert and Paz, JHEP 08(2010) 099]

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- Obtain new estimate for $Q_1 Q_{7\gamma}$ contribution:

$$\frac{C_1}{C_{7\gamma}}\frac{\Lambda_{17}}{m_b}$$

At
$$\mu = 1.5 \text{ GeV}$$
 : $C_1(\mu) = 1.257$, $C_{7\gamma}(\mu) = -0.407$ and $m_b = 4.58 \text{ GeV}$

• New result: $Q_1 - Q_{7\gamma}$ contribution to nonperturbative uncertainty

$$Q_1 - Q_{7\gamma} \in \boxed{[-0.3, +1.6]\%}$$

- Compare with 2010 estimate $\in [-1.7, +4.0]\%$

• New Belle data: $Q_{7\gamma} - Q_{8g}$ nonperturbative uncertainty

$$Q_{7\gamma}-Q_{8g}\in[-1.4,+2]\%$$

[Watanuki et. al. PRD 99, 032012 (2019)] [Gunawardana and Paz, JHEP 11(2019)141 [arXiv:1908.02812]]

2010 estimates	2019 estimates	
[-1.7, +4.0]%	[-0.3, +1.6]%	Gunawardana and Paz (20
[-4.4, +5.6]%	[-1.4, +2]%	Watanuki et al. (2019
[-0.3, +1.9]%	[-0.3, +1.9]%	Benzke et al. (2010)
[-6.4, +11.5]%	[-2.0, +5.5]%	Gunawardana and Paz (20
	[-1.7, +4.0]% [-4.4, +5.6]% [-0.3, +1.9]%	

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Operator pair	2010 estimates	2019 estimates	
$Q_1-Q_{7\gamma}$	[-1.7, +4.0]%	[-0.3, +1.6]%	Gunawardana and Paz (20
$Q_{7\gamma}-Q_{8g}$	[-4.4, +5.6]%	[-1.4, +2]%	Watanuki et al. (2019
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- Total uncertainty is reduced by half!
- Total uncertainty obtained by scanning over each uncertainty

CP asymmetry: phenomenological estimates • Resolved photon contribution to CP asymmetry

$$\begin{split} \mathcal{A}_{X_s\gamma}^{\mathrm{SM}} &= \left(1.15 \times \frac{\tilde{\Lambda}_{17}^u - \tilde{\Lambda}_{17}^c}{300 \mathrm{MeV}} + 0.71\right)\% \\ \tilde{\Lambda}_{17}^u &= \frac{2}{3} \, h_{17}(0) \\ \tilde{\Lambda}_{17}^c &= \frac{2}{3} \, \int_{4m_c^2/m_b}^{\infty} \frac{d\omega_1}{\omega_1} \, \underbrace{f\left(\frac{m_c^2}{m_b\omega_1}\right)}_{\text{Perturbative}} \underbrace{h_{17}(\omega_1)}_{\text{Non-perturbative}} \end{split}$$

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Comparison between 2010 and new values

Operator pair	2010 estimates	2019 estimates
$\tilde{\Lambda}^u_{17}$	[-330 MeV, 525 MeV]	[-660 MeV, 660 MeV]
$\tilde{\Lambda}^{c}_{17}$	[-9 MeV, 11 MeV]	[-7 MeV, 10 MeV]
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Direct CP Asymmetry experimental estimate:

$$A_{CP} = (1.5 \pm 2.0) \%$$

[Amhis et. al. EPJC 77, 895 (2017)]

Conclusion

- $\bar{B} o X_s \gamma$ is a important New Physics probe
- Non perturbative error of the decay rate is 5%
- $Q_1^c Q_{7\gamma}$ is the largest contribution to the error
- Better estimates for $Q_1^c-Q_{7\gamma}$ obtained from moments of h_{17}
- New estimates for CP asymmetry

Work in progress

- Applications to Inclusive $b \rightarrow u$ transitions
 - Systematic parameterization of shape functions using new moment information (work in progress)
 - Update to the BLNP method